

Data assimilation for chaotic geophysical dynamics

A very brief overview

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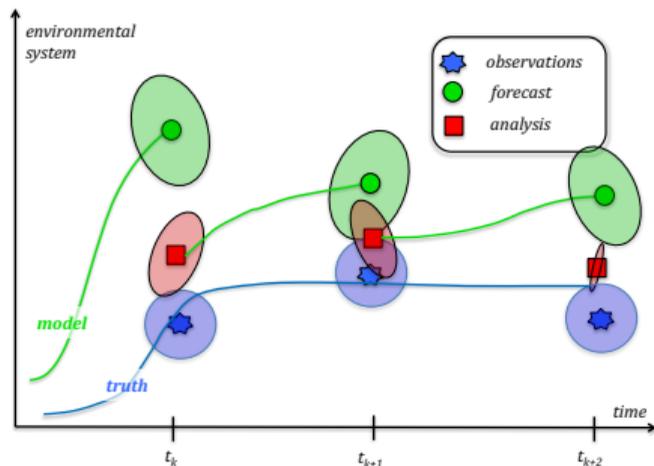
(4): CEREAs, joint lab École des Ponts ParisTech and EdF R&D, IPSL, France



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Data Assimilation: *What and Why?*



- ▶ Large family of methods to perform state/parameter/model estimation by combining (taking the best of) models and data.
- ▶ The quantities of interest are **probability density functions** (PDFs).
- ▶ They quantify the uncertainty on the estimate.
- ▶ The goal is to sequentially estimate the conditional PDF $p(\mathbf{x}|\mathbf{y})$, the posterior.
- ▶ The PDFs are evolved in time and updated at analysis times using **Bayes' rule**.

$$p(\mathbf{x}_{K:0}|\mathbf{y}_{K:1}) \propto p(\mathbf{x}_0) \prod_{k=1}^K p(\mathbf{y}_k|\mathbf{x}_k)p(\mathbf{x}_k|\mathbf{x}_{k-1})$$

DA methods in geosciences: *key challenges*

- ▶ Huge models, $m \geq 10^9$ & massive dataset, $d \geq 10^7$ daily obs (yet not enough!) \implies Quest for computationally affordable solutions.
- ▶ Fully Bayesian DA very difficult (*curse of dimensionality*).
- ▶ Gaussian/Linear hypotheses allow to derive computationally tractable methods \implies **Kalman filter/smoothen**.
- ▶ And their Monte Carlo, (still Gaussian) nonlinear approx, **ensemble Kalman filter/smoothen**.
- ▶ The transition density, $p(\mathbf{x}_k | \mathbf{x}_{k-1})$, approximated by running an ensemble of N trajectories.
- ▶ EnKF highly rank deficient in geophysical applications: $N = \mathcal{O}(10^2) \ll m$. Yet it works!

DA methods for chaotic dynamics: *key challenges*

- ▶ Atmosphere and ocean, are examples of chaotic dissipative dynamics \implies Highly state-dependent error growth.
- ▶ DA must track and incorporate this flow-dependency in the quantification of the uncertainty (*i.e.* error covariance).
- ▶ Dissipation induces an “effective” dimensional reduction \implies The error dynamics is confined to a subspace of much smaller dimension, $n_0 \ll m$: the **unstable subspace**
- ▶ The existence of the underlying *unstable-stable splitting of the phase space* expected to have enormous impact on DA.

Motivations

- 1 Is there any fingerprint of the unstable subspace on the fate of (En)KF and (En)KS?
- 2 Can dynamical properties be used to design computationally cheap DA strategies?

Deterministic linear case: *behavior of the KF and KS*

(Some) key **analytic results** (without controllability):

