

Point Spread Function Reconstruction in Ground-based Astronomy

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Outline

- 1. Ground-based Astronomy**
- 2. Models and Solution Methods**
- 3. High-resolution Image Reconstruction**

What is (gray-scale) image?



pixels

color image $\times 3$

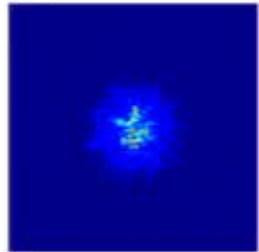
| | | | | | | |
|---|-----|-----|-----|-----|-----|---|
| | | | ⋮ | | | |
| | 218 | 215 | 226 | 223 | 226 | |
| | 225 | 231 | 243 | 237 | 128 | |
| ⋯ | 135 | 136 | 145 | 136 | 31 | ⋯ |
| | 30 | 30 | 30 | 31 | 31 | |
| | 0 | 0 | 0 | 0 | 1 | |
| | | | ⋮ | | | |

pixel values between
0 and 255 (8-bit)

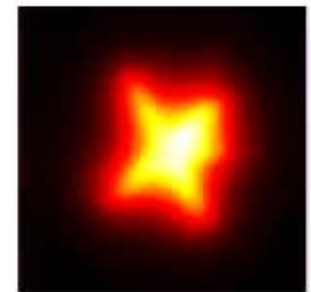
1000-by-1000 image = 1000-by-1000 matrix
concatenate into 1M-vector

Ground-Based Astronomy

true image
 $f(x, y)$



point spread
function
 $k(x, y)$



observed image
 $g(x, y)$

Unknown Point Spread Functions

Some well-known approaches for getting $k(x, y)$:

- Blind-deconvolution to simultaneously obtain $k(x, y)$ and $f(x, y)$:

$$g(x, y) = k^{(i+\frac{1}{2})}(x, y) * f^{(i)}(x, y) + n(x, y), \quad i = 1, 2, \dots$$

[T. Chan & Wong IEEE TIP 98]



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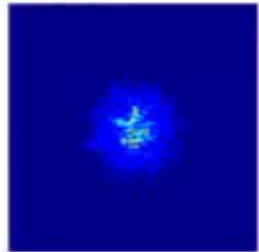
[T. Chan & Wong IEEE TIP 98]

- Reconstruct $k(x, y)$ by some means (e.g. natural or artificial guide-star)

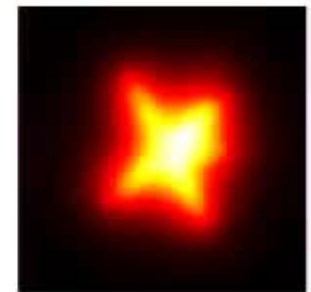


Ground-Based Astronomy

true image
 $f(x, y)$



point spread
function
 $k(x, y)$



observed image
 $g(x, y)$

Point-spread Function Reconstruction

Planar waves change across atmospheric turbulence

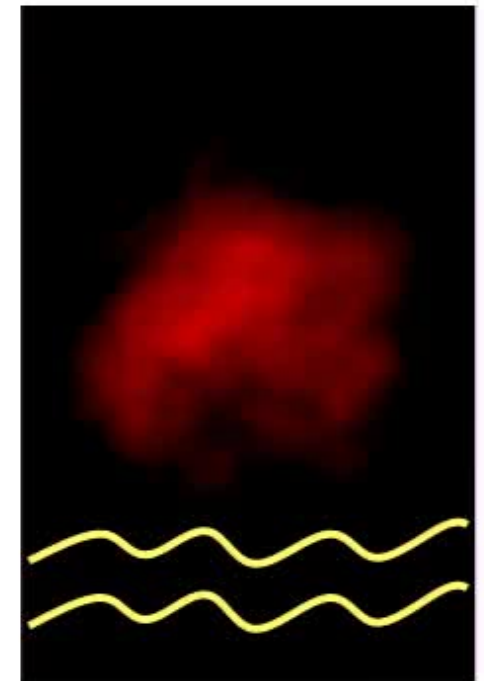
- $\phi(x, y)$: deviation from planarity is called **phase error** or **phase**

Fourier optics model:

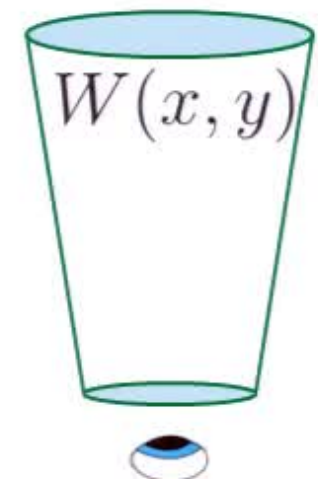
$$k(x, y) = \left| \mathcal{F}^{-1} \{ W(x, y) e^{i\phi(x, y)} \} \right|^2$$

- $W(x, y)$: aperture of the telescope
- \mathcal{F} : Fourier transform

[Goodman 96, Bardsley SIMAX, 08]



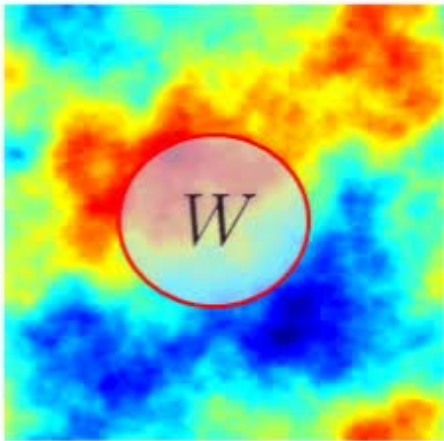
$\phi(x, y)$



Wavefront to Wavefront Gradient

Phase $\phi(x, y)$ cannot be directly measured, only its gradients by wavefront sensors:

phase ϕ

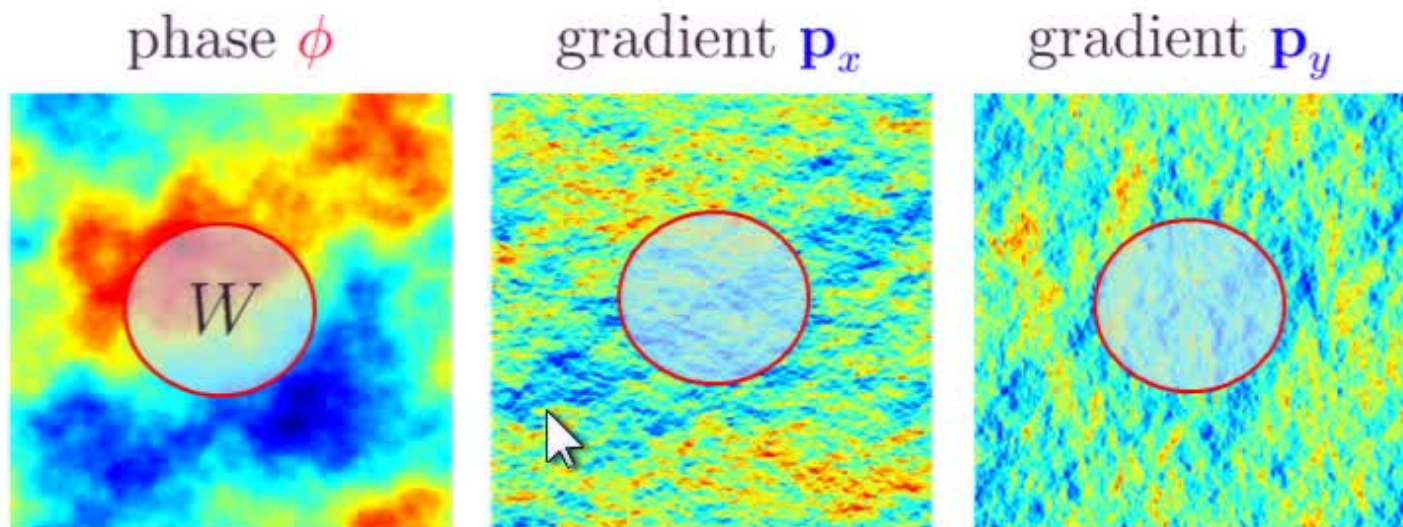


Wavefront to Wavefront Gradient

Phase $\phi(x, y)$ cannot be directly measured, only its gradients by wavefront sensors:

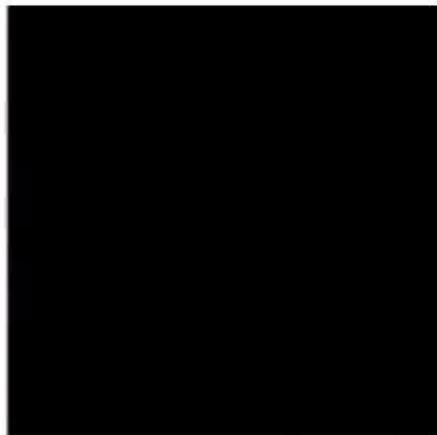
- $\mathbf{p}_x = D_x \phi(x, y)$: horizontal wavefront gradient
- $\mathbf{p}_y = D_y \phi(x, y)$: vertical wavefront gradient

D_i : 1st-order derivative operator modeling the sensor

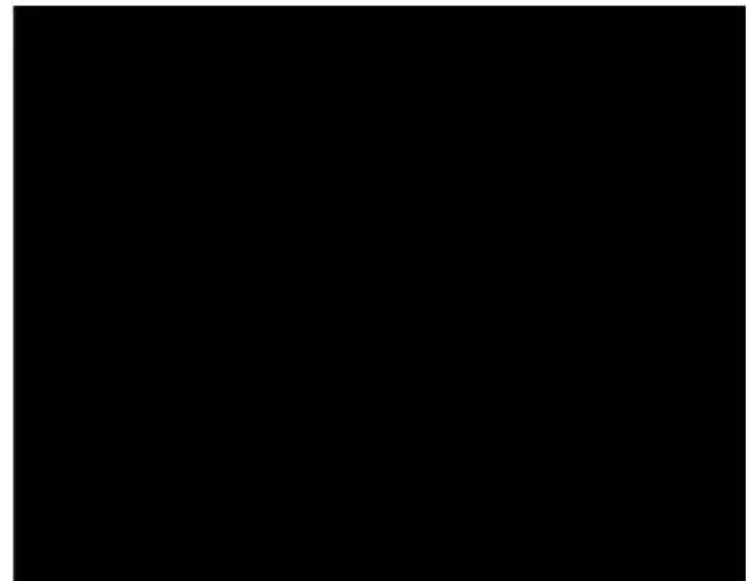


The Problem

- Wavefront sensors collect wavefront gradients $D_i\phi(x, y)$, not the phase $\phi(x, y)$
- $D_i\phi(x, y)$ are collected on coarse grids



