

REVISITING CLASSICAL PROBLEMS OF IMAGE PROCESSING:

Looking for new ways to address longstanding problems

Mauricio Delbracio

Duke University

SIAM Conference on Imaging Science
24 May 2016

Thanks to...

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UdelaR



Pablo Musé
UdelaR



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NC State University



Cecilia Aguerrebere
Duke University

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Looking for simpler ways to address longstanding problems:

- 1 Recovering the Camera Point Spread Function
- 2 Removing Camera Shake via Fourier Burst Accumulation
- 3 Boosting Stochastic Renderers by Auto-similarity Filtering

Act one

Recovering the Camera Point Spread Function

Joint work with: A. Almansa¹, P. Musé² and J.-M. Morel³

¹Telecom Paristech, ²UdelaR, ³ENS-Cachan

Image blur can be a consequence of

- Camera misusing or scene configuration (Extrinsic):
 - Wrongly setting the camera focus
 - Only an specific interval of depths in focus
 - Camera shake, scene motion

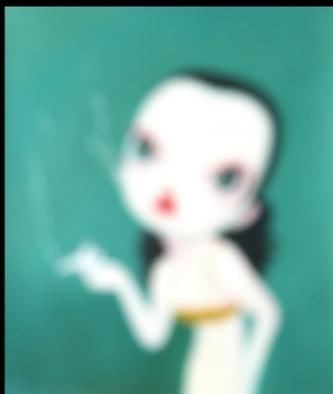
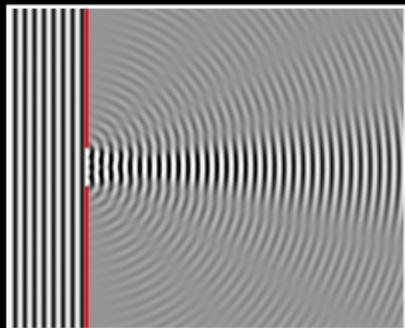


Image blur can be a consequence of

- Physical camera phenomena (Intrinsic):
 - Light diffraction
 - Sensor averaging
 - Lens aberration
 - Optical anti-aliasing filter



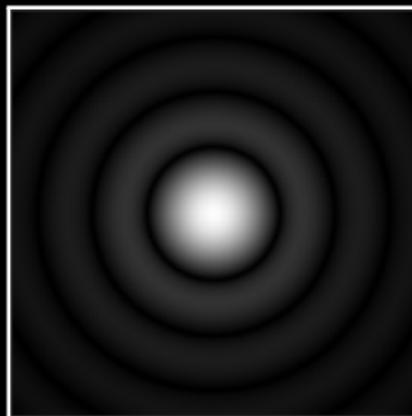
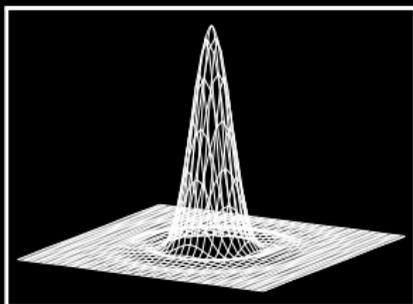
Our Goal

Accurately estimate a function, called **Point Spread Function (PSF)**, that models the blur due to intrinsic camera phenomena.

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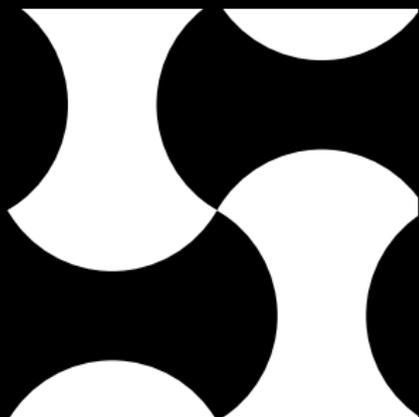
Accurately estimate a function, called **Point Spread Function (PSF)**, that models the blur due to intrinsic camera phenomena.

Image ideally obtained from a null-area point light source (impulse response).



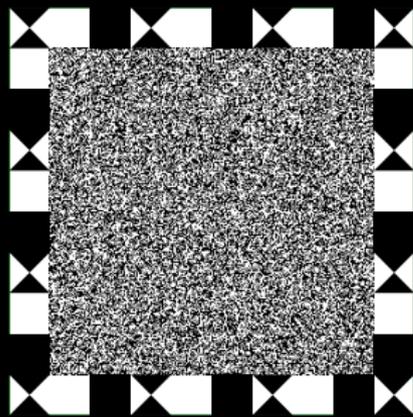
Non-blind estimation: use a calibration pattern

- How do we choose the calibration pattern?
- Local, accurate, subpixel PSF estimation - is it possible?



"knife-edge"

Joshi, et al. '08



"Bernoulli pattern"

Delbracio, et al. '12

Problem statement

Image Formation Model



$$\mathbf{v} = \mathbf{S}_1 (D(u) * h) + \mathbf{n}$$

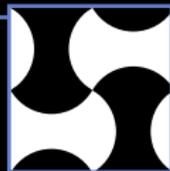
- \mathbf{v} is the acquired image.
- u is the continuous image.
- h local convolution kernel due to all PSF like effects.
- $D(\cdot)$ geometrical transformation.
- \mathbf{S}_1 bi-dimensional ideal ideal sampling operator (sensor array).
- \mathbf{n} models all noise sources of the acquisition process.

Problem statement

Image Formation Model



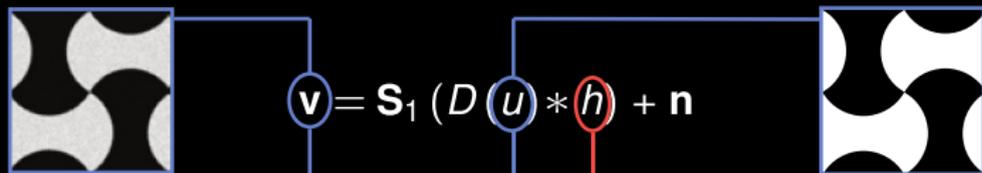
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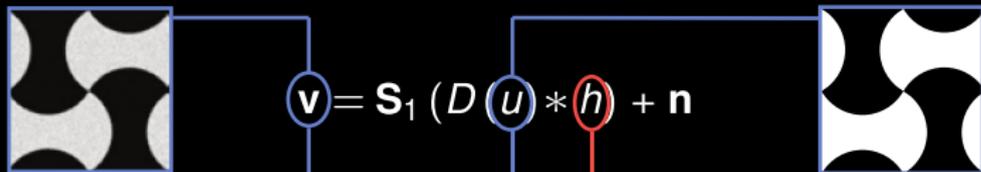
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Problem statement

Image Formation Model



- \mathbf{v} is the acquired image.
 - \mathbf{u} is the continuous image.
 - \mathbf{h} local convolution kernel due to all PSF like effects.
 - $D(\cdot)$ geometrical transformation.
 - \mathbf{S}_1 bi-dimensional ideal ideal sampling operator (sensor array).
 - \mathbf{n} models all noise sources of the acquisition process.
- Model is local: \mathbf{h} may change all over the image.

Problem statement (cont.)

Discrete Image Formation Model

- h is band-limited in $\text{supp}(\hat{h}) = [-s\pi, s\pi]^2$, e.g., $s = 3 - 4$
- Take samples at rate at least $s \times$ to correctly sample the PSF

Problem statement (cont.)

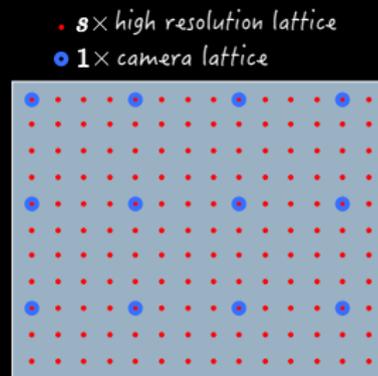
Discrete Image Formation Model

- h is band-limited in $\text{supp}(\hat{h}) = [-s\pi, s\pi]^2$, e.g., $s = 3 - 4$
- Take samples at rate at least $s\times$ to correctly sample the PSF

Continuous model can be replaced by a $s\times$ oversampled discrete model,

$$v = \mathcal{S}_s(u * h) + n$$

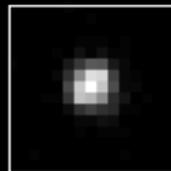
- u : samples at $s\times$ of the low-pass filtered distorted pattern image.
- \mathcal{S}_s : s -subsampling operator ($\mathcal{S}_s u)(x) = u(sx)$



Mathematical Formulation

Solve an inverse problem based on prior information about the small spatial support of the PSF.

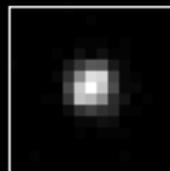
$$N = r \times r$$



Mathematical Formulation

Solve an inverse problem based on prior information about the small spatial support of the PSF.

$$N = r \times r$$



$$\tilde{h} = \arg \min_{h \in \mathbb{R}^N} \left\| S_s C[u] h - v \right\|^2$$

- v : observation
- u : $s \times$ resolution rasterized distorted pattern
- $C[u]$: convolution with u in matrix form
- S_s : s -sub-sampling operator in matrix form



Mathematical Formulation (cont.)

Solution to the inverse problem

$$\tilde{\mathbf{h}} = (\mathbf{S}_s \mathbf{C}[\mathbf{u}])^\dagger \mathbf{v}.$$

Mathematical Formulation (cont.)

Solution to the inverse problem

$$\tilde{\mathbf{h}} = (\mathbf{S}_s \mathbf{C}[\mathbf{u}])^\dagger \mathbf{v}.$$

The mean square error is given by

$$\text{MSE}(\tilde{\mathbf{h}}) = \underbrace{\|(\mathbf{S}_s \mathbf{C}[\mathbf{u}])^\dagger\|_{\text{F}}^2}_{\gamma} \underbrace{\sigma^2}_{\text{noise level}}$$

Mathematical Formulation (cont.)

Solution to the inverse problem

$$\tilde{\mathbf{h}} = (\mathbf{S}_s \mathbf{C}[\mathbf{u}])^\dagger \mathbf{v}.$$

The mean square error is given by

$$\text{MSE}(\tilde{\mathbf{h}}) = \underbrace{\|(\mathbf{S}_s \mathbf{C}[\mathbf{u}])^\dagger\|_{\text{F}}^2}_{\gamma} \underbrace{\sigma^2}_{\text{noise level}}$$

To minimize the error, one has to minimize

$$\gamma(\mathbf{S}_s \mathbf{C}[\mathbf{u}]) := \|(\mathbf{S}_s \mathbf{C}[\mathbf{u}])^\dagger\|_{\text{F}}^2 = \sum_{i=1}^N \sigma_i^{-2},$$

where $\{\sigma_1, \sigma_2, \dots, \sigma_N\}$ are the singular values of $\mathbf{S}_s \mathbf{C}[\mathbf{u}]$.

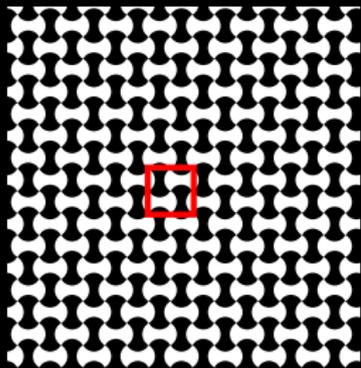
- γ controls the noise amplification,
- should be as low as possible.