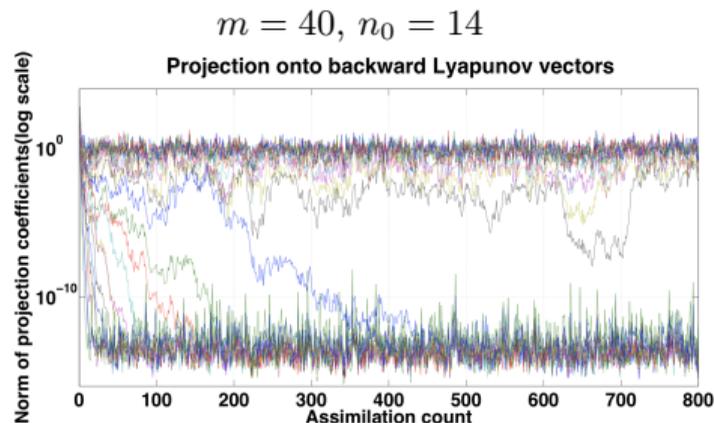
► **Collapse of the uncertainty**: KF error covariance asymptotically in the span of the unstable-neutral backward Lyapunov vectors (BLVs^u) [Gurumoorthy *et al* 2017]

► **Convergence of the covariance**: Low rank, n_0 , KF covariance, initialized in the span of BLVs^u, converges to the true KF one

$$\lim_{k \rightarrow \infty} \|\mathbf{P}_k - \hat{\mathbf{P}}_k\| = 0$$

if the unstable-neutral subspace is observed [Bocquet *et al* 2017]. **Warning**: *neutral modes are tricky!*

► Likewise demonstrated for Kalman smoother [Bocquet & Carrassi 2017].

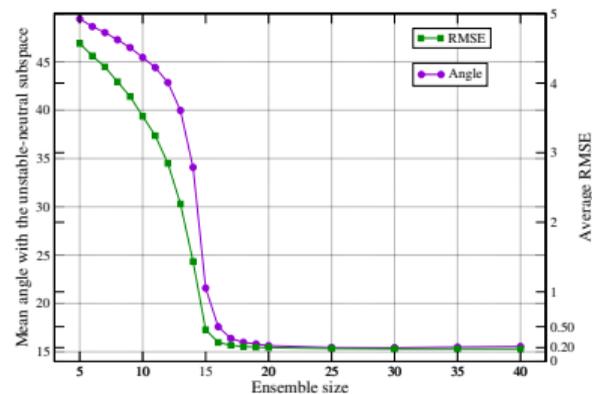
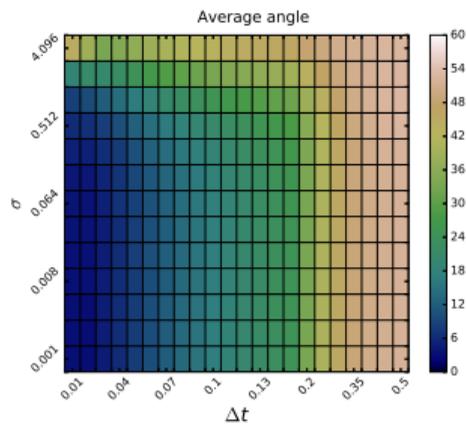


KF/KS reduced rank surrogates based on BLVs are possible.

Deterministic nonlinear case: *behavior of the EnKF and EnKS*

- ▶ Asymptotic rank of EnKF covariances related to multiplicity and strength of unstable Lyapunov exponents (LEs) [Carrassi *et al* 2009; Gonzalez-Tokman & Hunt 2013].
- ▶ When the EnKF/EnKS ensemble subspace recovers the unstable subspace the unknown system state is estimated with high accuracy (sudden drop of RMSE) [Bocquet & Carrassi, 2017].

- Lorenz 96 model, $m = 40$, $n_0 = 14$
- **Left** - Angle Unstable/Ensemble subspaces *vs* $(\Delta t^{obs}, \sigma^{obs})$.
- **Right** - EnKF RMSE (green) and Angle (purple) *vs* N .



Nonlinear systems, with “weakly nonlinear” error dynamics, need only n_0 members!

Error in stochastic models: *What the role of the instabilities?*

$$\mathbf{x}_k = \mathcal{M}_{k:k-1}(\mathbf{x}_{k-1}) + \boldsymbol{\eta}_k, \quad \boldsymbol{\eta}_k \in \mathcal{N}(\mathbf{0}, \mathbf{Q}_k)$$