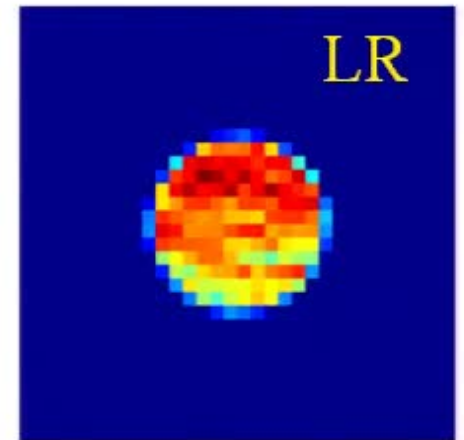
Phase $\phi(x, y)$



$D_x\phi(x, y)$

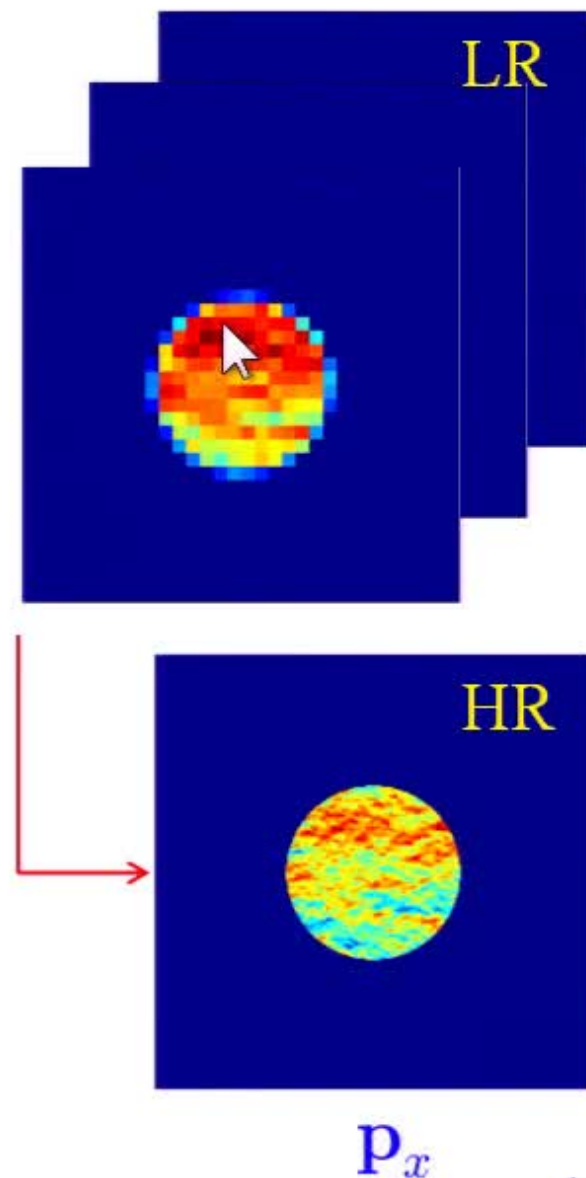
- Not accurate to compute ϕ from $D_i\phi(x, y)$

The Aim



The Aim

- Use a sequences of low-resolution (LR) frames of wavefront gradients to obtain the high-resolution (HR) wavefront gradients $\mathbf{p}_i = D_i\phi(x, y)$
- From HR wavefront gradient $D_i\phi(x, y)$ reconstruct more accurate $\phi(x, y)$
- From $\phi(x, y)$ reconstruct $k(x, y)$
- Using $k(x, y)$, deblur $g(x, y)$ to get $f(x, y)$



Frozen Flow Hypothesis

Within a short time interval, phase does not change.

Move telescope to get a sequence of LR frames of wavefront gradients to reconstruct the HR wavefront gradients.

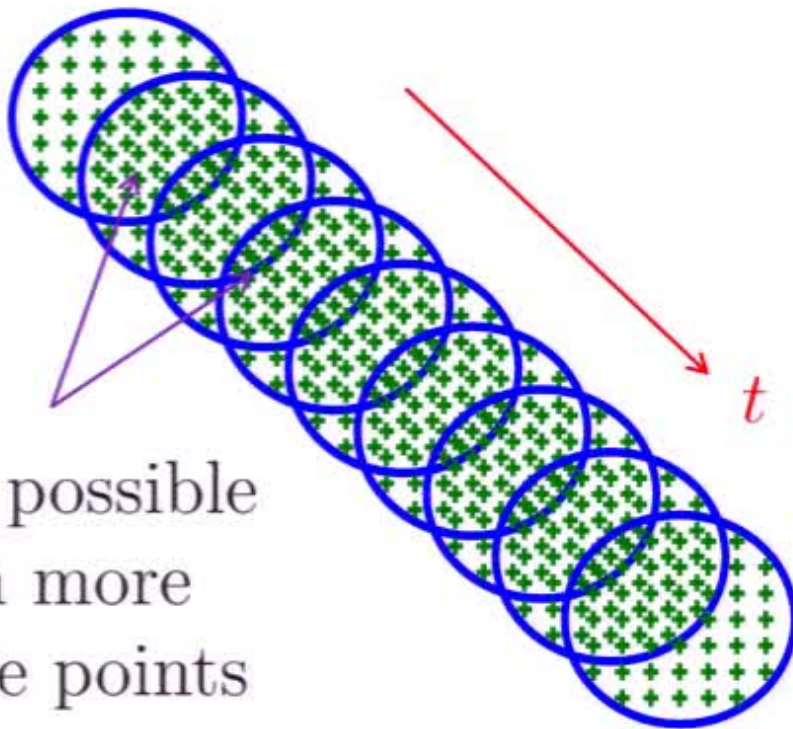
A LR wavefront
gradients
sensor



Frozen Flow Hypothesis

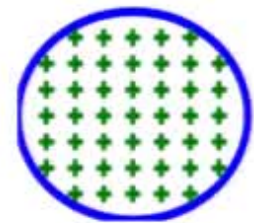
Within a short time interval, phase does not change.

Move telescope to get a sequence of LR frames of wavefront gradients to reconstruct the HR wavefront gradients.



HR is possible
with more
sample points

A LR wavefront
gradients
sensor



[Jefferies & Hart, 10]

Video Enhancement



*A 352-by-288 video
from a video recorder*

20 to 30 frames/second



*Bilinear interpolation
from 1 frame*

Video Enhancement

Reference frame



*Improving
resolution
of LR
reference
frame*

Video Enhancement

Reference frame

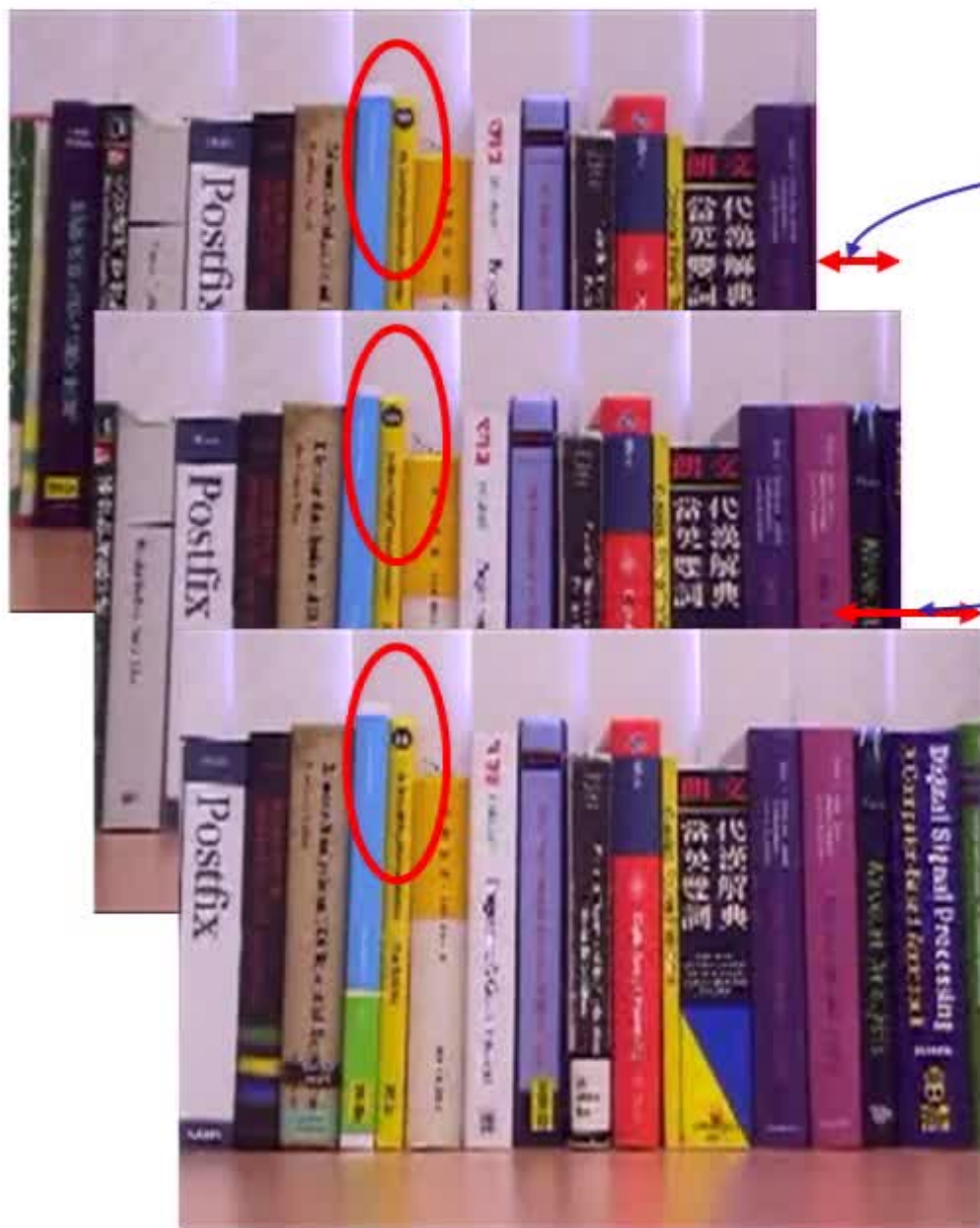


*Improving
resolution
of LR
reference
frame*

t

Video Enhancement

Reference frame



Displacement

ϵ

*Improving
resolution
of LR
reference
frame*

Video Enhancement



*A 352-by-288 video
from a video recorder*

Video Enhancement



*A 352-by-288 video
from a video recorder*



*Tight-frame method
using 21 frames*

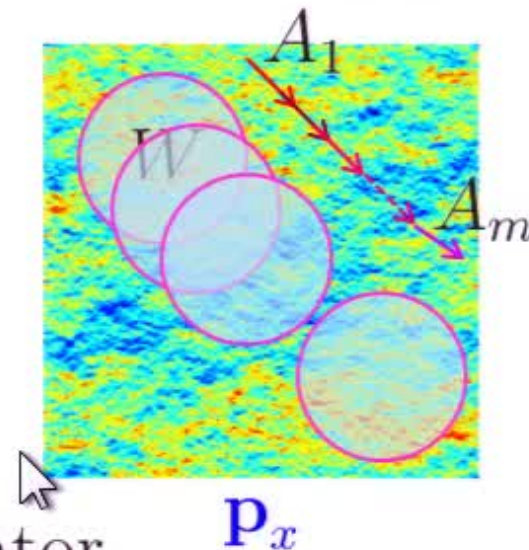
[C., Shen, & Xia, ACHA 08]