Proposition (Lower bound for optimal patterns)

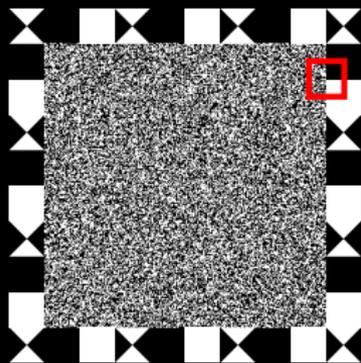
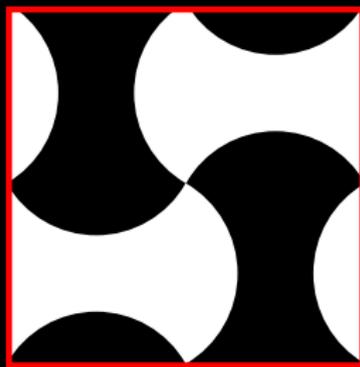
$$\min_{a \leq u_{ij} \leq b} \gamma(S_s C[u]) \geq \frac{1}{MN} \left(\frac{1}{b^2} + \frac{4(N-1)^2}{(b-a)^2} \right),$$

- $M = m \times n$ is the observation window size
- $N = r \times r$ is the kernel size
- Constraints $a \leq u_{ij} \leq b$ are linked to the physical realization and dynamic range of the sensors.

Comparing calibration patterns



Slant-edge pattern - Joshi et al. [2008]



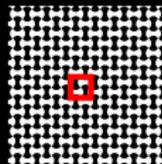
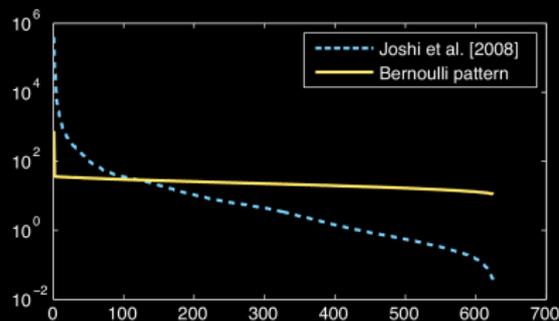
Bernoulli pattern - Delbracio et al. [2012]



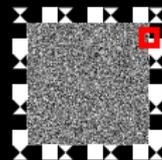
Optimality of the Bernoulli pattern

Why this i.i.d Bernoulli(0.5) random noise pattern?

Singular values of $S_s C[u]$



Slant-edge pattern - Joshi et al. [2008]

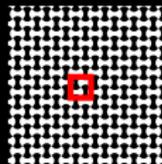
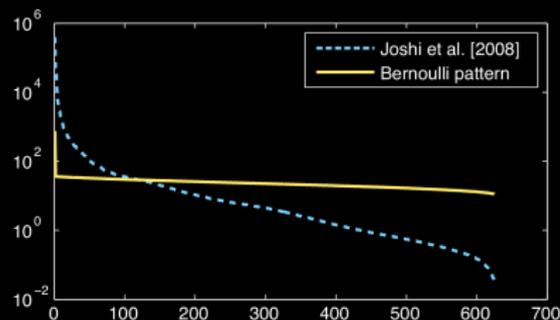


Bernoulli pattern - Delbracio et al. [2012]

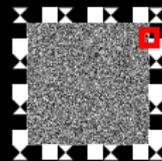
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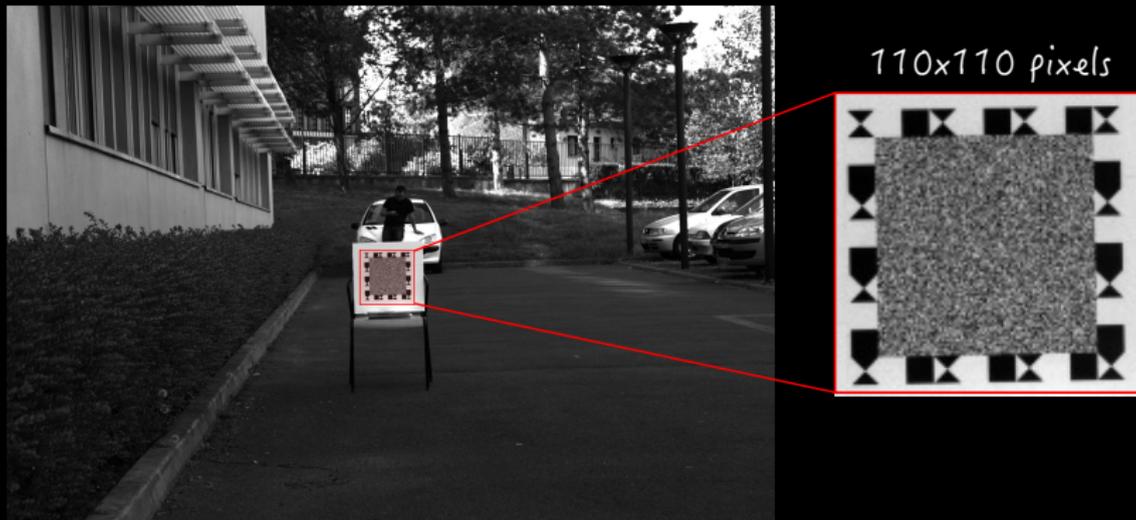
γ value:

	9×9	17×17	25×25	33×33
Theoretical bound	0.10	0.35	0.70	1.15
Bernoulli pattern	0.19	0.69	1.54	2.98
Joshi et al. [2008]	99.44	1133.05	6445.87	58419.08

Experiments

Real camera examples

Canon EOS 400D - Tamron AF 17-50mm F/2.8 XR Di-II lens, 50mm, Green channel 1.



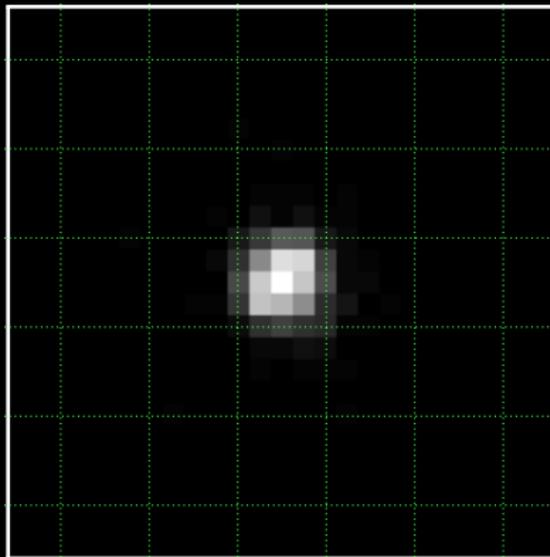
IPOL: Image Processing Online – ipol.im

Detailed Description + Online Demo + Source Code

Experiments

Real camera examples

Estimation at $4\times$ the sensor resolution for the Green channel 1.

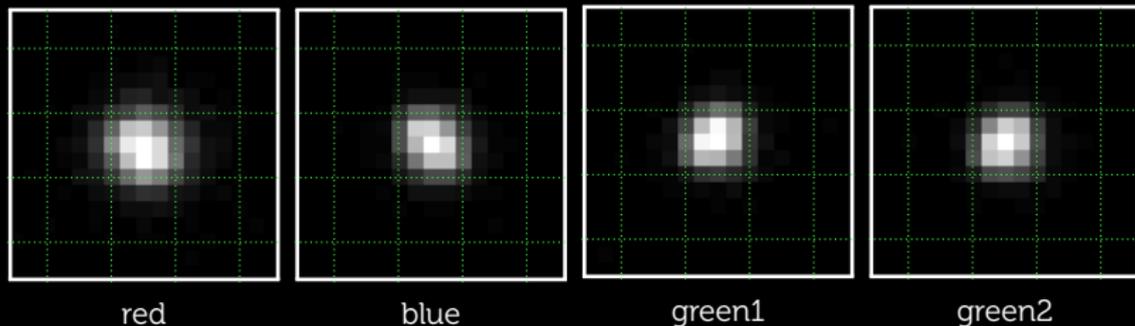


Canon EOS 400D - Tamron AF 17-50mm F/2.8 XR Di-II lens, f/5.6, Green channel 1.

Experiments

Different color channels

4× PSF estimation for the four Bayer pattern channels (RAW output).

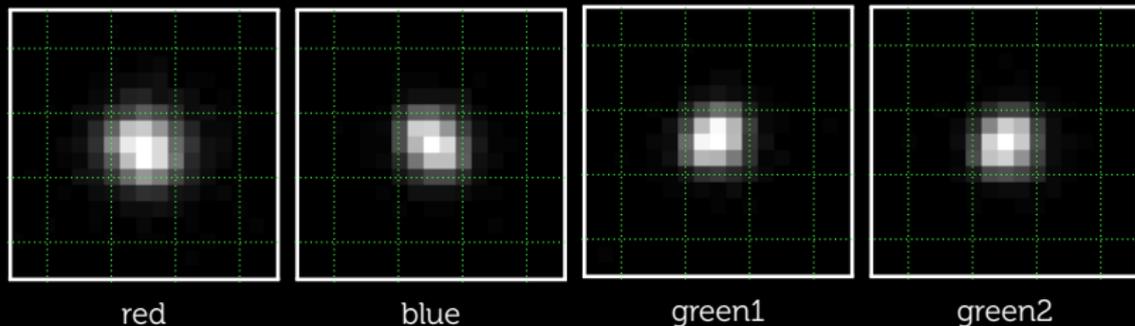


- Red PSF larger than green and blue ones (diffraction)

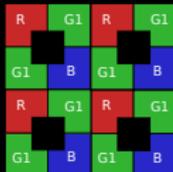
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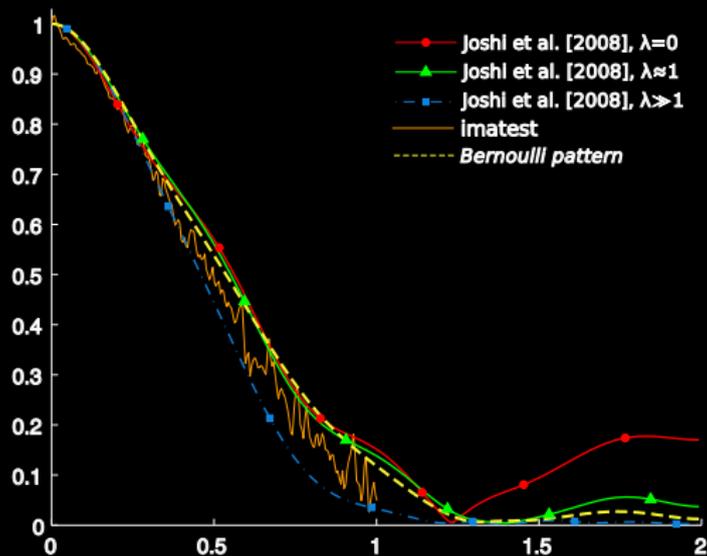
- Red PSF larger than green and blue ones (diffraction)
- Blue and green are symmetric (sensor active area L-shaped)



Comparison of several methods

- Joshi et al. [2008] very sensitive to regularization $+\lambda\|\nabla h\|^2$

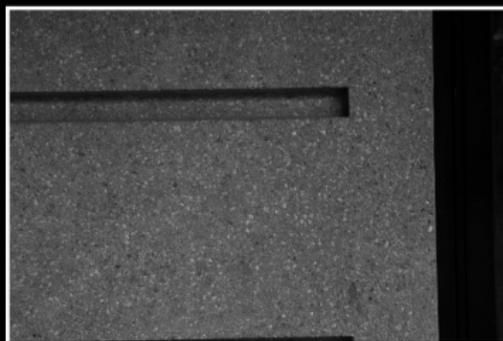
MTF horizontal profile



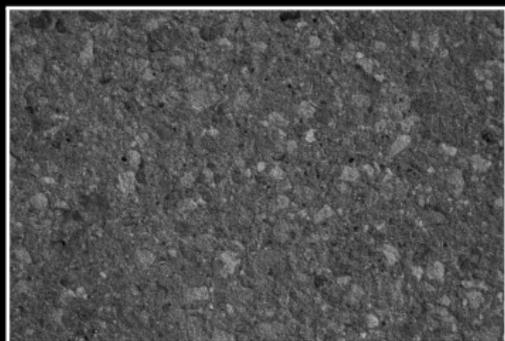
Can we avoid the calibration pattern?