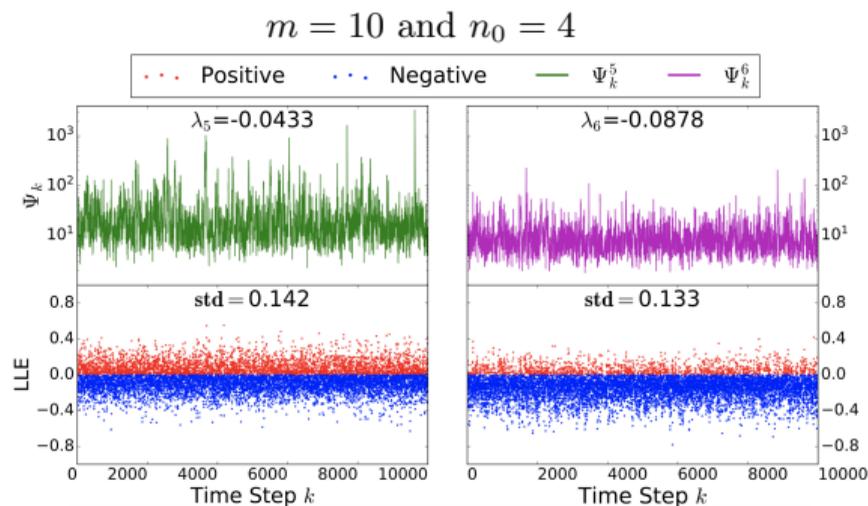
► By re-introducing perturbations, the error covariance is generally full rank (as \mathbf{Q}_k).

► Asymptotic uncertainty in the stable BLVs **no longer zero, but still bounded**.

► However, the bounds, Ψ_k^i , depend on [Grudzien *et al* 2018a]

- 1 the model error size (*i.e.* $\|\mathbf{Q}\|$),
- 2 the variance of the local LEs (LLEs).

► If model error is large and/or the LLEs have high variance, the bounds will be impractically large.



In stochastic systems it is *necessary* to include weakly stable BLVs of high variance.

Error in stochastic models: *The upwelling effect*

- ▶ Will the *necessary* increase $\mathbf{N} = \mathbf{n}_0 \rightarrow \mathbf{n}_0 + \mathbf{n}_{\text{ws}}$ also be *sufficient*?
- ▶ To answer this, write the model propagator in the basis of the BLVs using the recursive QR decomposition

$$\mathbf{M}_k = \mathbf{E}_k \mathbf{U}_k \mathbf{E}_k^T, \quad \mathbf{E}_k = (\mathbf{E}_k^f \ \mathbf{E}_k^u) \quad \text{with} \quad \mathbf{U}_k = \begin{pmatrix} \mathbf{U}_k^{\text{ff}} & \mathbf{U}_k^{\text{fu}} \\ 0 & \mathbf{U}_k^{\text{uu}} \end{pmatrix}$$

and partition the error into **filtered/unfiltered** variables $\epsilon_k = \mathbf{E}_k^f \epsilon_k^f + \mathbf{E}_k^u \epsilon_k^u$

- ▶ The error in the filtered space (“seen” by DA) is given recursively by [Grudzien *et al* 2018b]

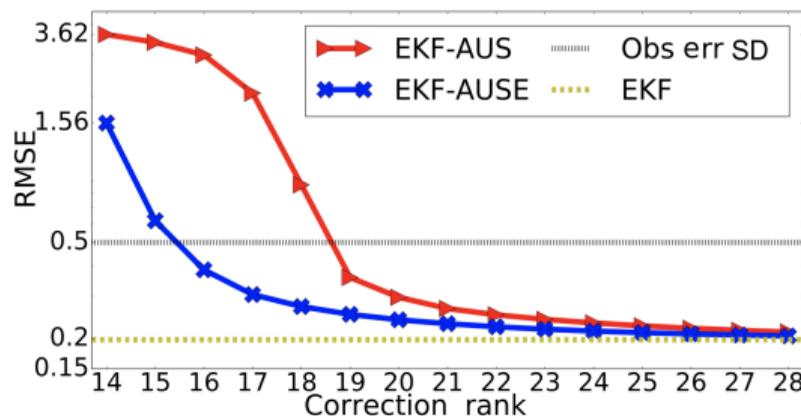
$$\epsilon_{k+1}^f = (\mathbf{U}_{k+1}^{\text{ff}} - \mathbf{U}_{k+1}^{\text{ff}} \mathbf{K}_k \mathbf{H}_k \mathbf{E}_k^f) \epsilon_k^f - \mathbf{U}_{k+1}^{\text{ff}} \mathbf{K}_k \epsilon_k^{\text{obs}} + \boldsymbol{\eta}_k^f + (\mathbf{U}_{k+1}^{\text{fu}} - \mathbf{U}_{k+1}^{\text{ff}} \mathbf{K}_k \mathbf{H}_k \mathbf{E}_k^u) \epsilon_k^u$$

- ▶ **The terms in black** correspond to the usual KF-like recursion and highlight the stabilizing effect of the DA [Carrassi *et al* 2008b].

