

Shape Reconstruction by Photometric Stereo with Unknown Lighting

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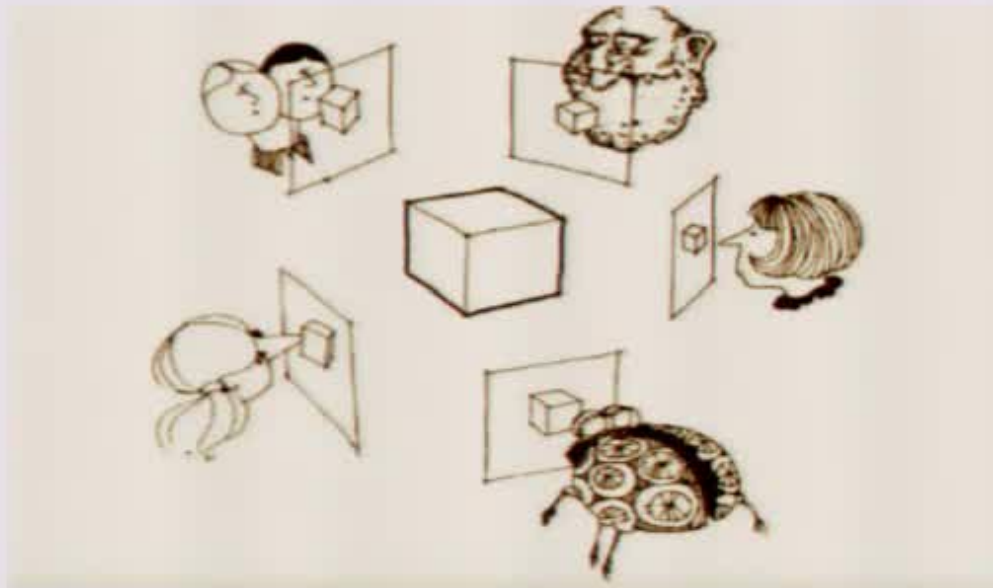
Minisymposium MS14: Recent Advances in Numerical Linear Algebra for Image Processing

Atlanta, U.S.A., October 26–30, 2015

A typical problem in Computer Vision consists of reconstruct the 3D shape of an object, starting from a set of pictures.

There are two main settings:

- **Stereo vision** or **Multivision**: a set of pictures is shot as the camera moves around the object; the lighting is static.