More on this later
in Part 3

Relation between HR and LR Gradients

$$\mathbf{q}_x^i = RW A_i \mathbf{p}_x + \mathbf{n}_x^i, \quad i = 1, 2, \dots, m$$
$$\mathbf{q}_y^i = RW A_i \mathbf{p}_y + \mathbf{n}_y^i, \quad i = 1, 2, \dots, m$$

- $\mathbf{p}_x, \mathbf{p}_y$: HR wavefront gradients
- $\mathbf{q}_x^i, \mathbf{q}_y^i$: sequences of LR wavefront gradients
- $\mathbf{n}_x^i, \mathbf{n}_y^i$: noise
- A_i : motion operator
- W : aperture operator
- R : down-sampling operator



Tikhonov ℓ_2 - ℓ_2 Model

[Nagy, Jefferies, & Chu, Maui Conf. 10, SISC 13]:

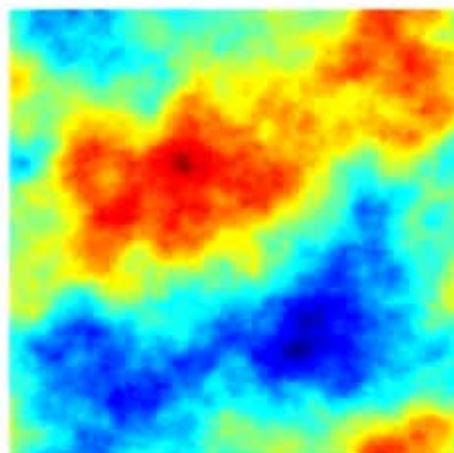
$$\min_{\mathbf{p}_x} \|\mathbf{p}_x\|_2^2 + \frac{\alpha}{2} \sum_{i=1}^m \|RW A_i \mathbf{p}_x - \mathbf{q}_x^i\|_2^2$$

$$\min_{\mathbf{p}_y} \|\mathbf{p}_y\|_2^2 + \frac{\alpha}{2} \sum_{i=1}^m \|RW A_i \mathbf{p}_y - \mathbf{q}_y^i\|_2^2$$

□ linear solve with $[I + \alpha \sum_i (RW A_i)^T RW A_i]$

□ $\|\mathbf{p}_x\|_2^2, \|\mathbf{p}_y\|_2^2 \approx \|\nabla \phi\|_2^2$

□ may smooth the edges
in ϕ



phase ϕ

Combined Model for the Phase

□ Note that:

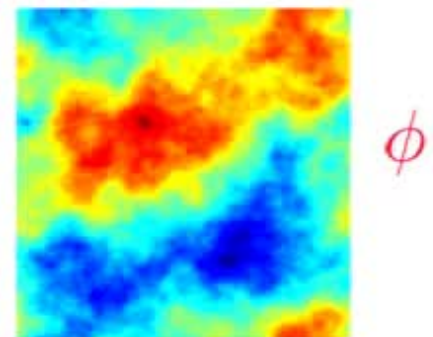
$$\begin{cases} \mathbf{q}_x^i = RW A_i D_x \phi + \mathbf{n}_x^i, \\ \mathbf{q}_y^i = RW A_i D_y \phi + \mathbf{n}_y^i, \end{cases} \quad i = 1, 2, \dots, m.$$

□ Treat ϕ as an “image” and regularize it:

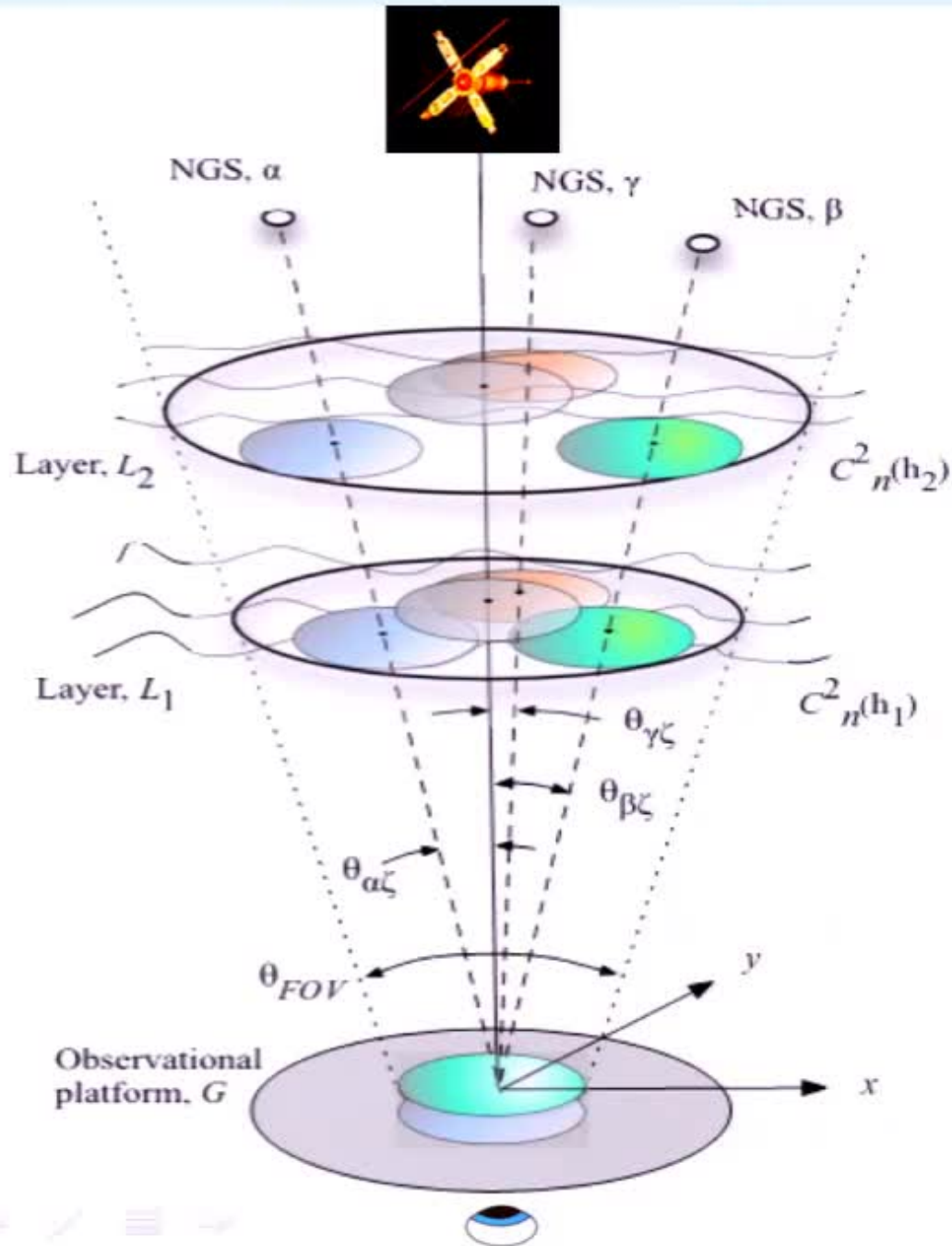
$$\min_{\phi} \|C\phi\|_1 + \frac{\alpha}{2} \sum_{i=1}^m \left\| \begin{bmatrix} RW A_i D_x \\ RW A_i D_y \end{bmatrix} \phi - \begin{bmatrix} \mathbf{q}_x^i \\ \mathbf{q}_y^i \end{bmatrix} \right\|_2^2.$$

□ Regularizer C can be TV, wavelet, tight-frame, fractional, ...

[C., Yuan, & Zhang, Science China, 13]

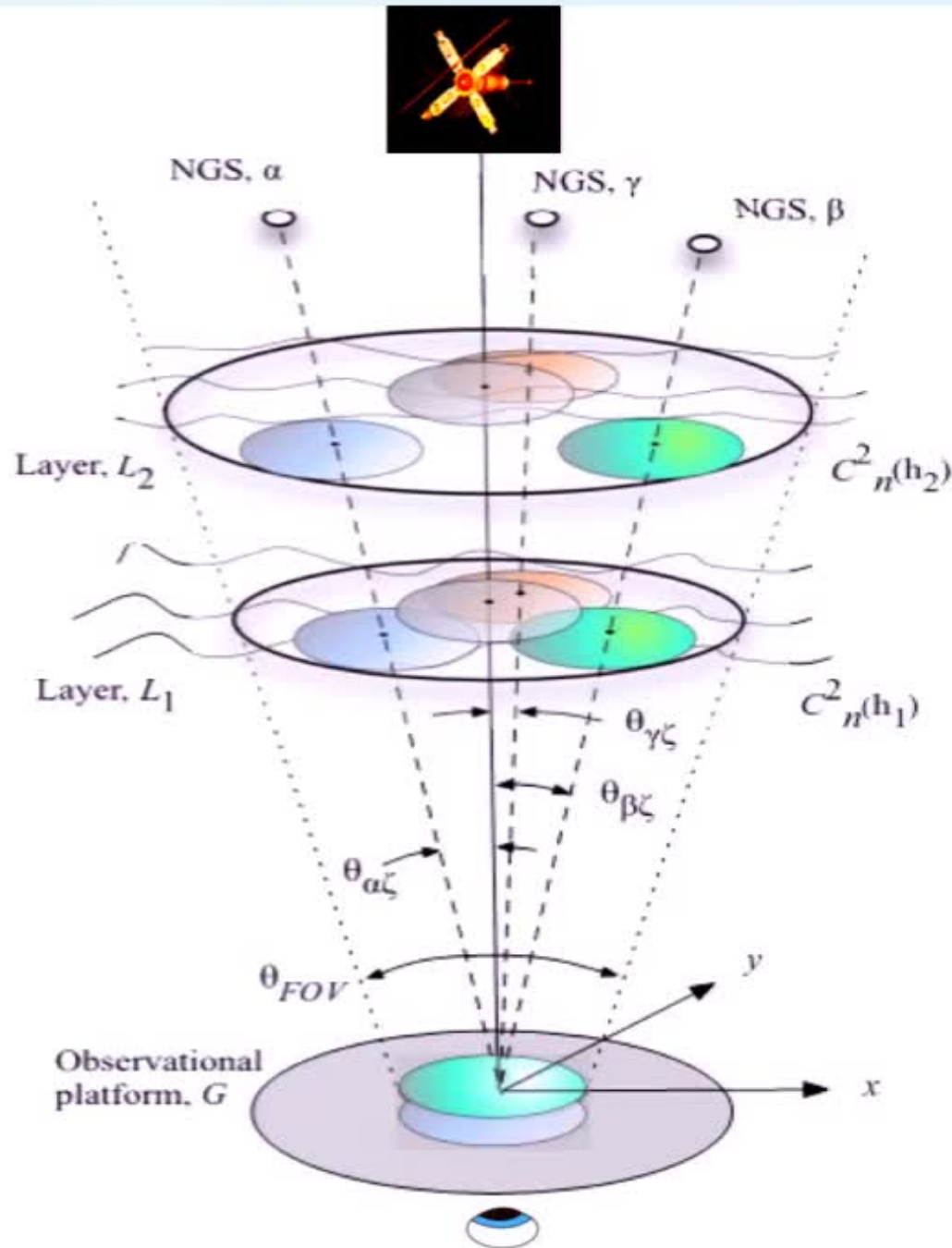


Multi-layered Atmosphere



S.J. Weddell, "Optical wavefront prediction with reservoir computing".