- ▶ **The terms in red** disappear when the filtered subspace is the entire state space ($n = m$).

Error in stochastic models: *The upwelling effect*

- ▶ When $n < m$, they represent the **dynamical upwelling** of the unfiltered error into the filtered variables [Grudzien *et al* 2018b].
- ▶ It moves uncertainty from unfiltered to filtered subspace, *i.e.* from the stabler to the unstable subspace.
- ▶ This phenomenon **occurs whenever** $n < m$, but is **exacerbated by stochastic noise**.
- ▶ Leads to underestimating the error in the (En)KF \Rightarrow Need for **inflation** to prevent divergence.



- **EKF** solves the *full-rank* recursion.
- **EKF-AUS** solves the *low-rank* recursion without upwelling (black terms only).
- **EKF-AUSE** solves the *low-rank* recursion with upwelling (black+red terms).

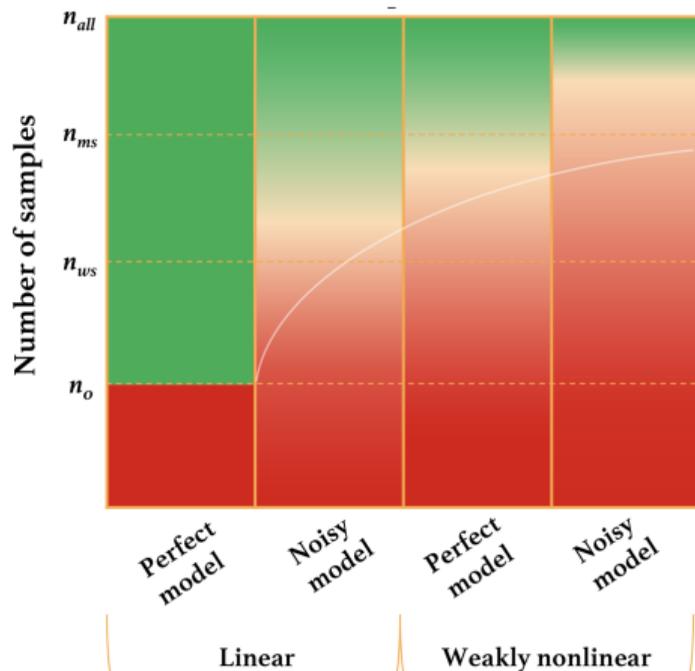
EnKF/KS with chaotic systems: *A summary on the required ensemble size*

► Illustration of the minimum number of ensemble members to achieve filter accuracy.

► Different **model scenarios** are given in the *x-axis*.

► The number of members (samples) is given in the *y-axis* with:

- n_0 : number of unstable-neutral BLVs.
- n_{ws} : number of unstable-neutral BLVs + weakly stable.
- n_{ms} : number of unstable-neutral BLVs + weakly stable + more stable.
- n_{all} : number of member = model dimension (*i.e.* full-rank filter).