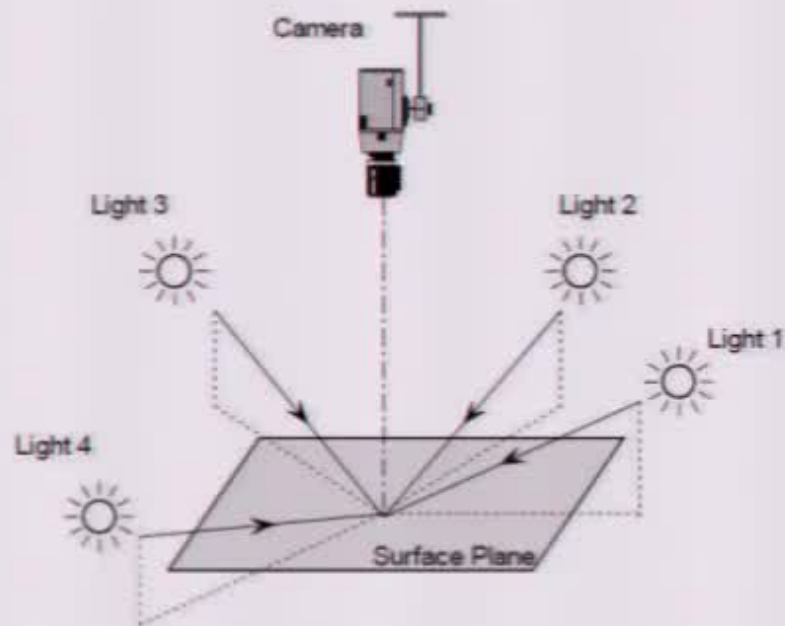


Shape from shading

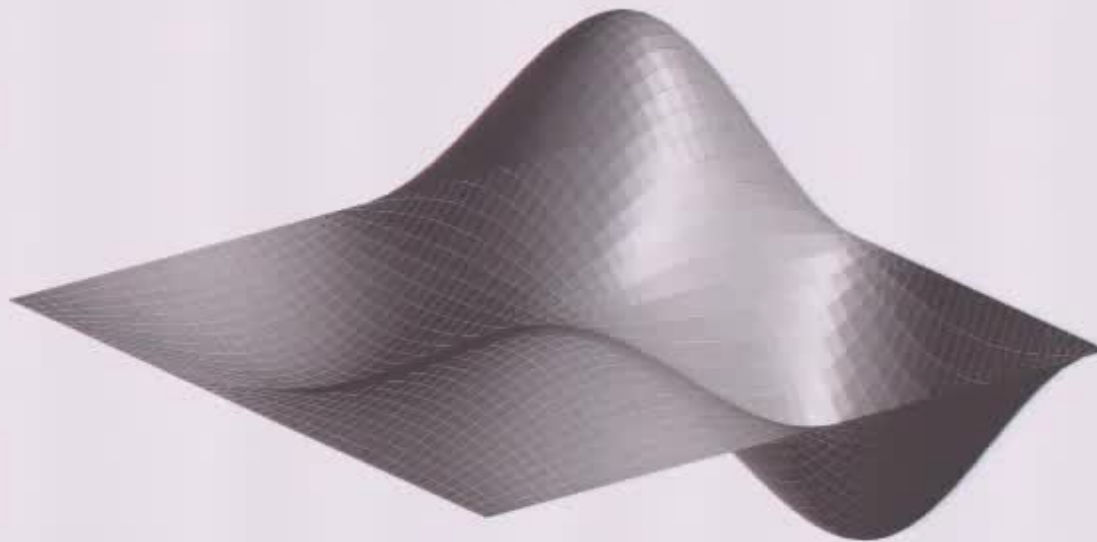
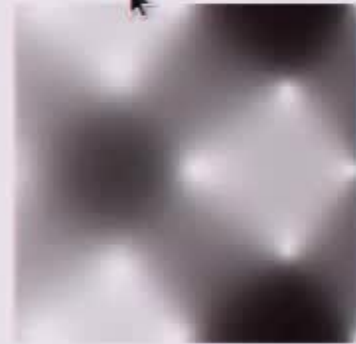
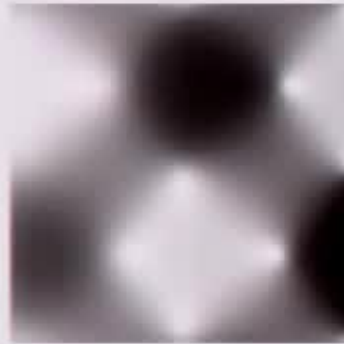
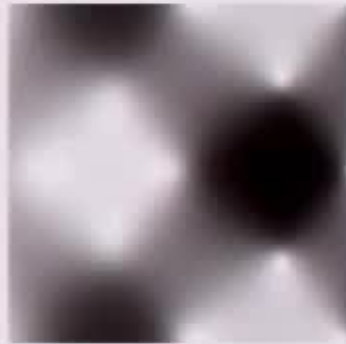
A typical problem in Computer Vision consists of reconstruct the 3D shape of an object, starting from a set of pictures.

There are two main settings:

- **Photometric stereo**: the camera and the object are at fixed positions, pictures correspond to different lighting conditions.



Input & Output



Application to Archaeology



Application to Archaeology

Benefits:

- possibility to document any site in full 3D and color by a commercial camera, a tripod, and a hand-positioned flash;
- instrumentation is low cost and easily transportable;
- it allows for simultaneous operation on different findings by a team of researchers;
- data can be processed on site, **if fast and reliable algorithms are available**;
- these requirements are not met by a **laser scanner**.

Aim:

- producing a **software** capable of **real time processing**, which does not rely on *hand tuning* of parameters, and whose accuracy has been assessed on reference datasets.