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Yes! Take two parallel photos (same scene) different distances



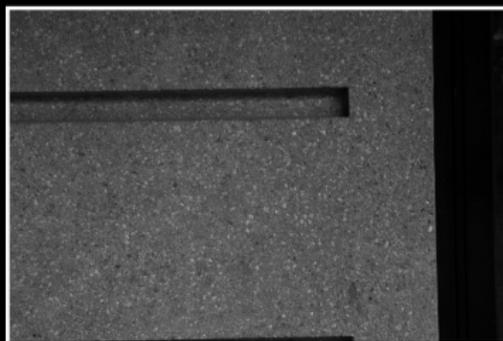
Farthest photograph v_2



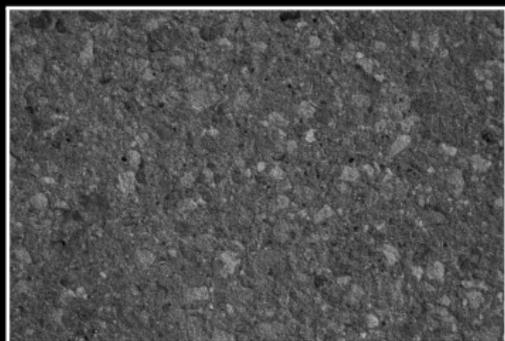
Closest photograph v_1

Can we avoid the calibration pattern?

Yes! Take two parallel photos (same scene) different distances



Farthest photograph v_2



Closest photograph v_1

- If acquired sufficiently far from each other: the PSF can be estimated from the relative blur between the two images

Relative blur between two images

- v_1, v_2 two fronto-parallel views (same scene), zooms $\lambda_1 < \lambda_2$.

Definition (inter-image kernel)

Any k satisfying

$$v_2 = H_\lambda v_1 * k, \quad \lambda := \lambda_2 / \lambda_1.$$

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Proposition

Under (mild assumptions), there is a unique inter-image kernel k ,

$$H_\lambda h * k = h,$$

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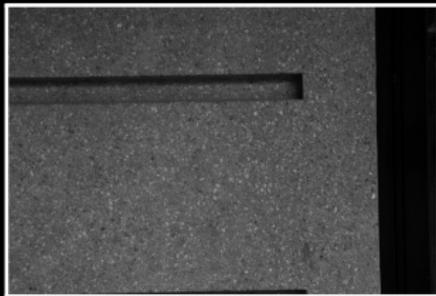
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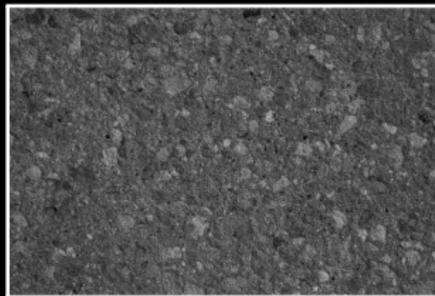
and h can be obtained from k as:

$$h = \lim_{n \rightarrow \infty} H_{\lambda^{n-1}} k * \dots * H_\lambda k * k.$$

Results: A running example

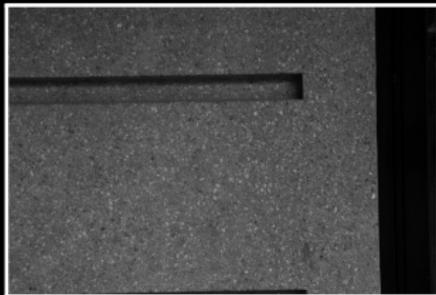


farthest image

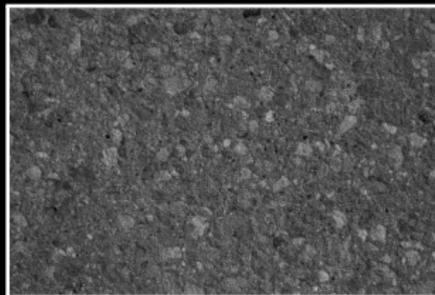


closest image

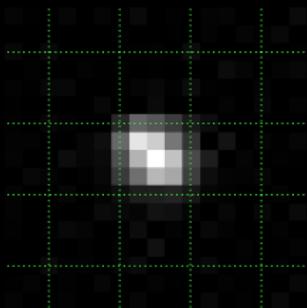
Results: A running example



farthest image

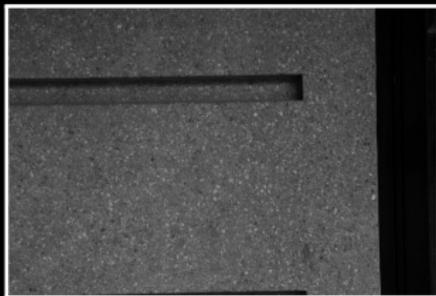


closest image

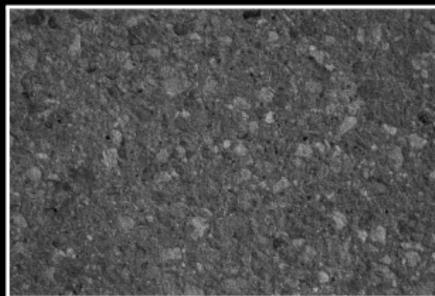


4 × Inter-image kernel

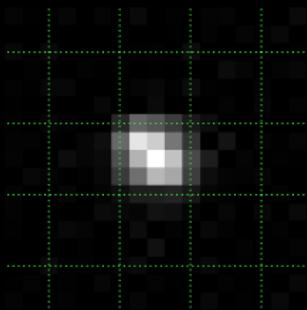
Results: A running example



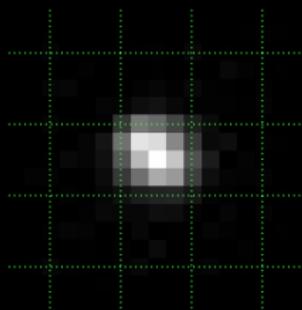
farthest image



closest image

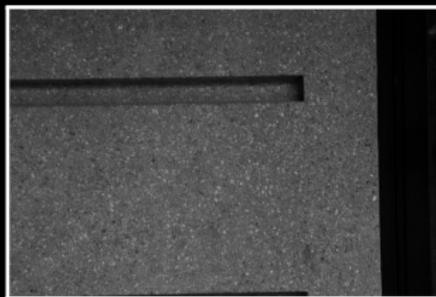


4 × Inter-image kernel

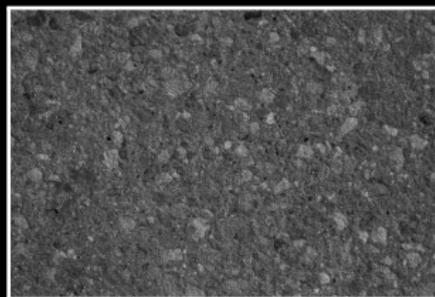


4 × PSF

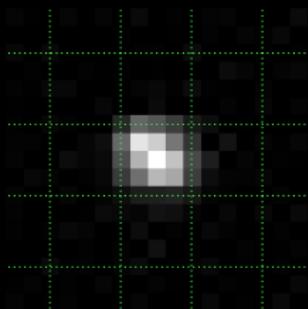
Results: A running example



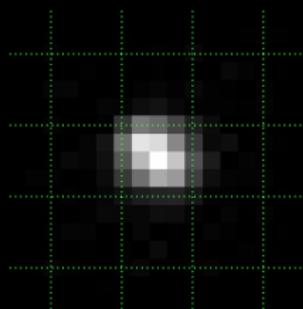
farthest image



closest image



4 × Inter-image kernel

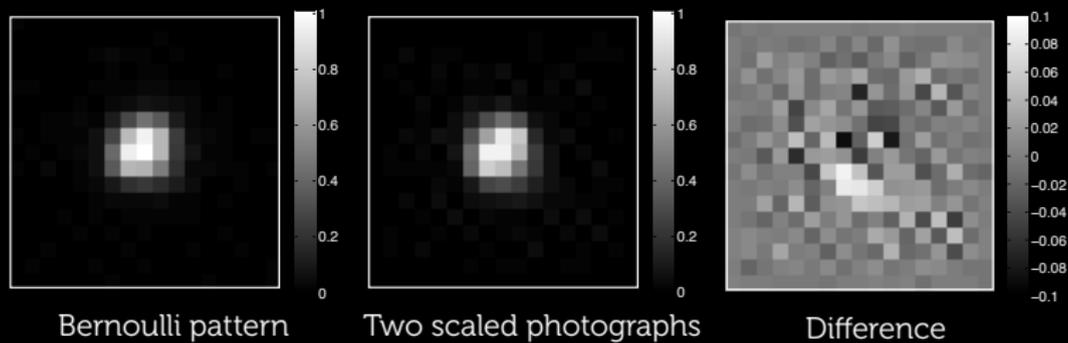


4 × PSF

IPOL: Image Processing Online – ipol.im

Detailed Description + Online Demo + Source Code

Bernoulli pattern vs. Two-scaled photographs



- Estimations at $4\times$ the camera resolution

- Avoid regularization: Chose the right calibration pattern
- Subpixel, accurate PSF estimation is well-posed if calibration pattern carefully chosen
- Bernoulli random pattern near optimal
- Calibration pattern can be “avoided” by taking two images of the same scene

Act two

Removing Camera Shake Blur
via Fourier Burst Accumulation

Joint work with: G. Sapiro
Duke University

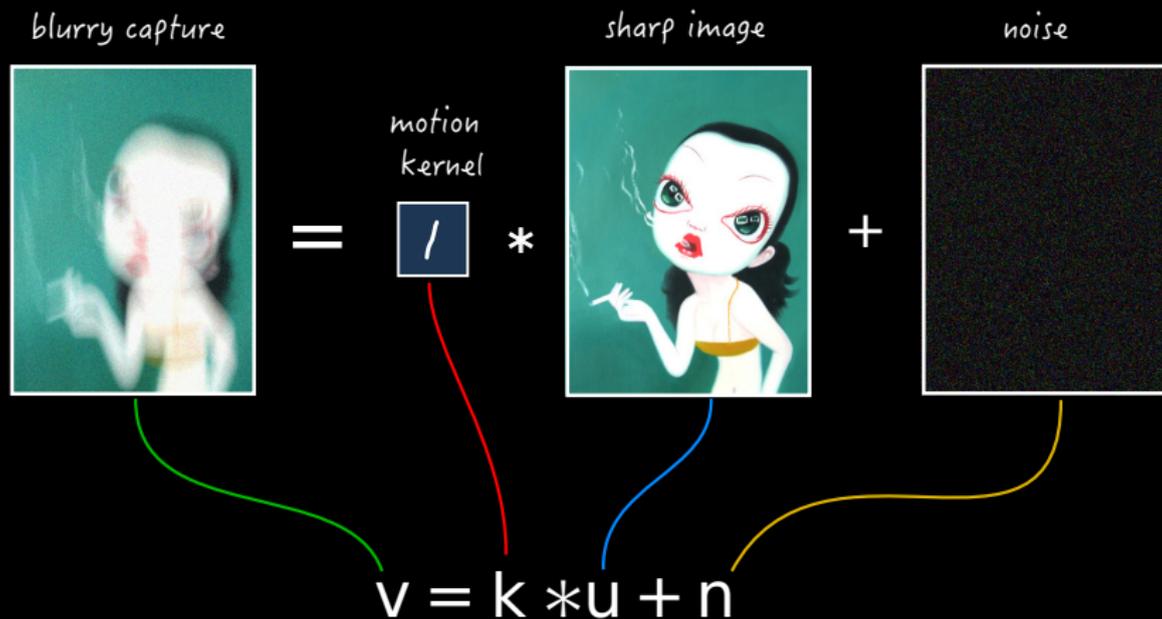
Deblurring



Deblurring



Shift-invariant blurring model

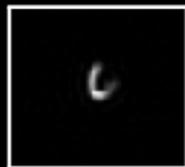


Standard Deblurring: Blind Deconvolution

- 1 Get a blurry image v



- 2 Estimate the motion kernel k



- 3 Perform nonblind deconvolution to recover \tilde{u}



Standard Deblurring: Blind Deconvolution

- 1 Get a blurry image v
- 2 Estimate the motion kernel k
- 3 Perform nonblind deconvolution to recover \tilde{u}



But...