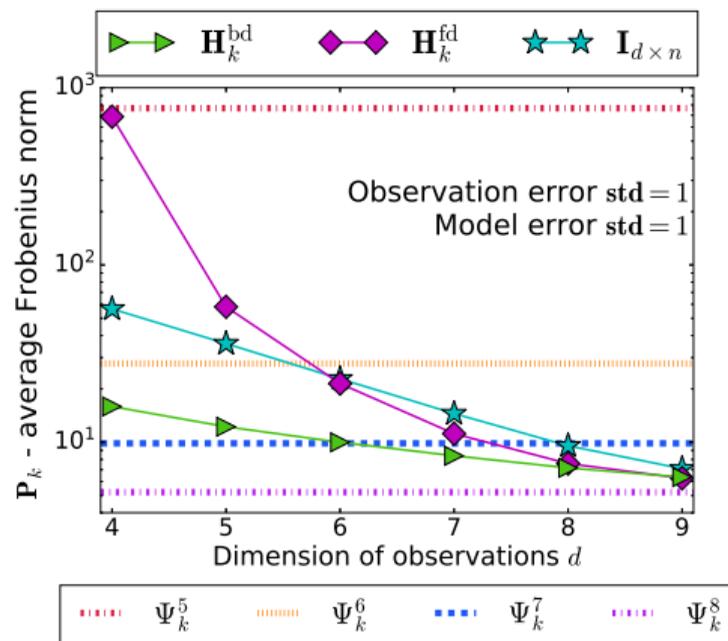
Grudzien *et al* 2018b

Assimilation in the unstable subspace - AUS

- These properties are at the basis of the **assimilation in the unstable subspace** (**AUS**, by *A. Trevisan & Collaborators*), where the unstable subspace is explicitly used in the DA to:
- parametrize the description (both temporally and spatially) of the uncertainty in the state estimate (*i.e.* the covariance) \iff Acting on \mathbf{K} [Trevisan *et al* 2010; Trevisan & Palatella 2011; Palatella & Trevisan 2015]
 - design of the observational network (types, distribution, frequency) \iff Acting on the operator \mathcal{H} [Trevisan & Uboldi, 2004; Carrassi *et al* 2007]
 - or both [Carrassi *et al* 2008a; 2008b]

Projecting the data in the unstable subspace - Conditions for filter stability

KF with data projected on a subspace of dimension d compared to a full KF.



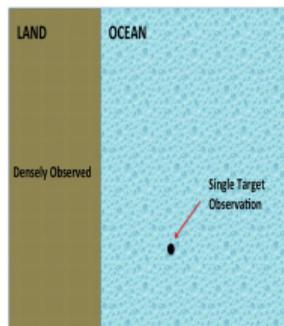
[Grudzien *et al* 2018a]

- ▶ \mathbf{H}_k^{fd} - Observe within the subspace of the d leading *FLV* \Rightarrow Satisfy a weaker necessary condition [Frank and Zhuk 2018]
- ▶ \mathbf{H}_k^{bd} - Observe within the subspace of the d leading *BLV*. \Rightarrow Satisfy a stronger sufficient condition [Bocquet *et al* 2017]
- ▶ $\mathbf{I}_{d \times n}$ - Observe the first d components.
- ▶ Projected observations based on dynamics was studied earlier by [Law *et al* 2016].

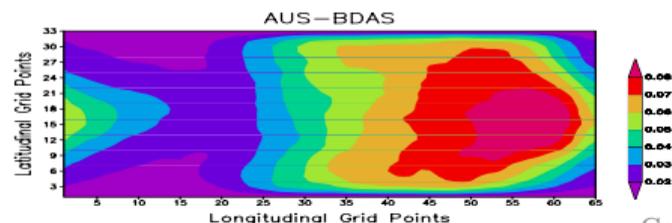
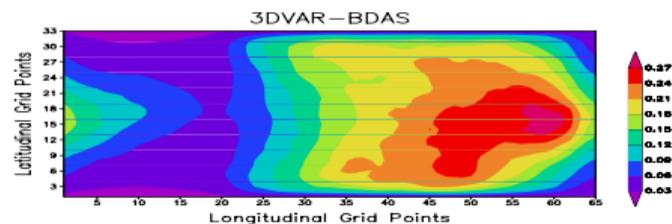
AUS and Target Observations

TARGET OBSERVATION STRATEGY: **Breeding on the Data Assimilation System** BDAS

- Quasi-geostrophic atmospheric model (Rotunno and Bao, 1996 MWR)
- Perfect model setup - Observation Dense area (1-20 Longitude) - Target Area, one obs between 21-64 Longitude