Multi-layered Atmosphere

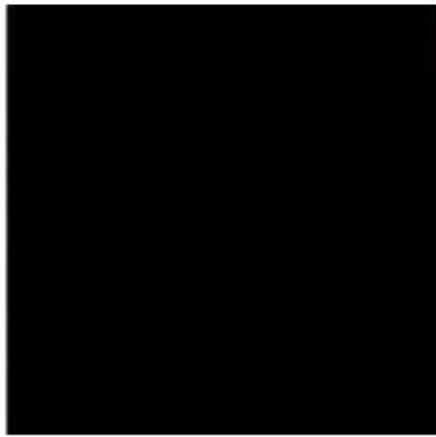


$$\phi_2(x, y)$$

$$\phi_1(x, y)$$

S.J. Weddell, "Optical wavefront prediction with reservoir computing".

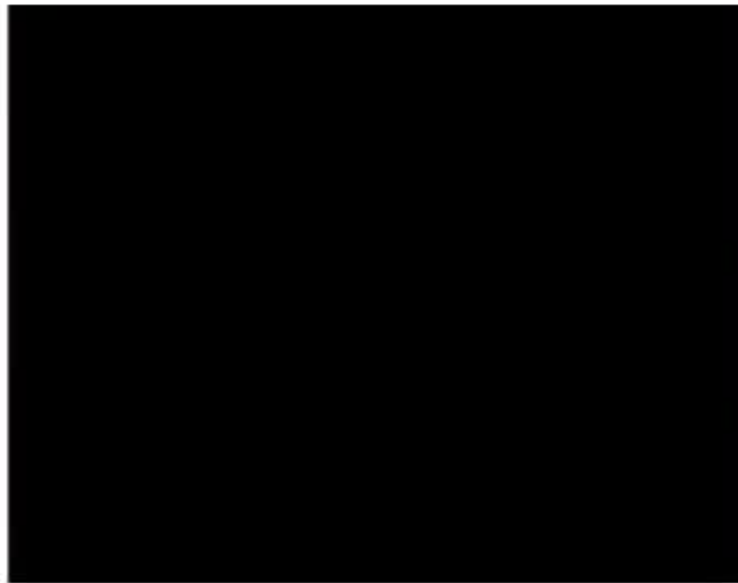
Multi-layered Phase



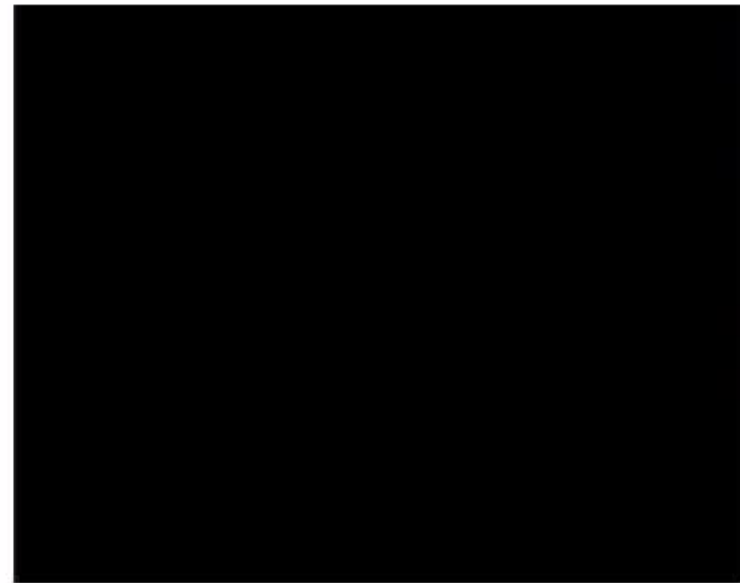
One-layer ϕ



Multi-layer $\{\phi_l\}$



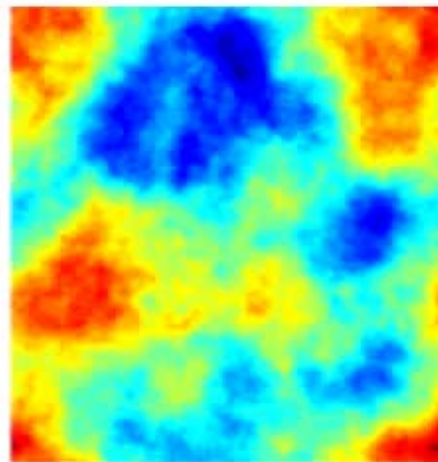
LR Horizontal \mathbf{q}_x



LR Horizontal \mathbf{q}_x

Experiment Setup

- generate true 256-by-256 ϕ [Nagy *et al.*, 10, 13]



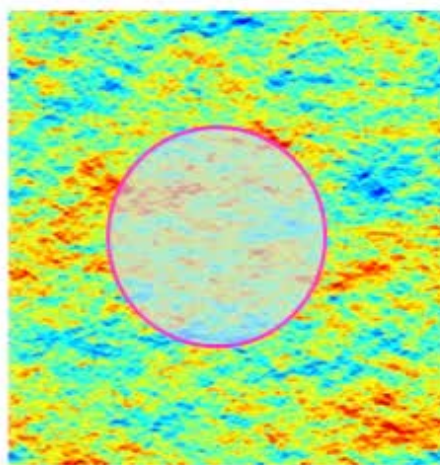
ϕ

Experiment Setup

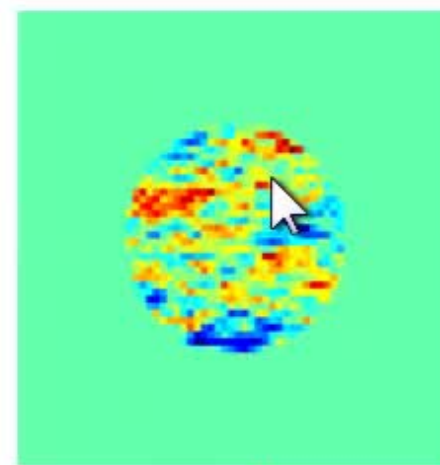
- generate true 256-by-256 ϕ [Nagy *et al.*, 10, 13]
- generate HR $\mathbf{p}_x = D_x\phi$ and $\mathbf{p}_y = D_y\phi$
- generate LR \mathbf{q}_j^i with 1% Gaussian noise by

$$\begin{aligned}\mathbf{q}_x^i &= RW A_i \mathbf{p}_x + \mathbf{n}_x^i, & i &= 1, 2, \dots, m \\ \mathbf{q}_y^i &= RW A_i \mathbf{p}_y + \mathbf{n}_y^i, & i &= 1, 2, \dots, m\end{aligned}$$

- downsample by a factor of 4 (64-by-64 LR)
- use $m = 16$ frames



\mathbf{p}_x



\mathbf{q}_x^i

Reconstructing the PSF

□ Given $\{\mathbf{q}_x^i\}_{i=1}^m$ and $\{\mathbf{q}_y^i\}_{i=1}^m$, solve

□
$$\min_{\mathbf{p}_x} \|\mathbf{p}_x\|_2^2 + \frac{\alpha}{2} \sum_{i=1}^m \|RW A_i \mathbf{p}_x - \mathbf{q}_x^i\|_2^2$$

□
$$\min_{\mathbf{p}_x} \|\mathbf{p}_x\|_1 + \frac{\alpha}{2} \sum_{i=1}^m \|RW A_i \mathbf{p}_x - \mathbf{q}_x^i\|_2^2$$

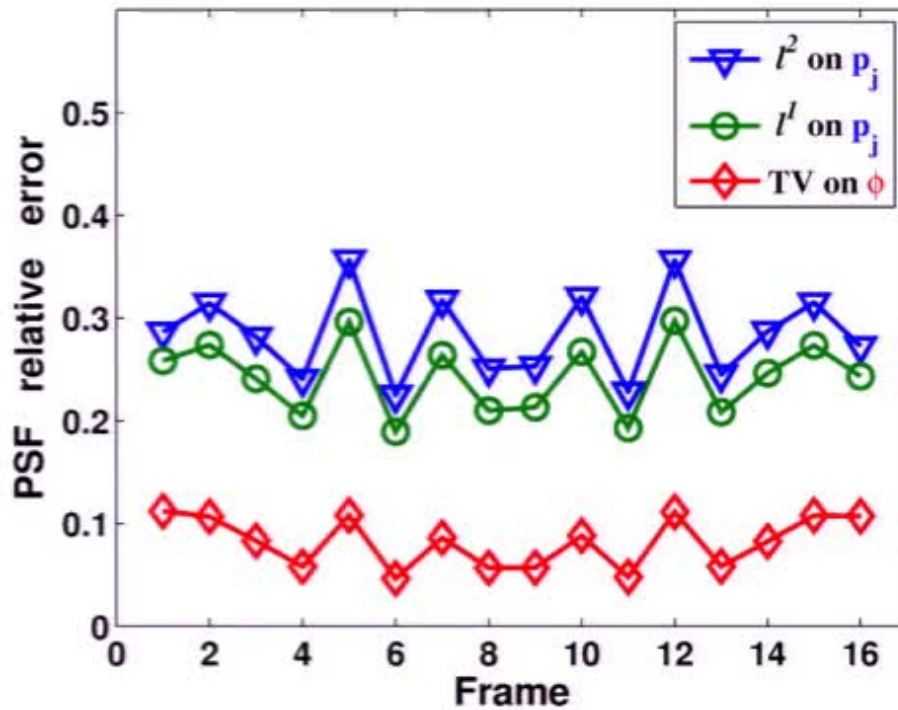
□
$$\min_{\phi} \|\nabla \phi\|_1 + \frac{\alpha}{2} \sum_{i=1}^m \left\| \begin{bmatrix} RW A_i D_x \\ RW A_i D_y \end{bmatrix} \phi - \begin{bmatrix} \mathbf{q}_x^i \\ \mathbf{q}_y^i \end{bmatrix} \right\|_2^2$$