- Who actually cares about the motion kernel?
- Even if the kernel is perfectly known the inversion is ill-posed

Standard Deblurring: Blind Deconvolution

- 1 Get a blurry image v
- 2 Estimate the motion kernel k
- 3 Perform nonblind deconvolution to recover \tilde{u}



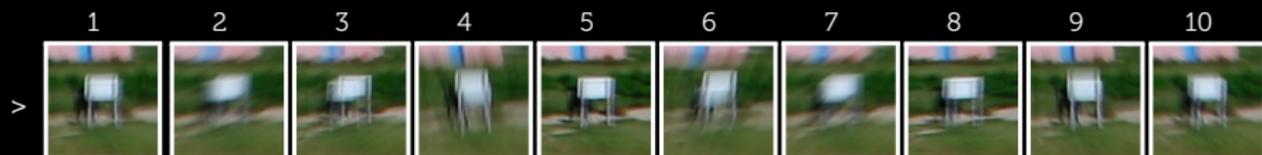
But...

- Who actually cares about the motion kernel?
- Even if the kernel is perfectly known the inversion is ill-posed

Is it possible to avoid explicit inversion?

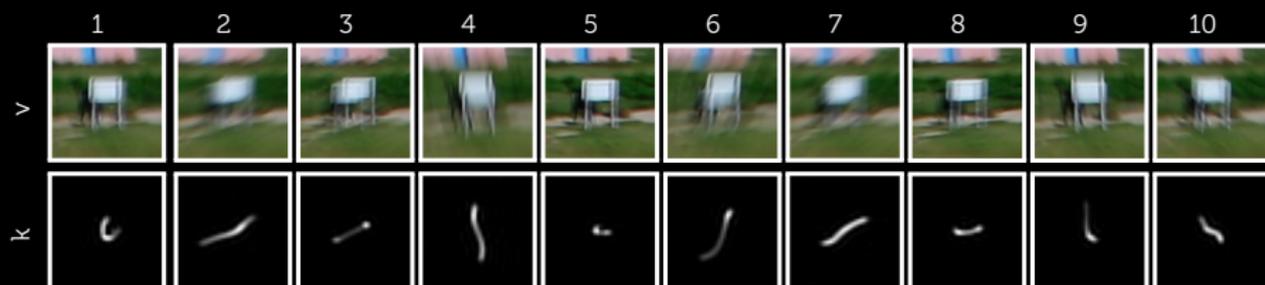
Burst photography

- In 2016... we can take a burst of 6-12 images



Burst photography

- In 2016... we can take a burst of 6-12 images



- Hand shake/tremor is random:
- Different images \rightarrow Different blur (in general)

A basic principle

Claim (Blurring kernels do not amplify the spectrum)

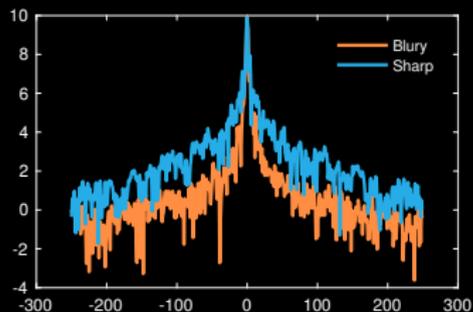
sharp



blury



Fourier spectrum (vertical)



A basic principle

Claim (Blurring kernels do not amplify the spectrum)

Let $k(x) \geq 0$ and $\int k(x) = 1$. Then, $|\hat{k}(\zeta)| \leq 1, \forall \zeta$.

Proof.

$$|\hat{k}(\zeta)| = \left| \int k(x) e^{ix \cdot \zeta} dx \right| \leq \int |k(x)| dx = \int k(x) dx = 1.$$

□

Deblurring by Fourier Burst Accumulation

The basic Fourier Burst Accumulation algorithm:

- 1 Take a burst

Deblurring by Fourier Burst Accumulation

The basic Fourier Burst Accumulation algorithm:

- 1 Take a burst
- 2 Align the images (with respect to the center one)

Deblurring by Fourier Burst Accumulation

The basic Fourier Burst Accumulation algorithm:

- 1 Take a burst
- 2 Align the images (with respect to the center one)
- 3 Combine the images in the Fourier domain (FBA)

Fourier Burst Accumulation (FBA)

$$\bar{u}(\mathbf{x}) = \mathcal{F}^{-1} \left(\sum_{i=1}^M w_i(\zeta) \cdot \hat{v}_i(\zeta) \right) (\mathbf{x}), \quad w_i(\zeta) = \frac{|\hat{v}_i(\zeta)|^p}{\sum_{j=1}^M |\hat{v}_j(\zeta)|^p},$$

Deblurring by Fourier Burst Accumulation

The basic Fourier Burst Accumulation algorithm:

- 1 Take a burst
- 2 Align the images (with respect to the center one)
- 3 Combine the images in the Fourier domain (FBA)
- 4 Inverse Fourier

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The basic Fourier Burst Accumulation algorithm:

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- 2 Align the images (with respect to the center one)
- 3 Combine the images in the Fourier domain (FBA)
- 4 Inverse Fourier
- 5 Unsharp masking (Optional)

Fourier Burst Accumulation (FBA)

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- (v_i) : M input aligned images, $\hat{v}_i = \mathcal{F}(v_i)$ – Fourier Transform
- The larger $|\hat{v}_i(\zeta)|$, the more $\hat{v}_i(\zeta)$ contributes to \bar{u}
- p controls the aggregation procedure (soft-max):
 - If $p = 0$: arithmetic average
 - If $p = \infty$: max
 - $0 < p < \infty$: Balance Max/Mean

Fourier Burst Accumulation: An example

Input Burst



Fourier Burst Accumulation: An example

Input Burst



Fourier Burst Accumulation

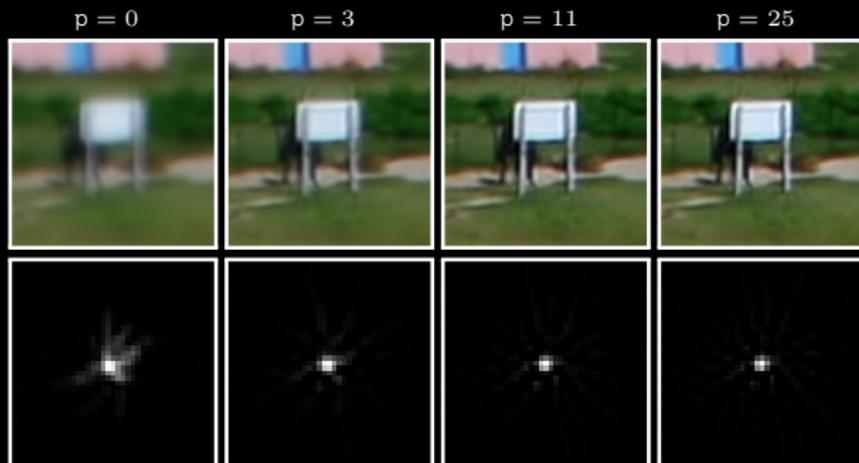


Fourier Burst Accumulation: An example

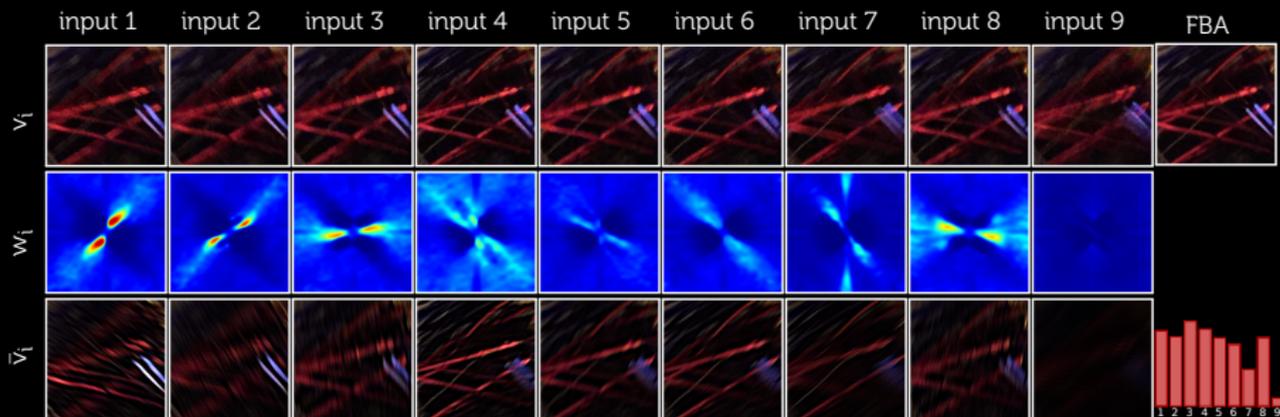
Input Burst



Fourier Burst Accumulation



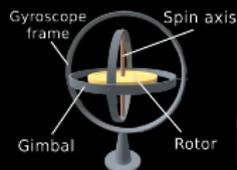
Anatomy of the Fourier Aggregation



$$\bar{u}(x) = \sum_{i=1}^M \underbrace{\mathcal{F}^{-1} \left(w_i(\zeta) \cdot \hat{v}_i(\zeta) \right)}_{\hat{v}_i(x)}(x), \quad w_i(\zeta) = \frac{|\hat{v}_i(\zeta)|^p}{\sum_{j=1}^M |\hat{v}_j(\zeta)|^p}, \quad p = 11.$$

Align images

- Can use gyroscope and accelerometer (Work in progress...)



- Current efficient approach
 - SIFT + Ransac (Homography)
 - Optical Flow (in low resolution)



More Results











Woods 6/12



Woods 7/12













Woods Align & Average ($p = 0$)







Cabo Polonio 2/14





Cabo Polonio 4/14



Cabo Polonio 5/14



Cabo Polonio 6/14



Cabo Polonio 7/14



Cabo Polonio 8/14



Cabo Polonio 9/14



Cabo Polonio 10/14



Cabo Polonio 11/14



Cabo Polonio 12/14







Cabo Polonio Align & Average ($p = 0$)





Cabo Polonio 5/14 (Best Frame)



More Results



Typical Shot

Best Shot

Align and average

Šroubek &
Milanfar [2012]

Zhang et al.
[2013]

FBA

Extension to videos



- Fourier weighted average to remove camera shake blur
- No (explicit) inversion, no kernel estimation, no deconvolution
- Not universal (blurs in the burst need to be different)

Act three

Boosting Stochastic Renderers by
Auto-similarity Filtering

Joint work with: P. Musé¹, T. Buades², J. Chauvier³, N. Phelps³, J.-M. Morel⁴

¹UdelaR, ²UiB, ³Eon-Software, ⁴ENS-Cachan

Realistic Image Synthesis

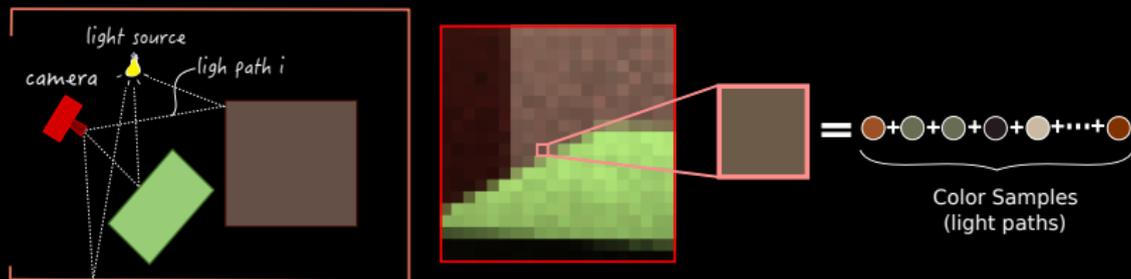
Goal: Generate images from a 3D virtual scene



Realistic Image Synthesis

Monte Carlo Rendering

- Ray-tracing: popular technique for resolving the equilibrium of light in a scene (rendering equation [Kajiya 1986]).
- Pixel color = average of values along light paths
 - cast from image pixel, through camera aperture, bouncing in the scene and reaching a light source.