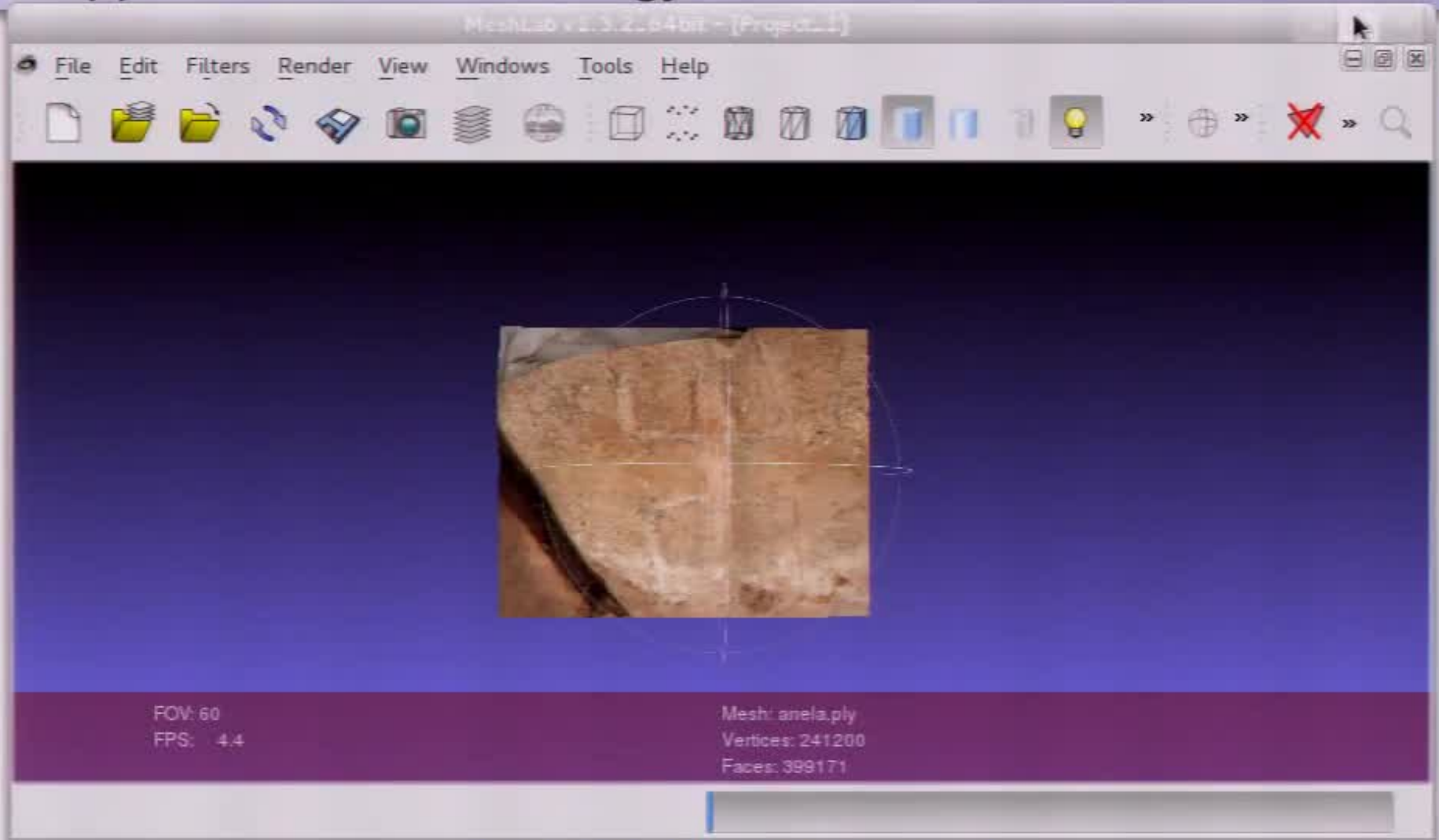
Application to Archaeology



“L'orante” (the praying man)

Necropolis (*Domus de Janas*) “Sos Furrighesos”, Anela, Sardinia, Italy

Application to Archaeology



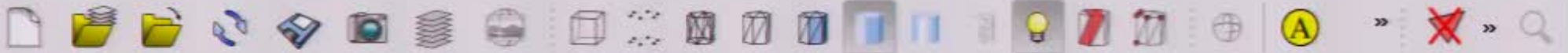
"L'orante" (the praying man)