□ Recover ϕ from $\mathbf{p}_x = D_x \phi$ and $\mathbf{p}_y = D_y \phi$

□ Recover PSF from $k(x, y) = |\mathcal{F}^{-1} \{W(x, y) e^{i\phi(x, y)}\}|^2$

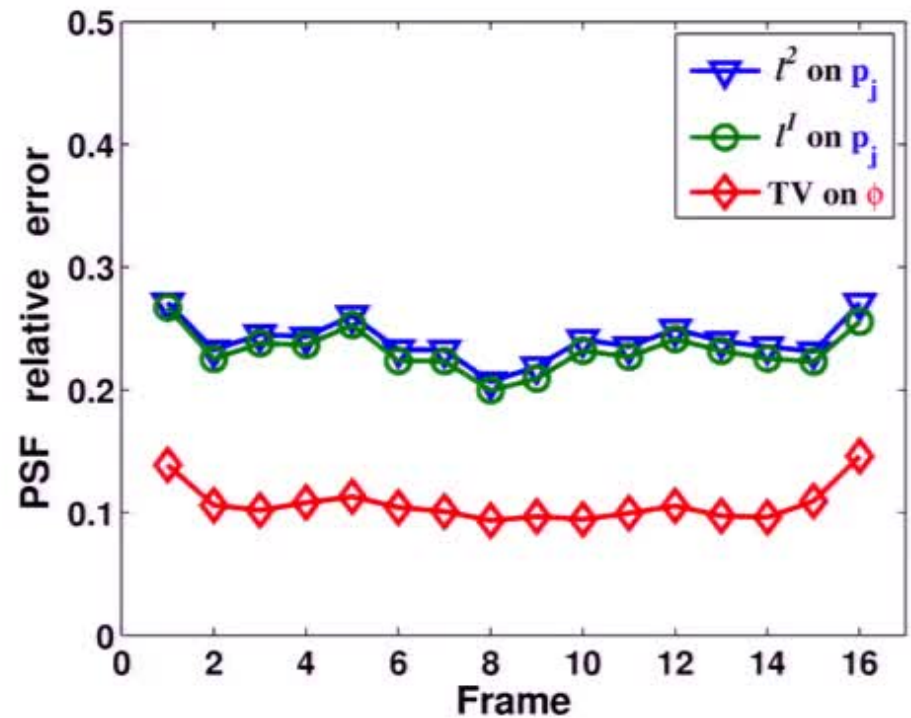
□ Compare computed k_c with true k from true ϕ

PSF Error Comparison



single-layer

seeing condition = 45



3-layer

seeing condition = 20

∇ : [Chu, Jefferies, & Nagy, SIAM J. Sci. Comput., 13]

\circ : [C., Yuan, & Zhang, J. Opt. Soc. Amer. A, 12]

\diamond : [C., Yuan, & Zhang, Science China A., 13]

How good is the Deblurring?

- Use true PSF $k(x, y)$ to generate blurred image:

$$g(x, y) = k(x, y) * f(x, y) + n(x, y)$$

with 1% Gaussian noise added.

- In matrix terminology:

$$\mathbf{g} = \mathbf{K}\mathbf{f} + \mathbf{n}$$

- Deblur \mathbf{g} with computed PSF $k_c(x, y)$:

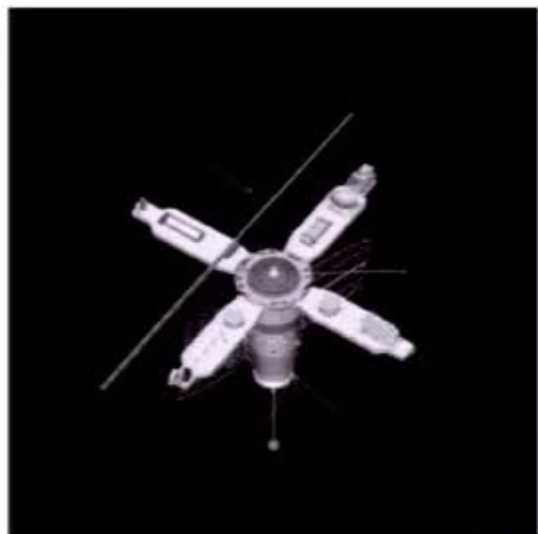
$$\min_{\mathbf{f}} \|\nabla \mathbf{f}\|_1 + \frac{\mu}{2} \|\mathbf{K}_c \mathbf{f} - \mathbf{g}\|_2^2$$



$g(x, y)$

Results for 1-Layer Case

1dB $\uparrow \approx 10\%$ \downarrow in relative error



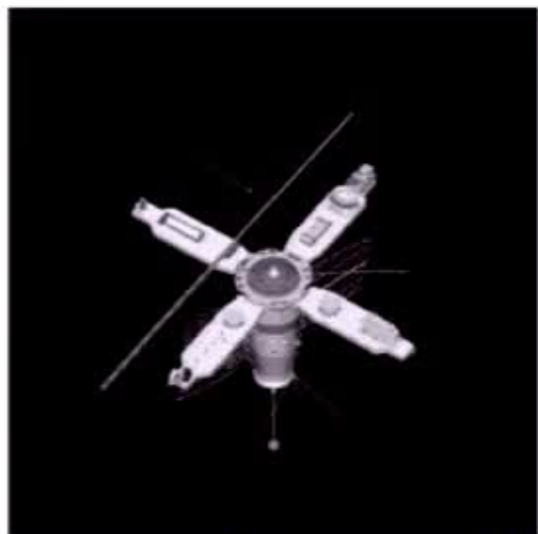
true image f



blurred image g

Results for 1-Layer Case

1dB $\uparrow \approx 10\%$ \downarrow in relative error



true image f



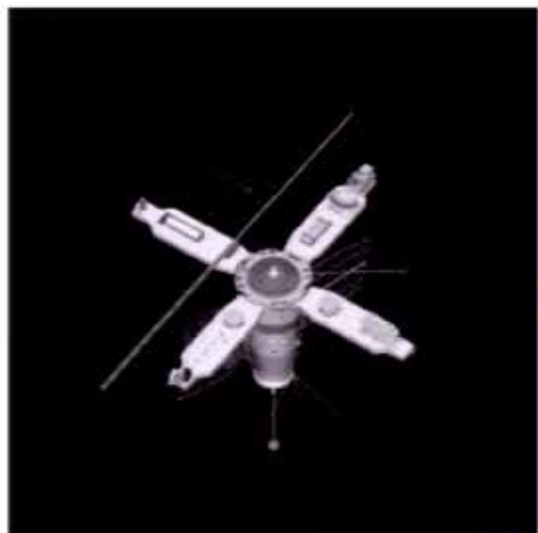
true psf



blurred image g

Results for 1-Layer Case

1dB $\uparrow \approx 10\%$ \downarrow in relative error

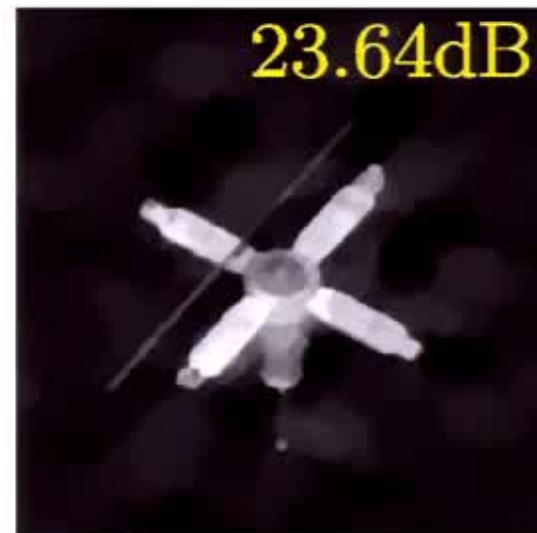


true image f



29.66dB

true psf



23.64dB

ℓ^2 on $\mathbf{p}_x, \mathbf{p}_y$



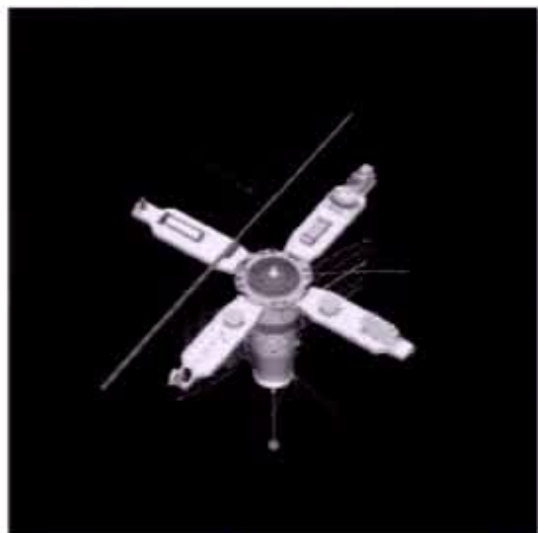
15.21dB

blurred image g



Results for 1-Layer Case

1dB $\uparrow \approx 10\%$ \downarrow in relative error

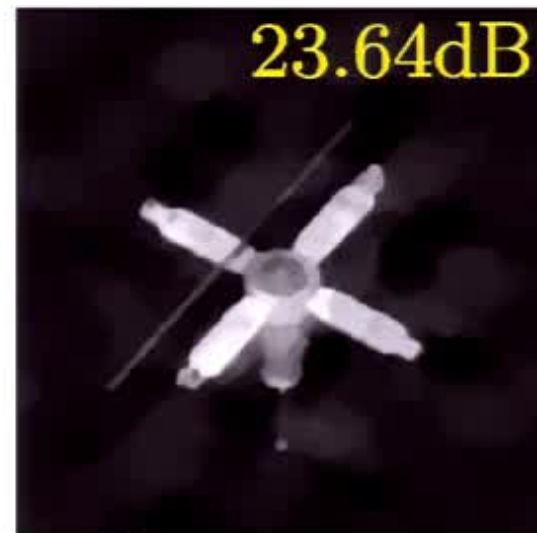


true image f



29.66dB

true psf



23.64dB

ℓ^2 on $\mathbf{p}_x, \mathbf{p}_y$



15.21dB

blurred image g

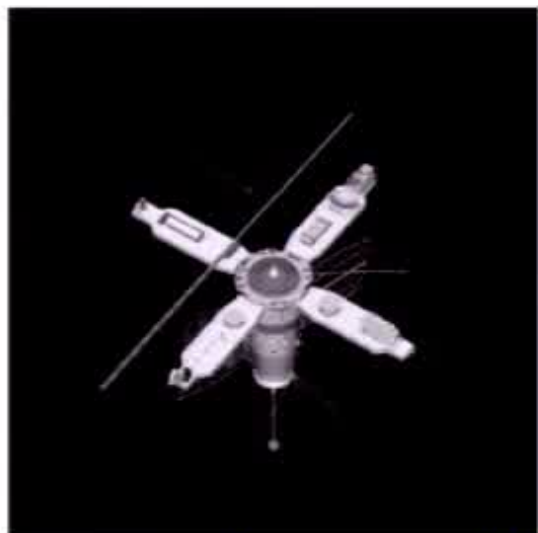


24.59dB

ℓ^1 on $\mathbf{p}_x, \mathbf{p}_y$

Results for 1-Layer Case

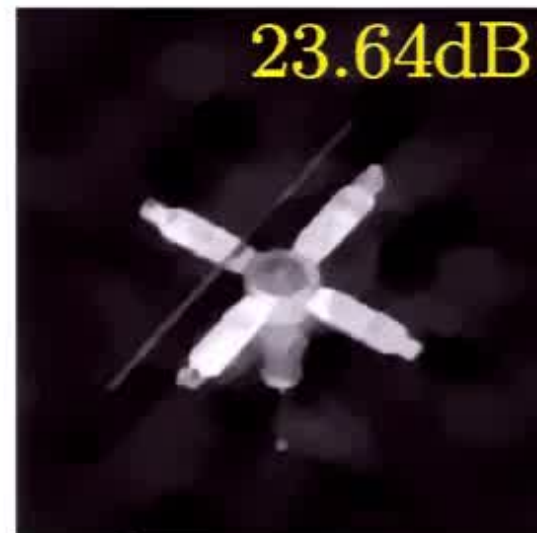
1dB $\uparrow \approx 10\%$ \downarrow in relative error