Realistic Image Synthesis

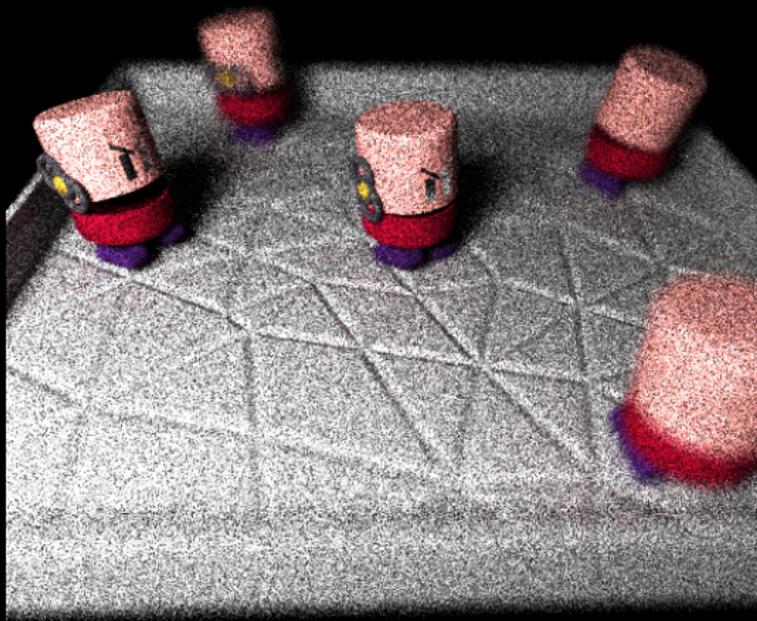
Monte Carlo Rendering

Unfortunately...

- Only a finite number of rays can be cast
- To avoid artifacts, rays are cast randomly
- Equivalent to solving the light equilibrium through a Monte Carlo integration procedure
- Variance converges linearly with number of samples

Realistic Image Synthesis

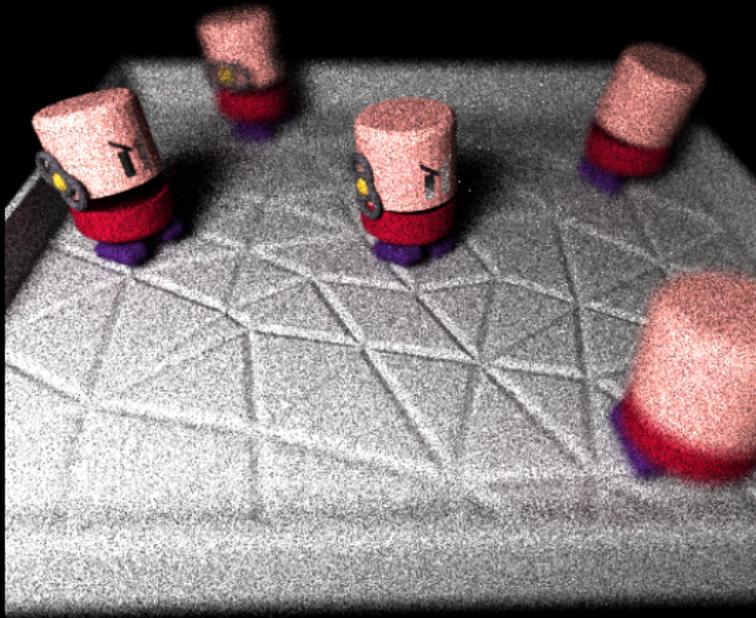
Monte Carlo Rendering Noise



32 samples per pixel (spp) [8s]

Realistic Image Synthesis

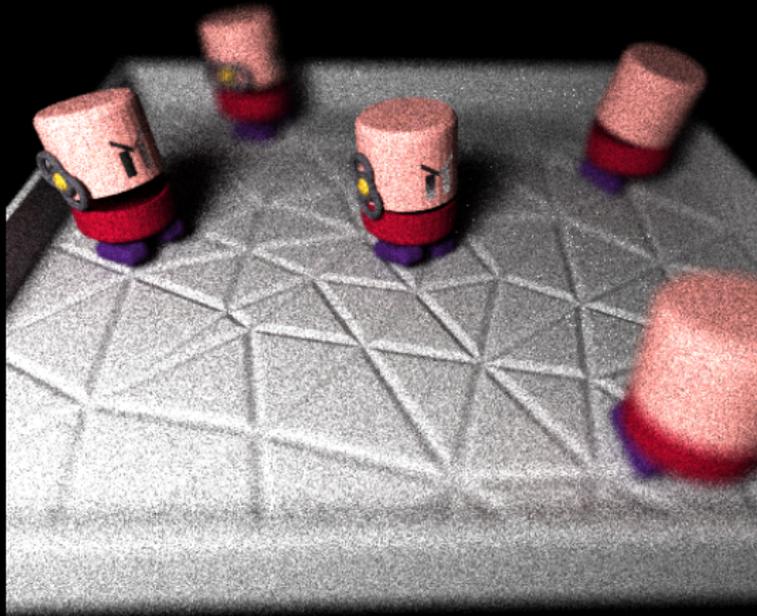
Monte Carlo Rendering Noise



64 samples per pixel (spp) [16s]

Realistic Image Synthesis

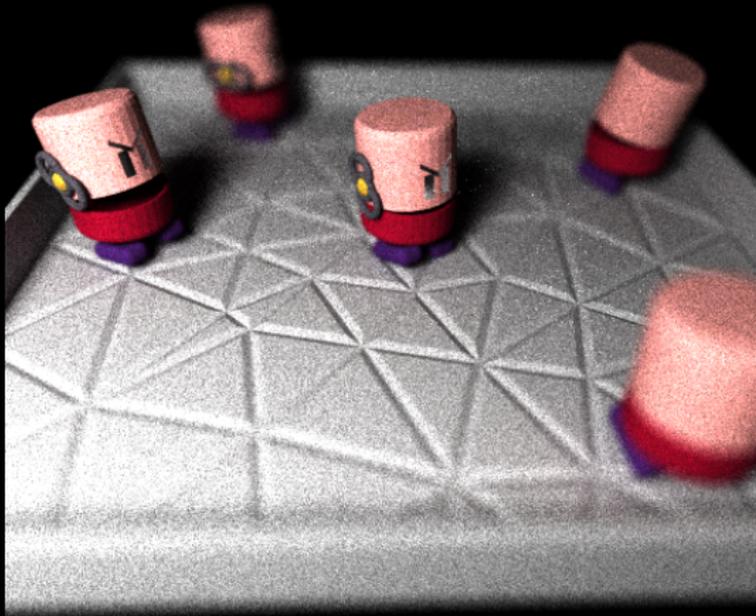
Monte Carlo Rendering Noise



128 samples per pixel (spp) [32s]

Realistic Image Synthesis

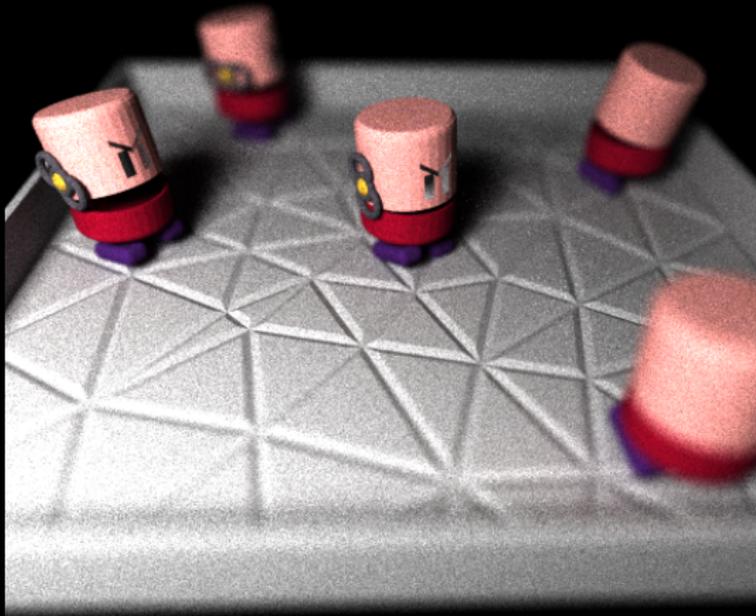
Monte Carlo Rendering Noise



256 samples per pixel (spp) [64s]

Realistic Image Synthesis

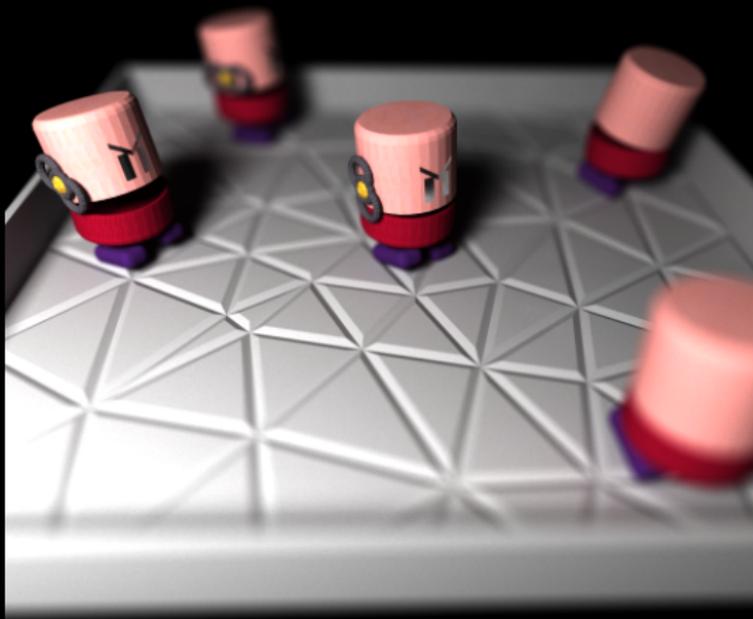
Monte Carlo Rendering Noise



512 samples per pixel (spp) [128s]

Realistic Image Synthesis

Monte Carlo Rendering Noise



65536 samples per pixel (spp) [16384s]

A General Principle: Auto-similarity

“Similar pixels must be denoised jointly,
being different samples of the same model.”