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FOV: 60
FPS: 13.5

Mesh: anela.ply
Vertices: 241200
Faces: 399171



FOV: 60
FPS: 13.4

Mesh: anela.ply
Vertices: 241200
Faces: 399171



FOV: 60
FPS: 13.8

Mesh: anela.ply
Vertices: 241200
Faces: 399171

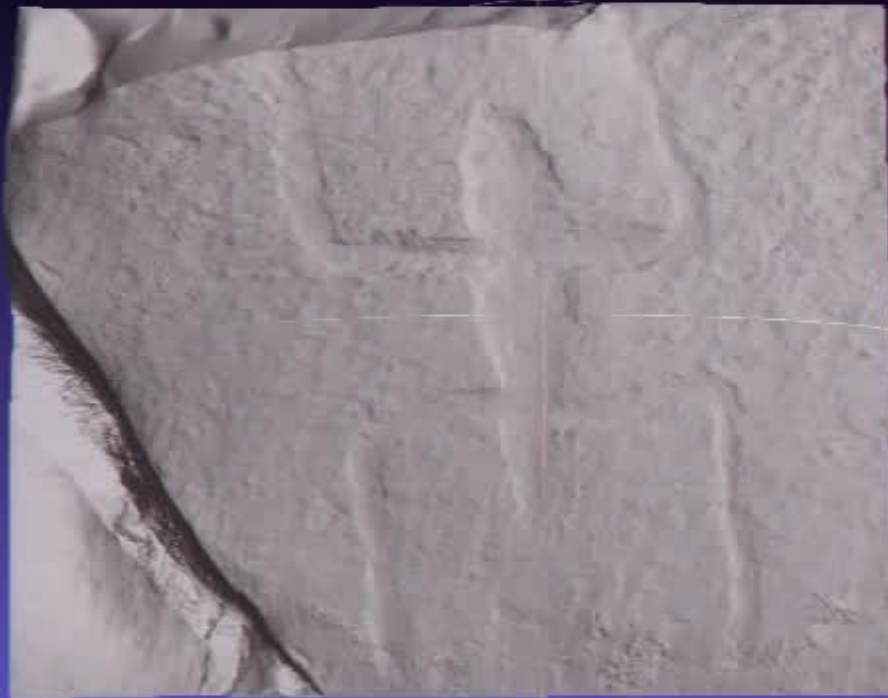


- Render Mode >
- Lighting >
- Color >**
- Shaders >

- None
- Per Mesh
- Per Vertex**
- Per Face

- Background Grid
- Show Vertex Dots
- Show Non-Faux Edges
- Show Boundary Edges
- Show Non Manif Edges
- Show Non Manif Vertices
- Show Face Normals
- Show Vertex Normals
- Show Vert Quality Histogram
- Show Face Quality Histogram
- Show Vertex Principal Curvature Directions
- Show Box Corners
- Show Box Corners (Abs)
- Show Axis
- Show Quoted Box
- Show Vertex Label
- Show Edge Label
- Show Face Label
- Show Camera
- Show UV Tex Param
- Show Texture Seams





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FPS: 20.3

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Application to Archaeology



“L'orante” (the praying man)

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Assumptions

- the surface is Lambertian
- there are no self-obstructions, light reflections, or shades
- the light sources are placed at ∞
- the camera is sufficiently far from the object to avoid perspective distortions

None of these assumptions is perfectly met in practice.

Working assumptions:

- the object is at the origin of a reference system in \mathbb{R}^3
- the camera is on the z-axis, aiming at the origin
- q pictures are available, with light sources at ℓ_t , $t = 1, \dots, q$
- each digital picture \mathbf{m}_t has resolution $r \times s$, with $p = rs$ pixels
- the pictures are vectorized (pixels in lexicographical ordering)

A nonlinear differential model

Let us assume that the directions ℓ_t are known, and write

$$\mathbf{n}(x, y) = \frac{(-u_x, -u_y, 1)^T}{\sqrt{1 + \|\nabla u\|^2}}, \quad \ell_t = \begin{pmatrix} \tilde{\ell}_t \\ \ell_{3t} \end{pmatrix}.$$