true image f



true psf



ℓ^2 on p_x, p_y



blurred image g



ℓ^1 on p_x, p_y



TV on ϕ

Results for 3-Layer Case



true image f



blurred image g

Results for 3-Layer Case



true image f



true psf



blurred image g



Results for 3-Layer Case



true image f



true psf



ℓ^2 on $\mathbf{p}_x, \mathbf{p}_y$



blurred image g

Results for 3-Layer Case



true image f



true psf



ℓ^2 on $\mathbf{p}_x, \mathbf{p}_y$



blurred image g



ℓ^1 on $\mathbf{p}_x, \mathbf{p}_y$

Results for 3-Layer Case



true image f



34.57dB

true psf



31.78dB

ℓ^2 on $\mathbf{p}_x, \mathbf{p}_y$



21.83dB

blurred image g



31.89dB

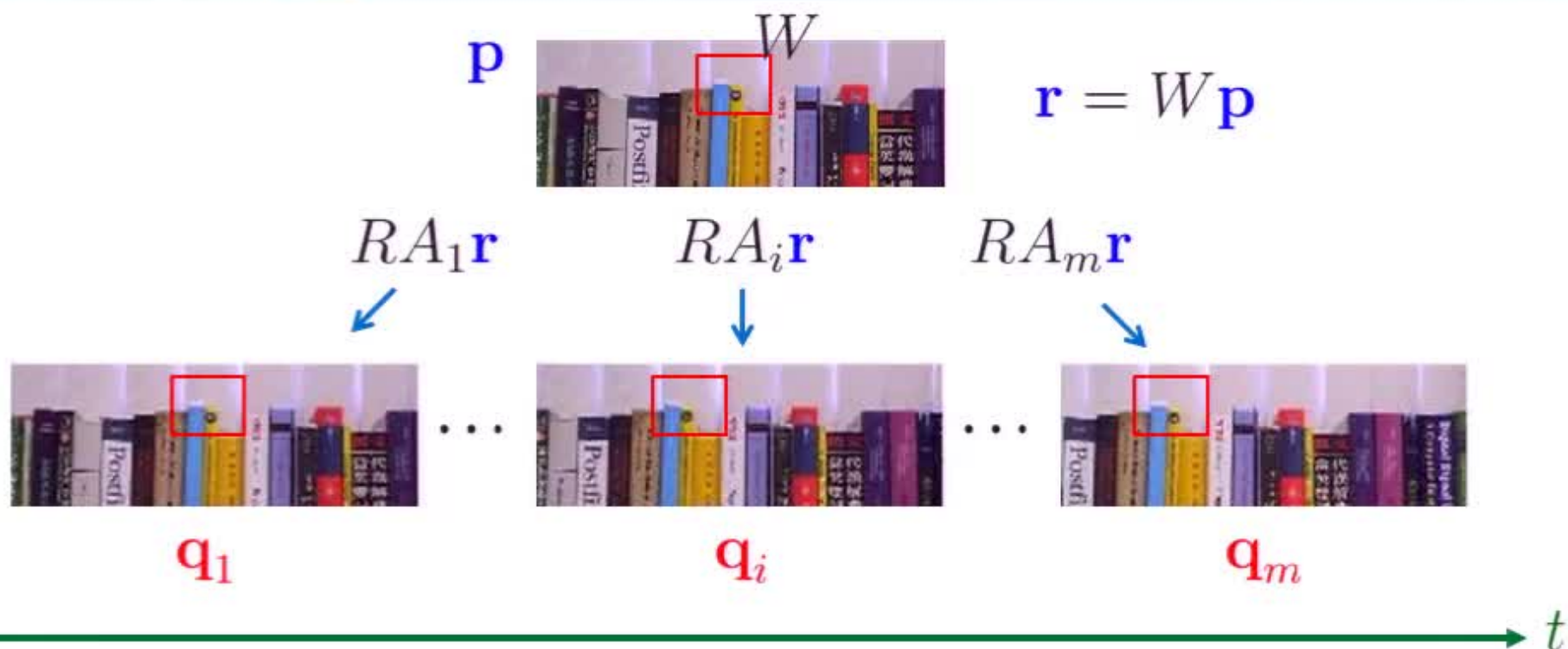
ℓ^1 on $\mathbf{p}_x, \mathbf{p}_y$



33.89dB

TV on ϕ

Classical Approach [Tsai & Huang, 84]

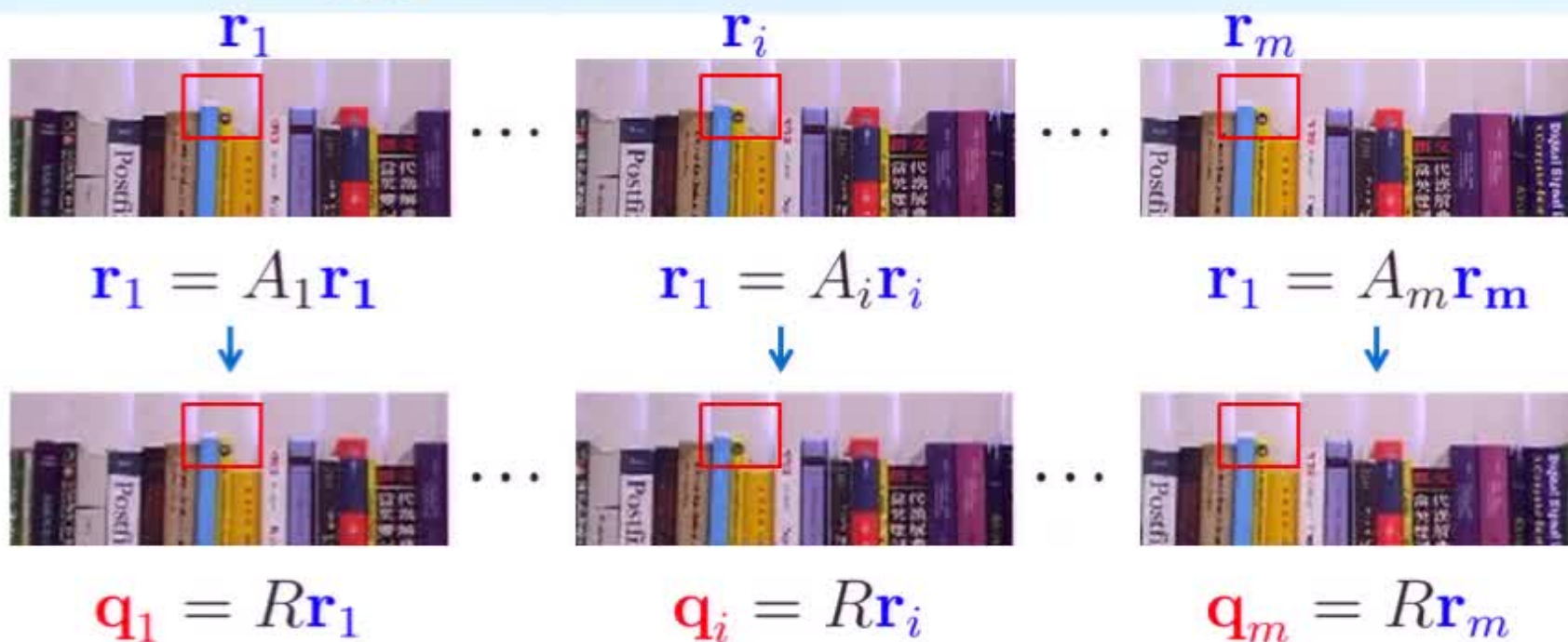


□ $\mathbf{q}_i = RBA_i\mathbf{r} + \mathbf{n}_i, i = 1, \dots, m$

□ Solve, e.g.

$$\min_{\mathbf{r}} \|\nabla\mathbf{r}\|_1 + \frac{\alpha}{2} \sum_{i=1}^m \|RA_i\mathbf{r} - \mathbf{q}_i\|_2^2.$$

Low-rank Approach



□ $\mathbf{q}_i = R\mathbf{r}_i + \mathbf{n}_i, i = 1, \dots, m.$

□ $[A_1 \mathbf{r}_1, \dots, A_m \mathbf{r}_m]$ low rank, so:

$$\min_{\mathbf{r}_i} \text{rank}[A_1 \mathbf{r}_1, \dots, A_m \mathbf{r}_m] + \frac{\alpha}{2} \sum_{i=1}^m \|R\mathbf{r}_i - \mathbf{q}_i\|_2^2.$$

Nuclear Norm [Candes, Recht, 09; Recht, Fazel, Parrilo, 10]

$$\min_{\mathbf{r}_i} \|[A_1 \mathbf{r}_1, \dots, A_m \mathbf{r}_m]\|_* + \frac{\alpha}{2} \sum_{i=1}^m \|R \mathbf{r}_i - \mathbf{q}_i\|_2^2.$$

- Nuclear norm: $\|U\|_* = \sum_j \sigma_j(U) = \|\boldsymbol{\sigma}(U)\|_1$.
- An ℓ^1 - ℓ^2 model. Can be solved by ADMM:
 - Auxiliary variables: $\mathbf{v}_i = A_i \mathbf{r}_i$.
 - \mathbf{r}_i -subproblem: $(\alpha R^t R + \beta A_i^t A_i) \mathbf{r}_i^{j+1} = \mathbf{b}^j$.
 - \mathbf{v}_i -subproblem: $[\mathbf{v}_1^{j+1}, \dots, \mathbf{v}_m^{j+1}] = \text{SVS}_{\frac{1}{\beta}}(U^j)$.