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Pixel similarity based on:

- Gaussian filter: spatial proximity
- Sigma/Bilateral filter: pixel color [Lee, et al., '83], [Tomasi et al., '98]
- NLmeans/BM3D: patch color [Buades et al. '05], Dabov et al, '07]
- LARK, GLIDE: kernels [Takeda, et al. '07], [Talebi, et al. '14]

A General Principle: Auto-similarity

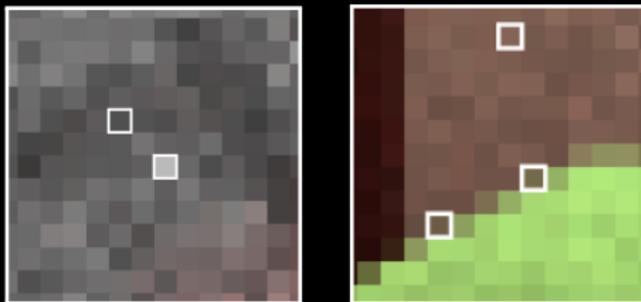
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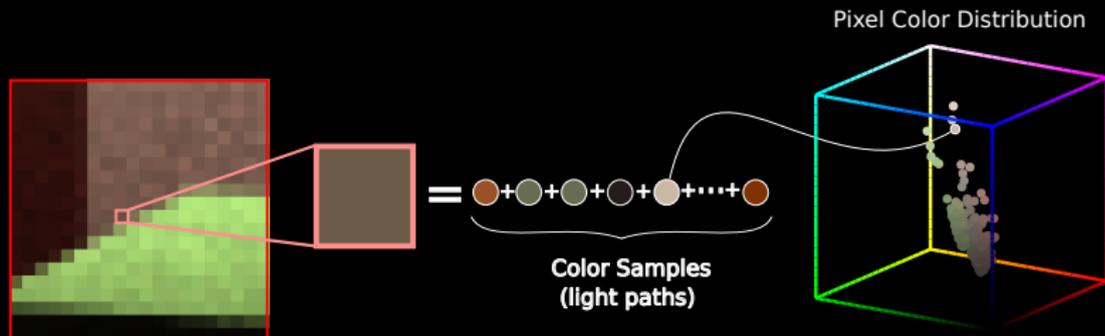
Main difficulty:

Hard to distinguish noise from intrinsic pixel variability (bias)



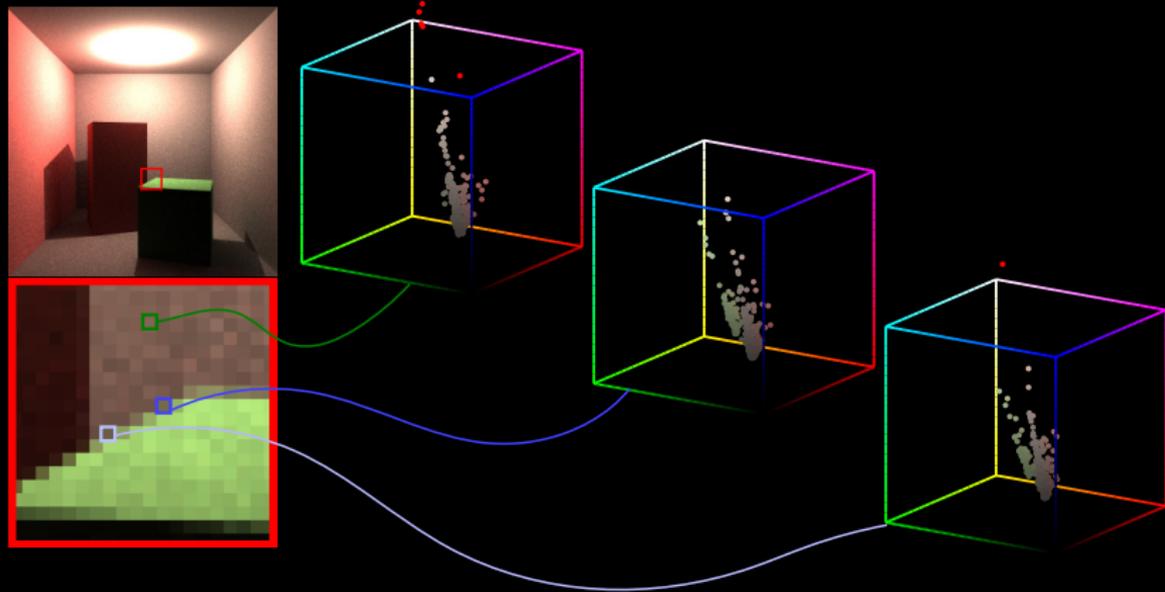
Empirical Ray Color Distribution

- During rendering a lot of information is computed
- In particular: RGB color of each ray hitting a given pixel
- Color histograms of rays cast from each pixel
- Use this histogram to define pixel similarity



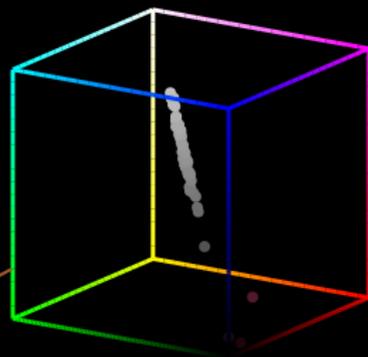
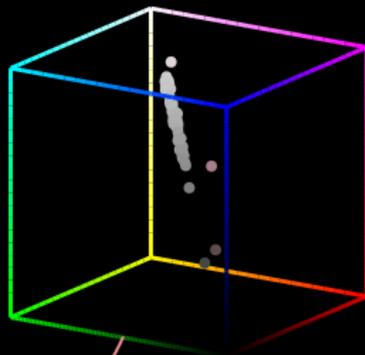
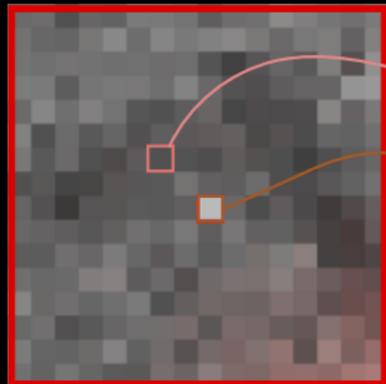
Empirical Ray Color Distribution

Measuring pixel similarity



Empirical Ray Color Distribution

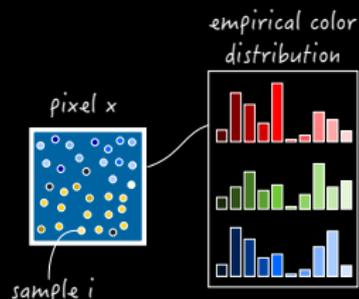
Measuring pixel similarity



Empirical Ray Color Distribution

Measuring pixel similarity

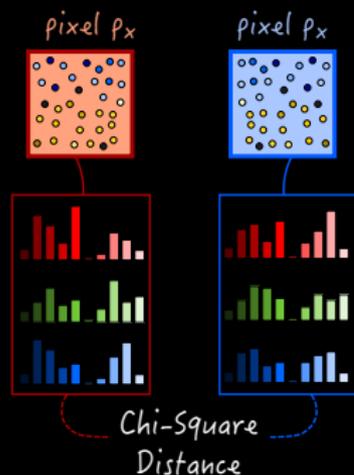
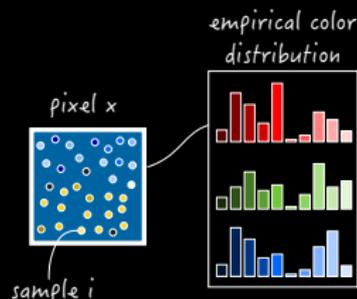
- Pixel similarity measured from the binned empirical color distributions
- Tri-stimulus color images: $3 \times 1D$ histograms (one per color channel)



Empirical Ray Color Distribution

Measuring pixel similarity

- Pixel similarity measured from the binned empirical color distributions
- Tri-stimulus color images: $3 \times 1D$ histograms (one per color channel)



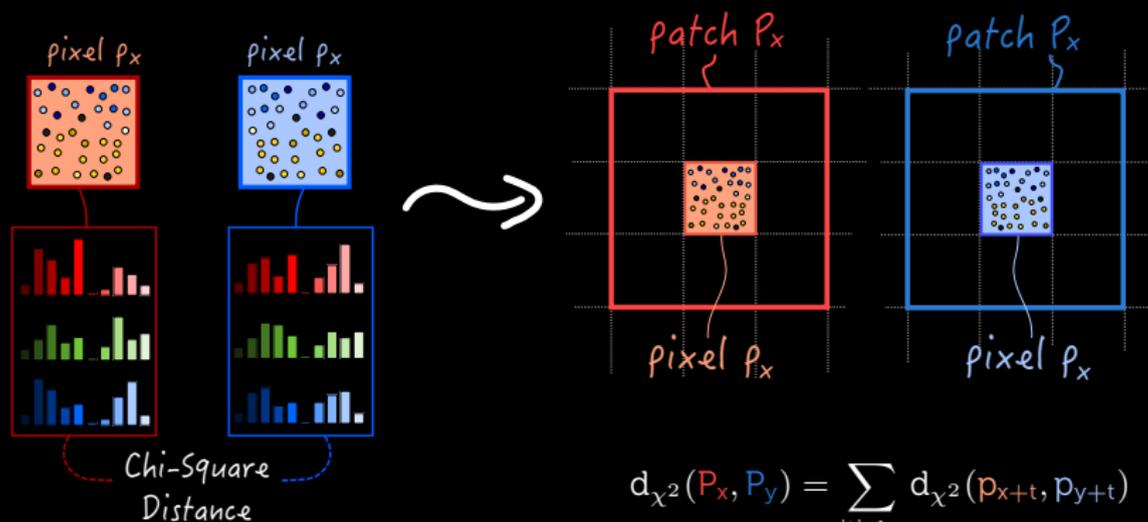
Given n_b -binned empirical color distributions ($h_i(x)$) and ($h_i(y)$) of pixels x and y , the Chi-Square distance is given by

$$d_{\chi^2}(p_x, p_y) = \sum_{k=1}^{n_b} \frac{(h_k(x) - h_k(y))^2}{h_k(x) + h_k(y)}$$

Empirical Ray Color Distribution

Measuring pixel similarity

- Extended to patches for spatial coherence



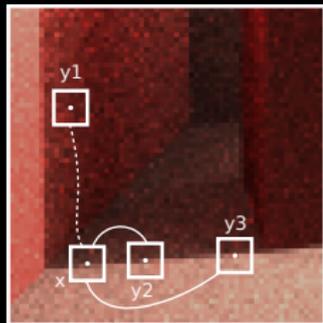
$$d_{\chi^2}(P_x, P_y) = \sum_{|t| \leq w} d_{\chi^2}(p_{x+t}, p_{y+t})$$

Distribution-driven average

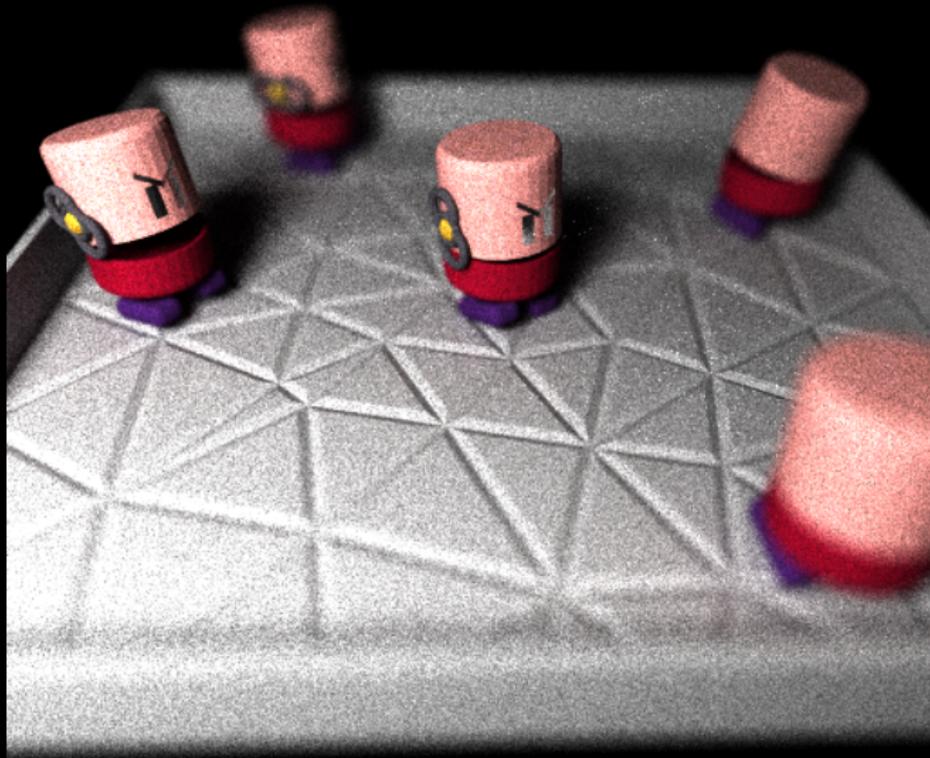
- Given a noisy patch, all the patches that are at a distance less than κ are considered to be similar.