Lambert's law becomes

$$\rho(x, y) \frac{\langle -\nabla u(x, y), \tilde{\ell}_t \rangle + \ell_{3t}}{\sqrt{1 + \|\nabla u(x, y)\|^2}} = I_t(x, y), \quad t = 1, \dots, q,$$

i.e., a system of first order nonlinear PDEs of Hamilton–Jacobi type

$$\begin{cases} H_t(x, y, \nabla u(x, y)) = 0, & t = 1, \dots, q, \\ u(x, y) = g(x, y), & (x, y) \in \partial\Omega. \end{cases}$$

Linearization of the differential model

Following [Mecca, Falcone, SJS 2013] we substitute

$$\sqrt{1 + |\nabla u(x, y)|^2} = \rho(x, y) \frac{\langle -\nabla u(x, y), \tilde{\ell}_1 \rangle + \ell_{31}}{l_1(x, y)}$$

in the equations for $t = 2, \dots, q$, to obtain

$$\begin{aligned} [\ell_{11} l_t(x, y) - \ell_{1t} h(x, y)] u_x + [\ell_{21} l_t(x, y) - \ell_{2t} h(x, y)] u_y \\ = [\ell_{31} l_t(x, y) - \ell_{3t} h(x, y)]. \end{aligned}$$



This shows that **the minimal number of images is 2.**

After $u(x, y)$ is computed, the albedo is given by

$$\rho(x, y) = \frac{l_t(x, y)}{\langle \mathbf{n}(x, y), \ell_t \rangle}, \quad \text{for any } t.$$

Well posedness and practical computation

- conditions for the existence of solutions are discussed in [Kozera, Math. Appl. Comput. 1991];
- [Mecca, Falcone, SJIS 2013] studied the problem under more realistic assumptions;
- the obvious choice for discretization is finite differences;
- the solution does not exist when a portion of the surface is shaded in all pictures;
- the solution may not exist for particular light orientations, when $q = 2$;
- taking $q > 2$ is a good choice for getting a solution and for noise reduction, as it does not make data acquisition significantly harder;
- knowing accurately the light positions ℓ_t is a strong requirement.

Poisson formulation

Since p (number of pixels) is large, to compute R and N from

$$RN^T L = M$$

it is required that $q \geq 3$ and that the ℓ_t vectors are independent.

In this case, the minimal number of images is 3.

Numerically differentiating the first two components of \mathbf{n}_k , $k \stackrel{\text{I}}{=} 1, \dots, p$, we get an approximation of the Laplacian of $u(x, y)$ on each point of the discretization.

Solving the Poisson equation

$$\Delta u(x, y) = f(x, y),$$

yields the solution [Dessi, Mannu, R, Tanda, Vanzi, DAACH 2015].

Photometric stereo under unknown lighting

The need for accurate information about the relative position of the lights and the object is a **severe limitation of the method**.

Being able to obtain the lights position from the available set of images opens the possibility of **freehand lighting**, removing the requirement for accurate positioning of lamps, one of the most difficult issues in practical PS.

Some papers conjecture that the problem can be uniquely solved by 4 images (authors often refer to *4-source photometric stereo*).

[Basri, Jacobs, Int. J. Comput. Vis 2007] propose an approach based on the decomposition of the light intensity, as a function of direction, into linear combinations of spherical harmonics.

The solution is not unique!

From the mathematical point of view, the problem consists of determining the factorization

$$\tilde{N}^T L = M,$$

where $\tilde{N} = NR$ is $3 \times p$, L is $3 \times q$ and M is $p \times q$ ($q \geq 3$).

There are infinite solutions: if (\tilde{N}, L) is a solution, any matrix pair $(A_{\mathbb{I}}^{-T} \tilde{N}, AL)$, with $A \in \mathbb{R}^{3 \times 3}$ nonsingular, gives another solution.

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Lemma

The matrices R , N , and L are determined up to a unitary transf., i.e., $(Q\tilde{N}, QL)$ is a solution for any orthogonal matrix $Q \in \mathbb{R}^{3 \times 3}$.

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Lemma

The matrices R , N , and L are determined up to a unitary transf., i.e., $(Q\tilde{N}, QL)$ is a solution for any orthogonal matrix $Q \in \mathbb{R}^{3 \times 3}$.