Comparison



*Single frame with
bilinear interpolation*

Comparison



*Single frame with
bilinear interpolation*



*21 frames with
nuclear norm*



*21 frames with
TV regularization*



*21 frames with
tightframe approach*

HS Image Reconstruction by Tightframe

- Not classical approach:

$$\min_{\mathbf{r}} \|\mathcal{T}\mathbf{r}\|_1 + \frac{\alpha}{2} \sum_{i=1}^m \|RA_i\mathbf{r} - \mathbf{q}_i\|_2^2.$$

- Tightframes are generalization of wavelets. Explain idea by the simplest tightframe, the Haar wavelet.
- Consider the simplest case: 4 low-resolution images merge into 1 high-resolution image:

HS Image Reconstruction by Tightframe

- **Not** classical approach:

$$\min_{\mathbf{r}} \|\mathcal{T}\mathbf{r}\|_1 + \frac{\alpha}{2} \sum_{i=1}^m \|RA_i\mathbf{r} - \mathbf{q}_i\|_2^2.$$

- Tightframes are generalization of wavelets. Explain idea by the simplest tightframe, the Haar wavelet.
- Consider the simplest case: 4 **low-resolution** images merge into 1 **high-resolution** image:

LR images \mathbf{q}
align exactly at
half-pixel

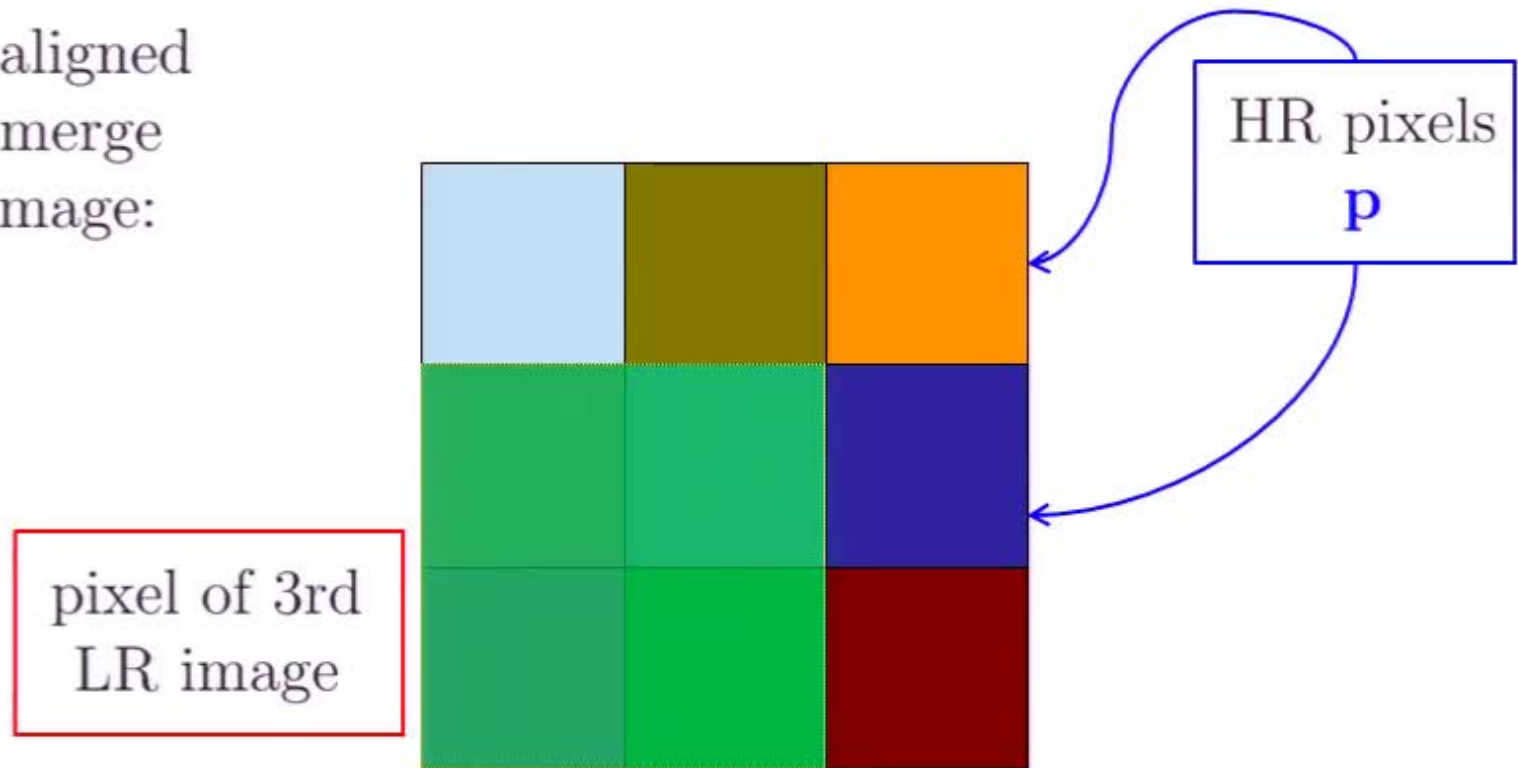


High-resolution
 \mathbf{p}



The Process from HR to LR

4 perfectly aligned
LR images merge
into 1 HR image:



□ HR \rightarrow LR process = convolution (blurring) with kernel

$$\left(\dots, 0, \frac{1}{2}, \frac{1}{2}, 0, \dots \right) \otimes \left(\dots, 0, \frac{1}{2}, \frac{1}{2}, 0, \dots \right)$$

□ $\left[\frac{1}{2}, \frac{1}{2} \right]$ is Haar's low-pass filter

Frequency Domain Inpainting

HR image reconstruction = frequency domain inpainting



p

\mathcal{T}

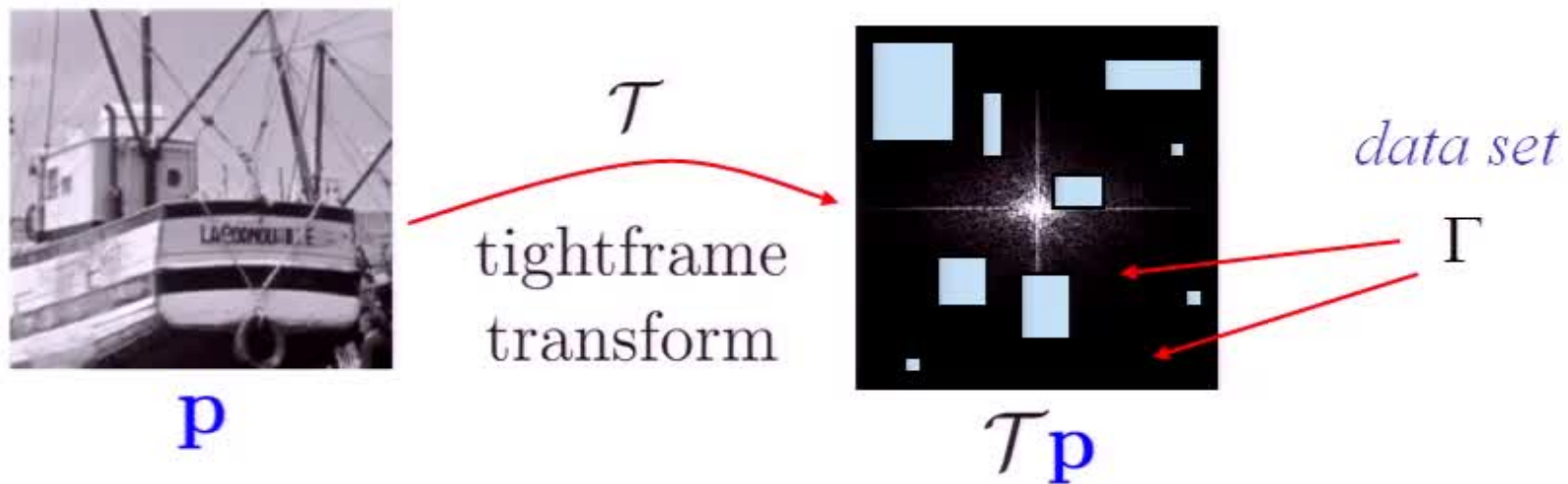


tightframe
transform

Frequency Domain Inpainting

HR image reconstruction = frequency domain inpainting

- The low-pass framelet coefficients \mathbf{q} are given at locations Γ .
- Find \mathbf{p} such that $P_{\Gamma}\mathcal{T}\mathbf{p} = P_{\Gamma}\mathbf{q}$.



Tightframe Algorithm for Inpainting

For $j = 0, 1, \dots$, until convergence:

1. Compute $\mathbf{c}^j = \mathcal{T}\mathbf{p}^j$.

2. Data fitting: set

$$[\mathbf{c}_d^j]_l = \begin{cases} [\mathbf{q}]_l, & l \in \Gamma \\ [\mathbf{c}^j]_l, & l \notin \Gamma. \end{cases}$$

3. Denoise \mathbf{c}_d^j by shrinkage to get \mathbf{c}_t^j .

4. Reconstruct $\mathbf{p}^{j+1} = \mathcal{T}^t \mathbf{c}_t^j$.

Note that for tightframes, we have $\mathcal{T}^t \mathcal{T} = I$.

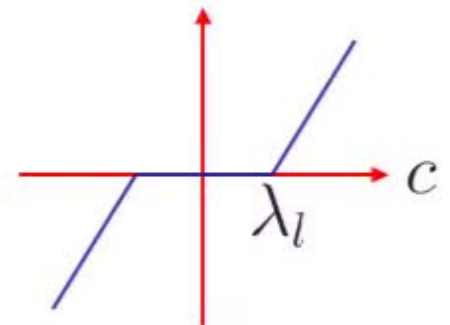
Tightframe Algorithm for Inpainting

The tight-frame frequency domain inpainting algorithm:

$$\mathbf{p}^{j+1} = \mathcal{T}^t \mathcal{S}_\lambda (P_{\Gamma^c} \mathcal{T} \mathbf{p}^j + P_\Gamma \mathbf{q})$$

where

- P_{Γ^c} : projection onto complement of Γ
- \mathcal{S}_λ : shrinkage operator with threshold λ



$$s_{\lambda_l}(c) \equiv \begin{cases} \text{sgn}(c)(|c| - \lambda_l), & \text{if } |c| > \lambda_l, \\ 0, & \text{if } |c| \leq \lambda_l. \end{cases}$$

16-to-1
sensor
array



16 LR images

16-to-1
sensor
array



16 LR images



recovered

Piecewise linear tightframe is used in our examples.

16-to-1
sensor
array



16 LR images



recovered

only
4 LR
images
given



4 LR images



recovered

Piecewise linear tightframe is used in our examples.