Distribution-driven average

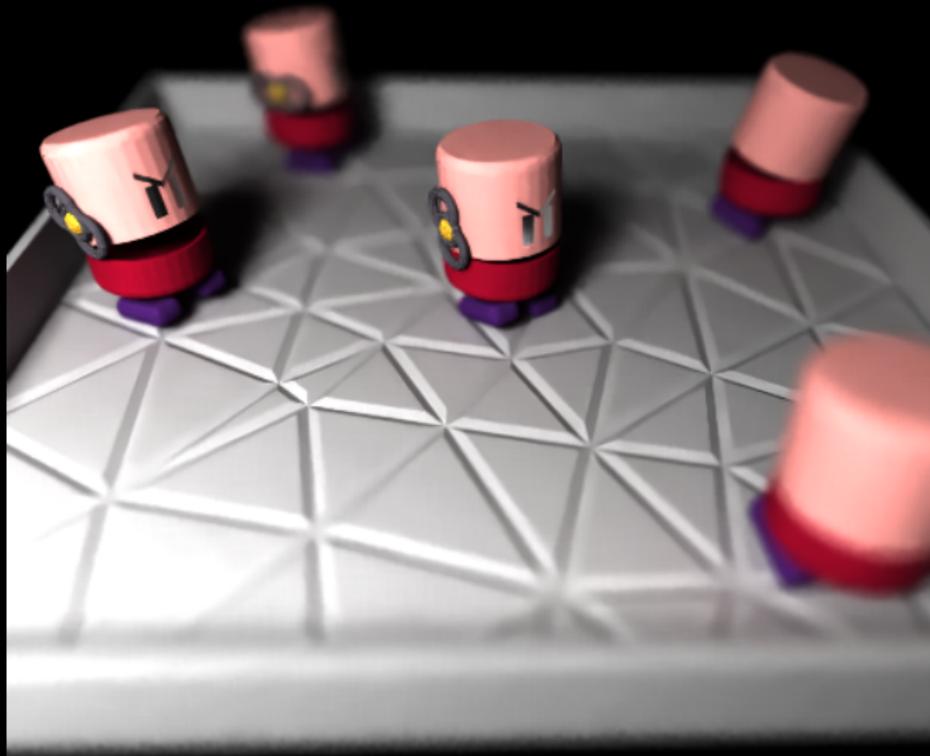
- Given a noisy patch, all the patches that are at a distance less than κ are considered to be similar.
- Replace a patch with the *average of the similar ones*.



$$\bar{u} = \text{[patch]} + \text{[patch]} + \dots + \text{[patch]}$$



Input toasters scene 256spp. psnr: 33.7dB



Filtered toasters scene. psnr: 38.1dB. Gain +14.7dB \approx 33 \times more samples.

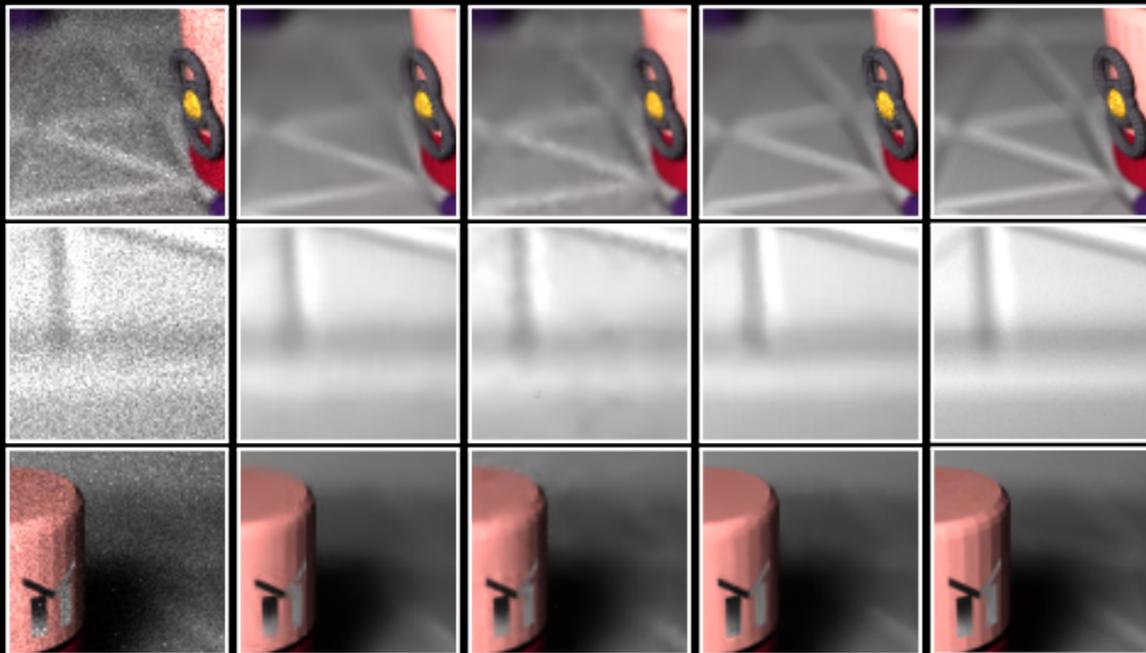
mc eq. time

NL-means eq. time

ASR eq. time

RHF

Reference MC

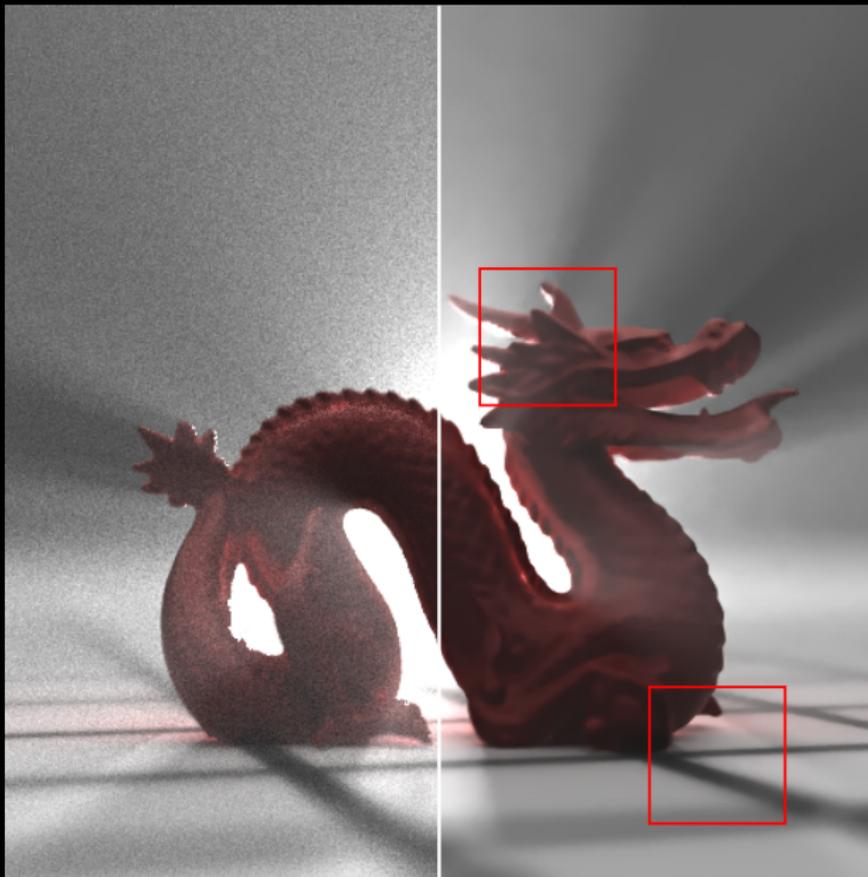


24.8 db [88.8s]

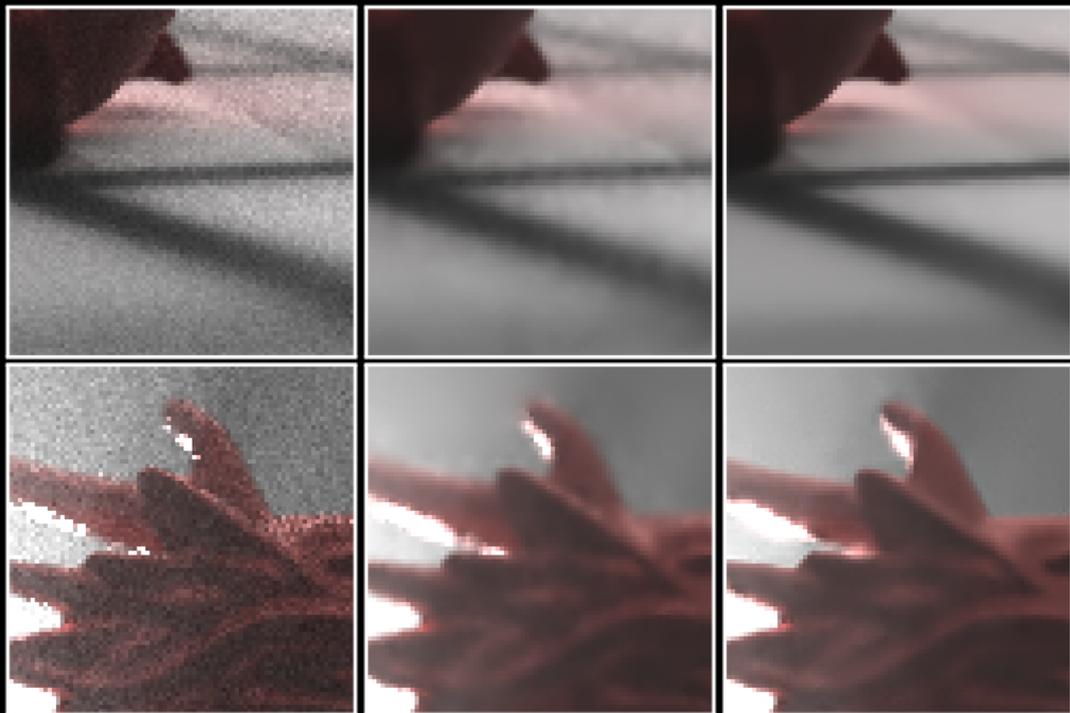
36.1db [88.9s]

35.7db [88.1s]

38.1db [88.8s]



Light interaction with participating media rendered through photon mapping



PM+FG

ASR

RHF

Results in dragon-fog scene (close-ups).



Input 256spp san-miguel scene. psnr: 24.1dB.