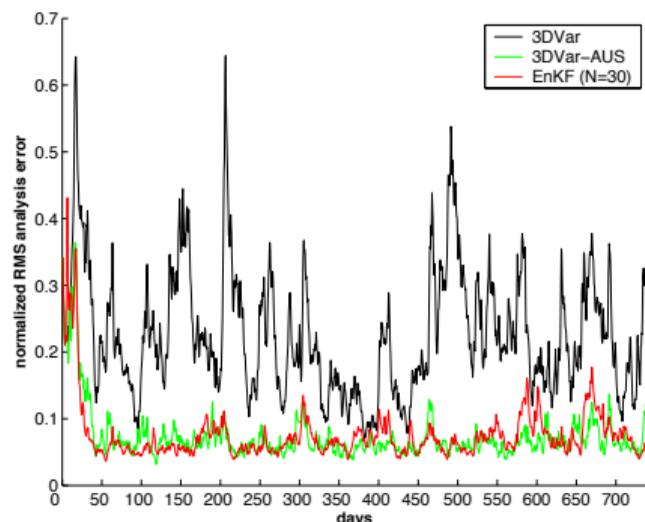


Experiment	Ocean Obs Type/Positioning/Assimilation	RMSE
LO	-	0.462
FO	vert.Prof/fixed(in the max(err))/3DVar	0.338
RO	vert.Prof/random/3DVar	0.311
3DVar-BDAS	vert.Prof/BDAS/3DVar	0.184
AUS-BDAS	temp.1-Level/BDAS/AUS	0.060



Carrassi et al 2007

Hybrid 3DVar-AUS: *Enhancing the performance of a 3DVar using AUS*

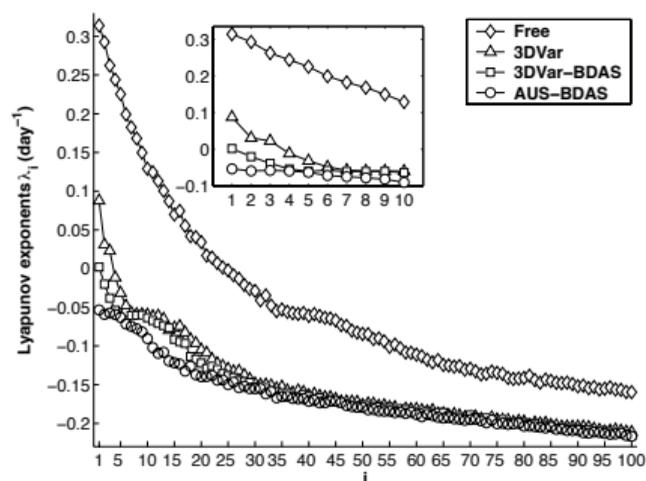


Carrassi *et al* 2008a

- ▶ QG model on a β -plane; network of randomly distributed obs (vertical soundings).
- ▶ 3DVar-AUS: (1) AUS assimilates the obs located on the unstable mode; (2) 3DVar process the remaining obs.
- ▶ 3DVar-AUS comparable to EnKF with only one unstable mode \Rightarrow Reduced computational cost and implementation on a pre-existing 3DVar scheme.

Does stabilization improve estimation?

Quasi-Geostrophic Model (Rotunno and Bao, 1996)



Experiment	RMSE
Free	1
3DVar	0.321
3DVar-BDAS	0.163
AUS-BDAS	0.058

- ▶ DA “always” provides a stabilizing effect (*e.g.* compare 3DVar with free system Lyapunov spectrum) but ...
- ▶ if the DA is designed to kill the instabilities, the estimation error is efficiently reduced

Carrassi *et al* 2008b

Conclusions

- ▶ We have shown that the (En)KF/(En)KS in deterministic dynamics naturally project the uncertainty on the unstable-neutral subspace $\Rightarrow N = n_0$ members are sufficient.
- ▶ These properties are at the basis of the **assimilation in the unstable subspace** (**AUS**, by *A. Trevisan & Collaborators*), where the unstable subspace is explicitly used in the DA process.
- ▶ AUS has been successfully applied to deterministic atmospheric, oceanic and traffic models [see *Palatella et al 2013* for a review].
- ▶ In stochastic dynamics we have shown that weakly stable modes of high variance must be included.
- ▶ Furthermore we have demonstrated the existence of an *upwelling* of uncertainty from unfiltered-to-filtered subspace that motivates the need for multiplicative inflation.
- ▶ All has been done within a Gaussian framework \Rightarrow Can the unstable subspace be used to develop efficient fully Bayesian (*Particle Filters*) methods? Maybe... [see *Maclean & Van Vleck 2019* - Next talk.]

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