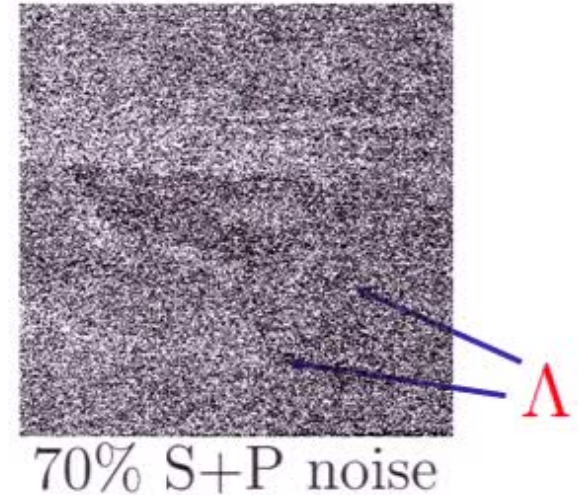
Iterative Thresholding Algorithms

- [$C^2 + S^2$, SISC, 03]:
Proximal forward-backward algorithm
- [Daubechies, Defrise, & De Mol, CPAM 04]
Iterative thresholding with sparsity constraint
- [Elad, Starck, Querre & Donoho, MCA 05]
Simultaneous cartoon and texture image inpainting
- [Combettes & Wajs, SIAM MMS, 05]
Signal recovery by proximal forward-backward splitting
- [Beck & Teboulle, SIIMS 09]
A fast iterative shrinkage-thresholding algorithm (FISTA)
- ...

Inpainting in Image Domain

Find image \mathbf{p} from data \mathbf{q} given on Λ ,
i.e $\mathcal{P}_\Lambda \mathbf{p} = \mathcal{P}_\Lambda \mathbf{q}$.

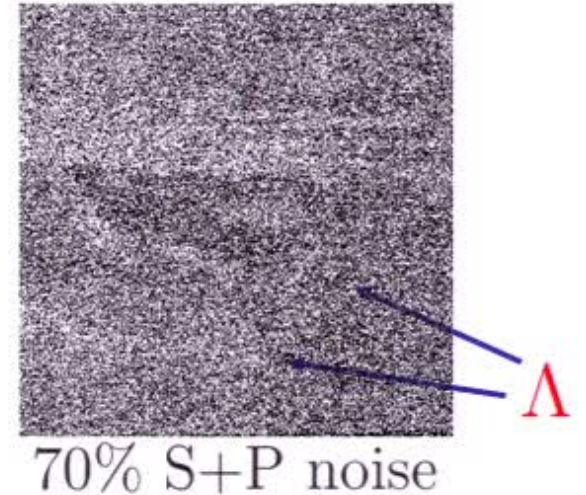
$$\mathbf{p}^{j+1} = \mathcal{P}_{\Lambda^c} \mathcal{T}^t \mathcal{S}_\lambda \mathcal{T} \mathbf{p}^j + \mathcal{P}_\Lambda \mathbf{q}$$



Inpainting in Image Domain

Find image \mathbf{p} from data \mathbf{q} given on Λ ,
i.e $\mathcal{P}_\Lambda \mathbf{p} = \mathcal{P}_\Lambda \mathbf{q}$.

$$\mathbf{p}^{j+1} = \mathcal{P}_{\Lambda^c} \mathcal{T}^t \mathcal{S}_\lambda \mathcal{T} \mathbf{p}^j + \mathcal{P}_\Lambda \mathbf{q}$$

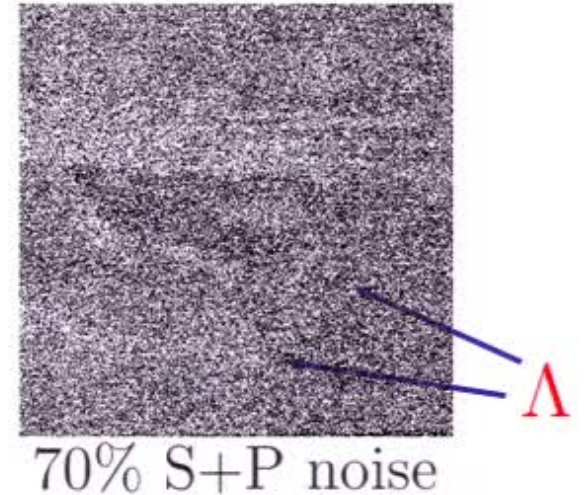


Adaptive Median Filter
(AMF)

Inpainting in Image Domain

Find image \mathbf{p} from data \mathbf{q} given on Λ ,
i.e $\mathcal{P}_\Lambda \mathbf{p} = \mathcal{P}_\Lambda \mathbf{q}$.

$$\mathbf{p}^{j+1} = \mathcal{P}_{\Lambda^c} \mathcal{T}^t \mathcal{S}_\lambda \mathcal{T} \mathbf{p}^j + \mathcal{P}_\Lambda \mathbf{q}$$



Adaptive Median Filter
(AMF)

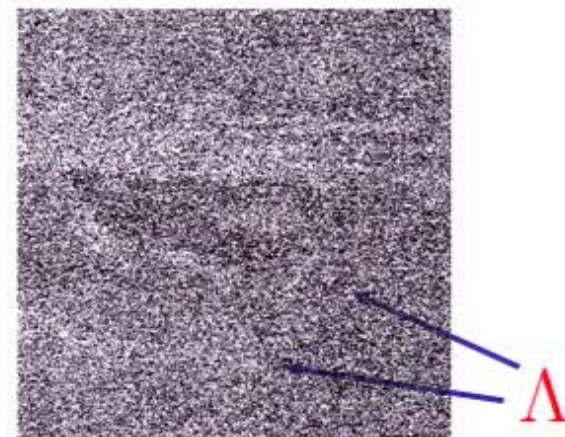


AMF + ℓ^1 - ℓ^1
[C., Ho, Nikolova, 05]

Inpainting in Image Domain

Find image \mathbf{p} from data \mathbf{q} given on Λ ,
i.e $\mathcal{P}_\Lambda \mathbf{p} = \mathcal{P}_\Lambda \mathbf{q}$.

$$\mathbf{p}^{j+1} = \mathcal{P}_{\Lambda^c} \mathcal{T}^t \mathcal{S}_\lambda \mathcal{T} \mathbf{p}^j + \mathcal{P}_\Lambda \mathbf{q}$$



70% S+P noise



Adapative Median Filter
(AMF)



AMF + ℓ^1 - ℓ^1
[C., Ho, Nikolova, 05]



AMF + tightframe
[Cai, C., Shen, Shen, 09]

Single Frame Upsampling

HR video from
LR video

Single Frame Upsampling

HR video from
LR video



Single Frame Upsampling



interpolate/inpaint



HR video from
LR video



Input Video



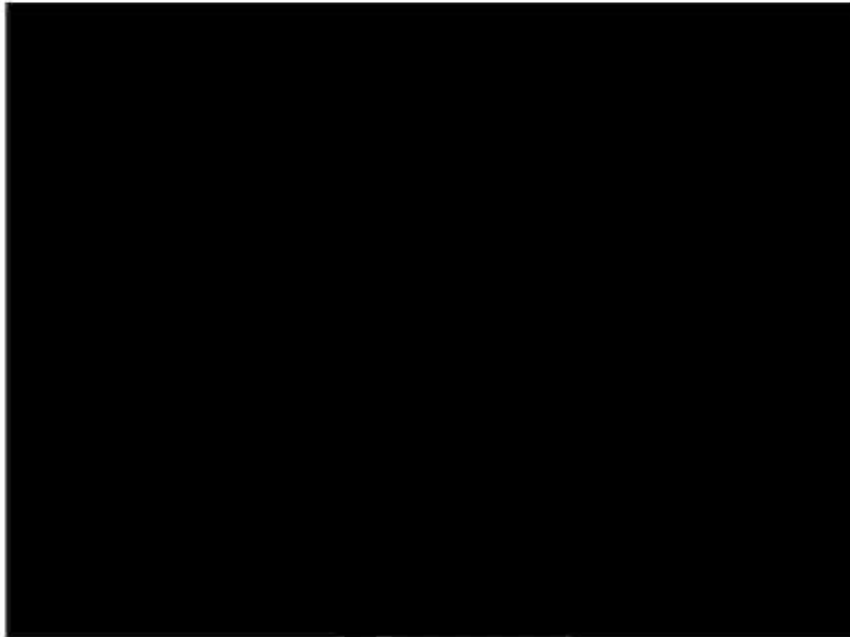
Qian *et al.* Siggraph 09: $\|\Psi(\nabla \mathbf{f})\|_1$



Upsampled by bicubic



Level 6 Tightframe



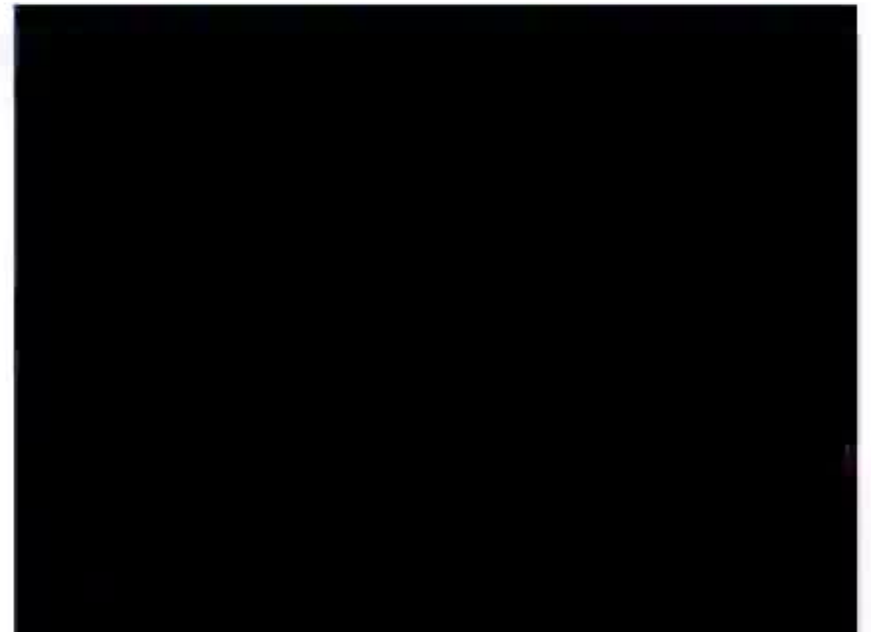
Input Video



Qian *et al.* Siggraph 09: $\|\Psi(\nabla \mathbf{f})\|_1$



Upsampled by bicubic



Level 6 Tightframe

Thanks to the Collaborators

- Jianfeng Cai (HK University of Science & Technology)
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- Xiaoming Yuan (HK Baptist University)
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Thank you for your attention!

Welcome to SIAM LA18 in Hong Kong

May 4-8, 2018



View from the Peak

Welcome to SIAM LA18 in Hong Kong

May 4-8, 2018



View from the Peak

Welcome to SIAM LA18 in Hong Kong

Salty Field Bay



May 4-8, 2018



View from the Peak

Chinese University of Hong Kong



Concluding Remarks

- ℓ^1 - ℓ^2 problems are common in image processing
- Efficient solvers by adding auxiliary variables
- Solution requires solving a linear system for every outer iteration
- Some requires computing an SVD for every outer iteration



Input Video



Qian *et al.* Siggraph 09: $\|\Psi(\nabla \mathbf{f})\|_1$



Upsampled by bicubic



Level 6 Tightframe