- the system object-camera-lights may result rotated in the reconstruction, perhaps with axes inversions
- this constitutes a problem if the object is represented by $z = u(x, y)$

Scheme of the algorithm

- ① find the light sources directions, possibly in a rotated system
- ② “straighten” the coordinates system
- ③ integrate the normal field

Some assumptions on the shooting technique are needed.

Determining the lights position

Let the “compact” singular value decomposition of the observations matrix be

$$M = U\Sigma V^T,$$

with $\Sigma = \text{diag}(\sigma_1, \dots, \sigma_q)$, $U \in \mathbb{R}^{p \times q}$, and $V \in \mathbb{R}^{q \times q}$.

In our case, $3 \leq q \ll p$, so standard svd may be sufficient. It can be computed by a Lanczos approach [Baglama, Reichel, BIT 2013].

Since it is expected that $\text{rank}(M) = 3$, we set $W = [\sigma_1 \mathbf{u}_1, \sigma_2 \mathbf{u}_2, \sigma_3 \mathbf{u}_3]^T$ and $Z = [\mathbf{v}_1, \mathbf{v}_2, \mathbf{v}_3]^T$, so that

$$W^T Z \simeq M.$$

This produces the best rank-3 approximation to the data matrix M with respect to the Euclidean norm.

Determining the lights position

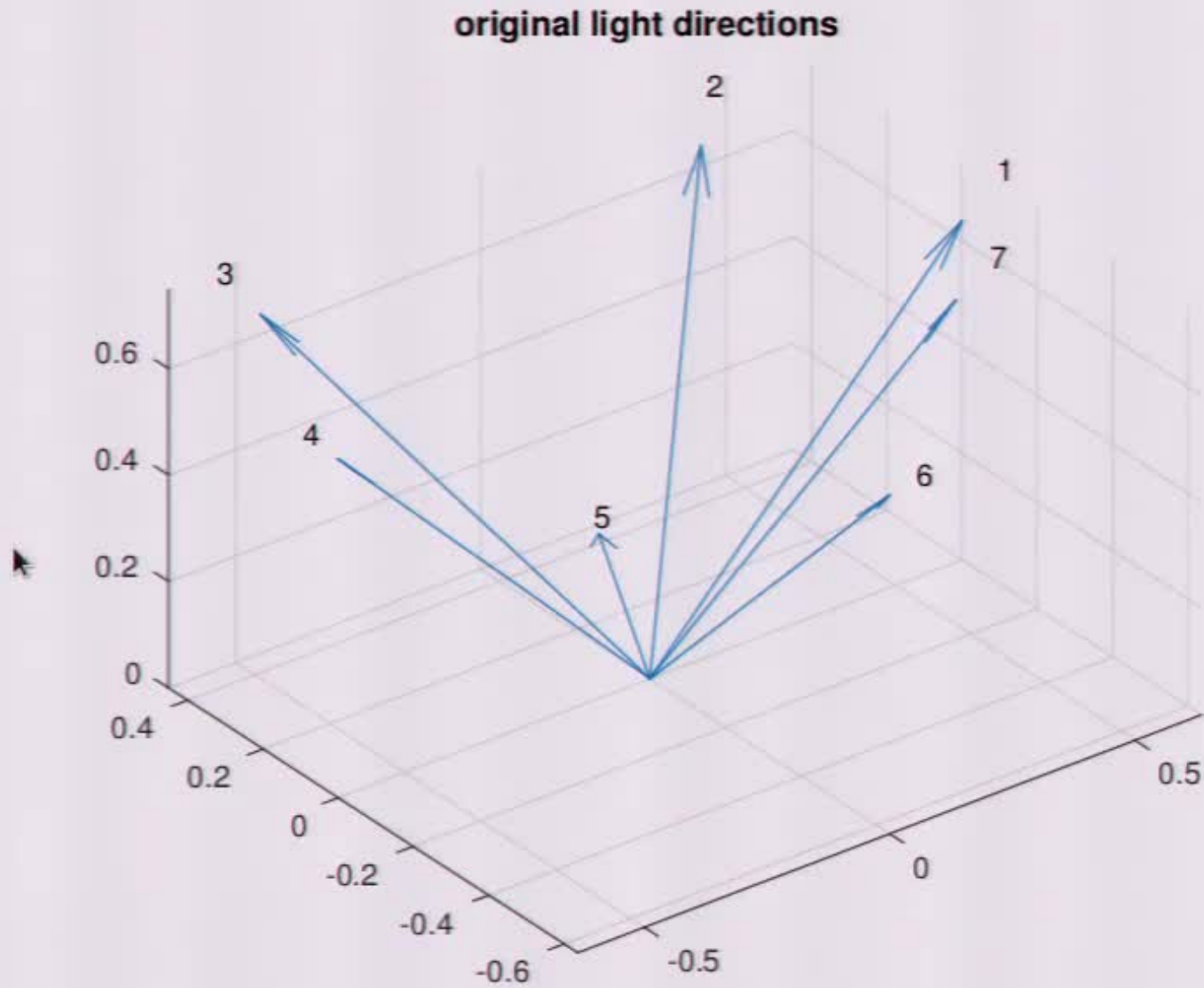
We observed that the initial factorization $W^T Z = M$ gives a good approximation of normal vectors and light directions in a particular situation, that is, when the light vectors span equal angles.