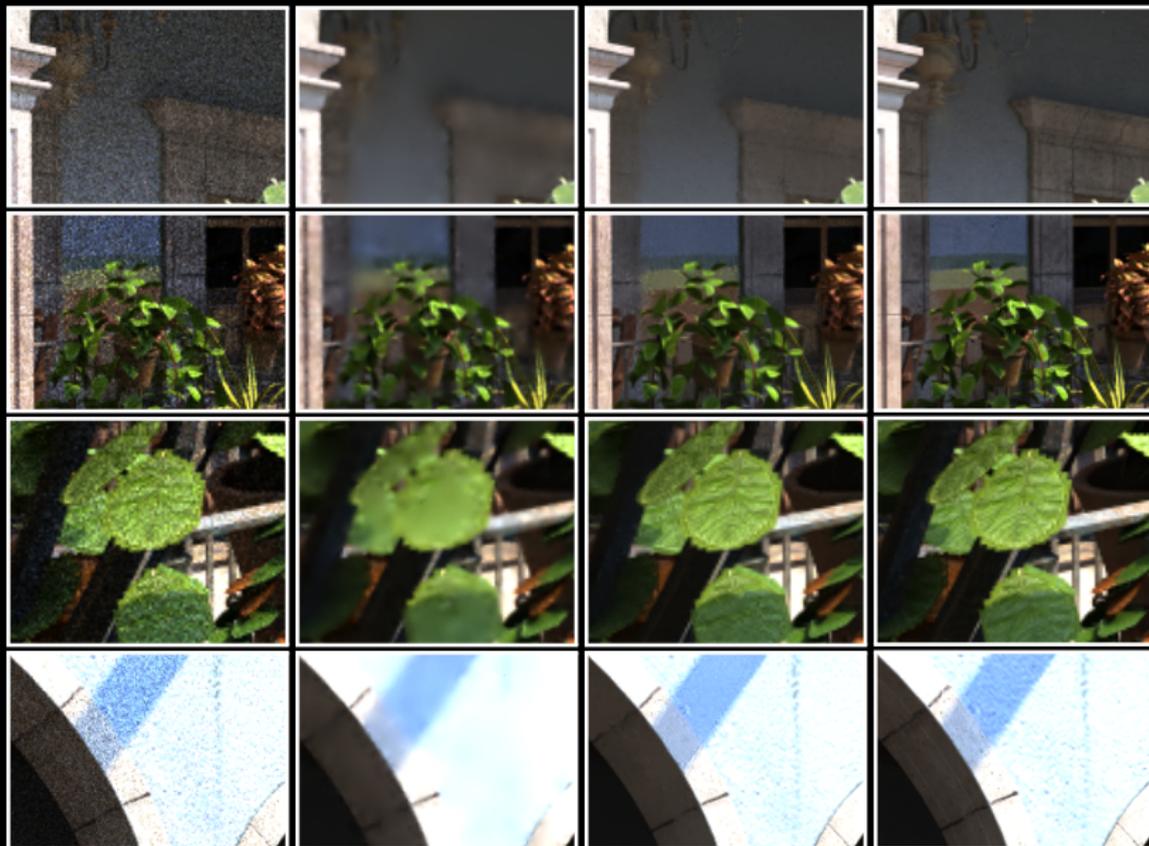
Filtered 256spp san-miguel scene. psnr: 29.8dB. Gain +5.7dB $\approx 4\times$ more samples.

mc noisy input

ASR eq. time

RHF

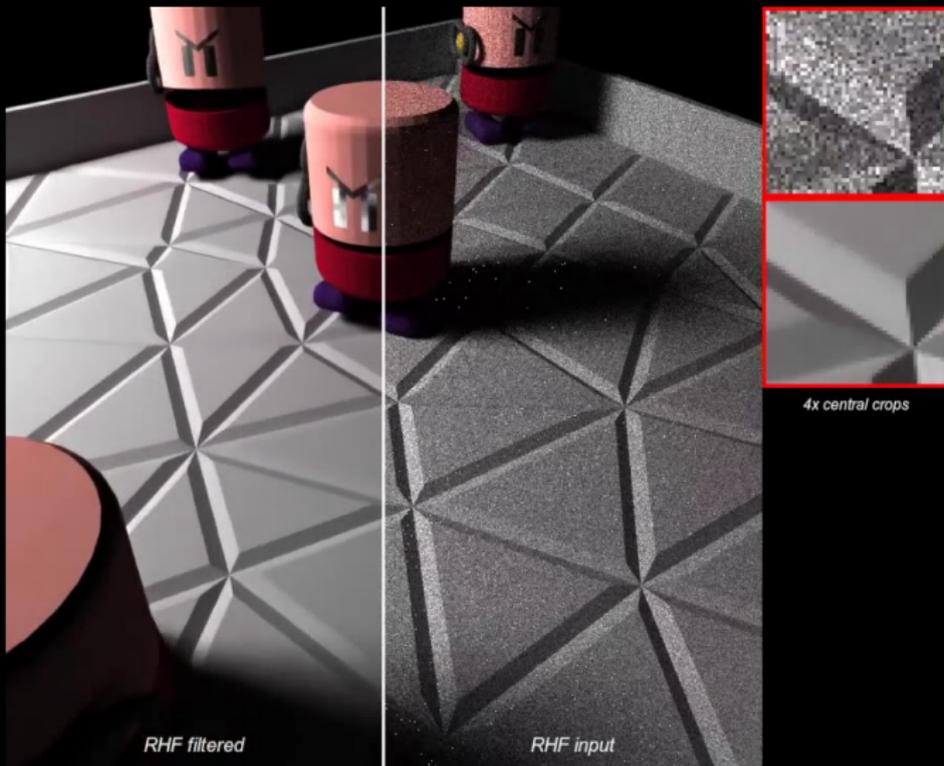
Reference MC



24.1dB [3521s]

26.2dB [4133s]

29.8dB [3952s]



IPOL: Image Processing Online – ipol.im
Detailed Description + Online Demo + Source Code

I presented three different problems where we found a different and simpler solution:

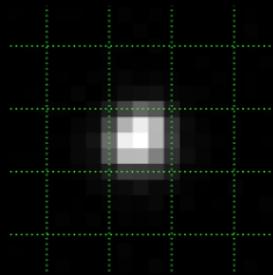
- 1 Instead of adding regularization we “avoid” it by choosing the best single capture (or pair of captures).
- 2 Instead of deblurring via deconvolution do a weighted fusion in the Fourier domain.
- 3 Instead of casting more rays, re-use them using the “auto-similarity” principle.

Thanks.

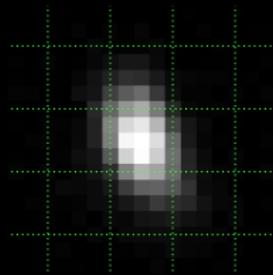
Experiments

Different locations

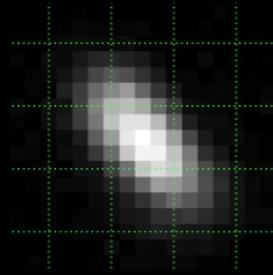
Lens aberration is more significant in image borders.



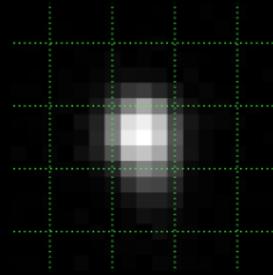
center



left



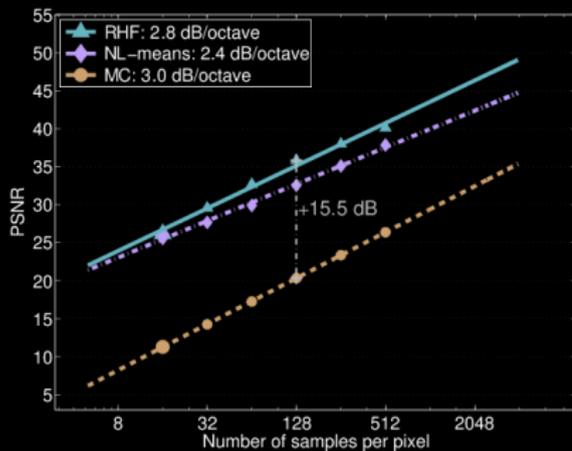
top-left



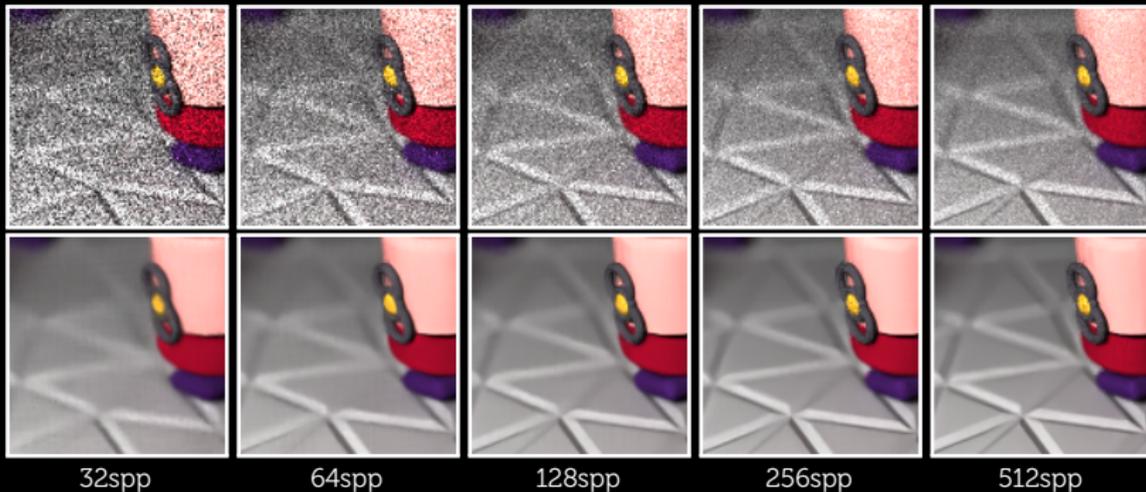
right

RHF: Time Complexity

- Complexity of the filtering at each scale is $O(N \cdot w \cdot b \cdot n_b)$ where N is the number of pixels, b search block size, w patch size, n_b number of bins.
- Computational cost is **independent** of the number of samples.
- In the case that two scales are used the computational cost increases by about 25%.
- If n_s scales are used the computational cost is bounded from above by 133% of the filtering time at the finest resolution.

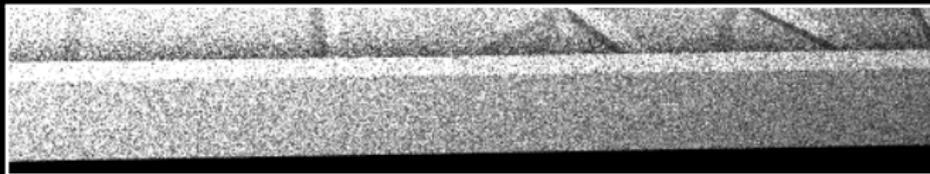


- Pure MC: duplicating spp \rightarrow +3dB
- RHF slope gain is +2.8dB
- RHF: +15dB \approx 35 \times more samples.



Multi-scale Extension

MC noise is white: To remove low-frequency noise filter at different scales and then fuse the filtered images.



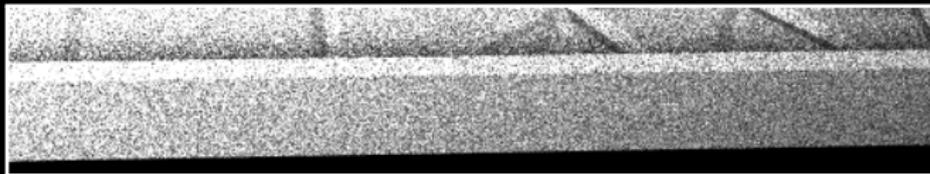
noisy input



single-scale denoised

Multi-scale Extension

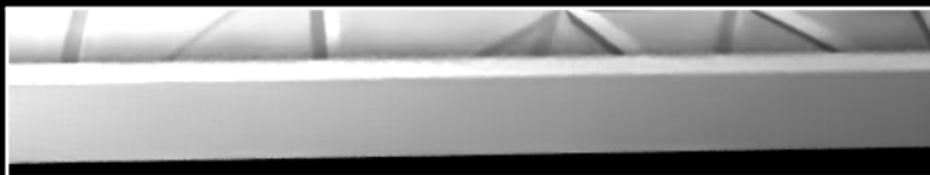
MC noise is white: To remove low-frequency noise filter at different scales and then fuse the filtered images.



noisy input



single-scale denoised



three-scale denoised