Exploiting the fact that the norms $\|\ell_t\|$ are proportional to the light intensities, which has to be known up to a proportionality constant, it is possible to prove the following result.

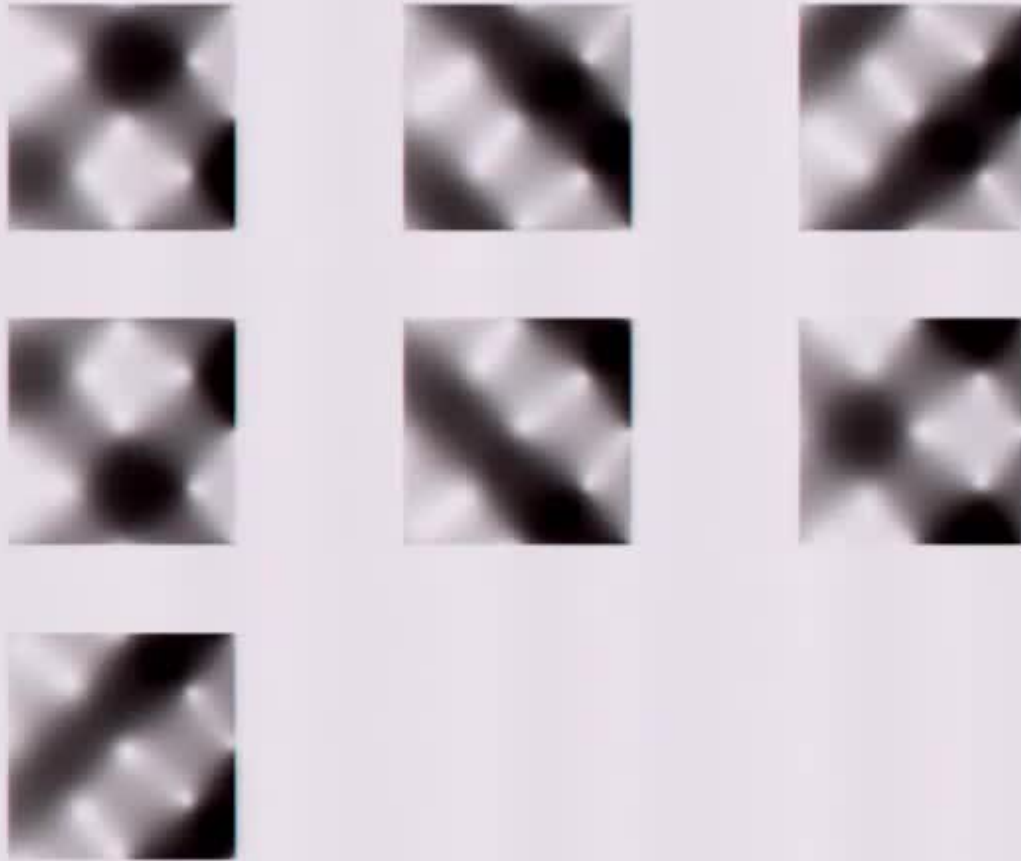
Theorem

The normal vectors and the lights position can be uniquely determined from $RN^T L = M$ up to a unitary transformation, only if at least 6 images taken in different lighting conditions are available.

A numerical experiment



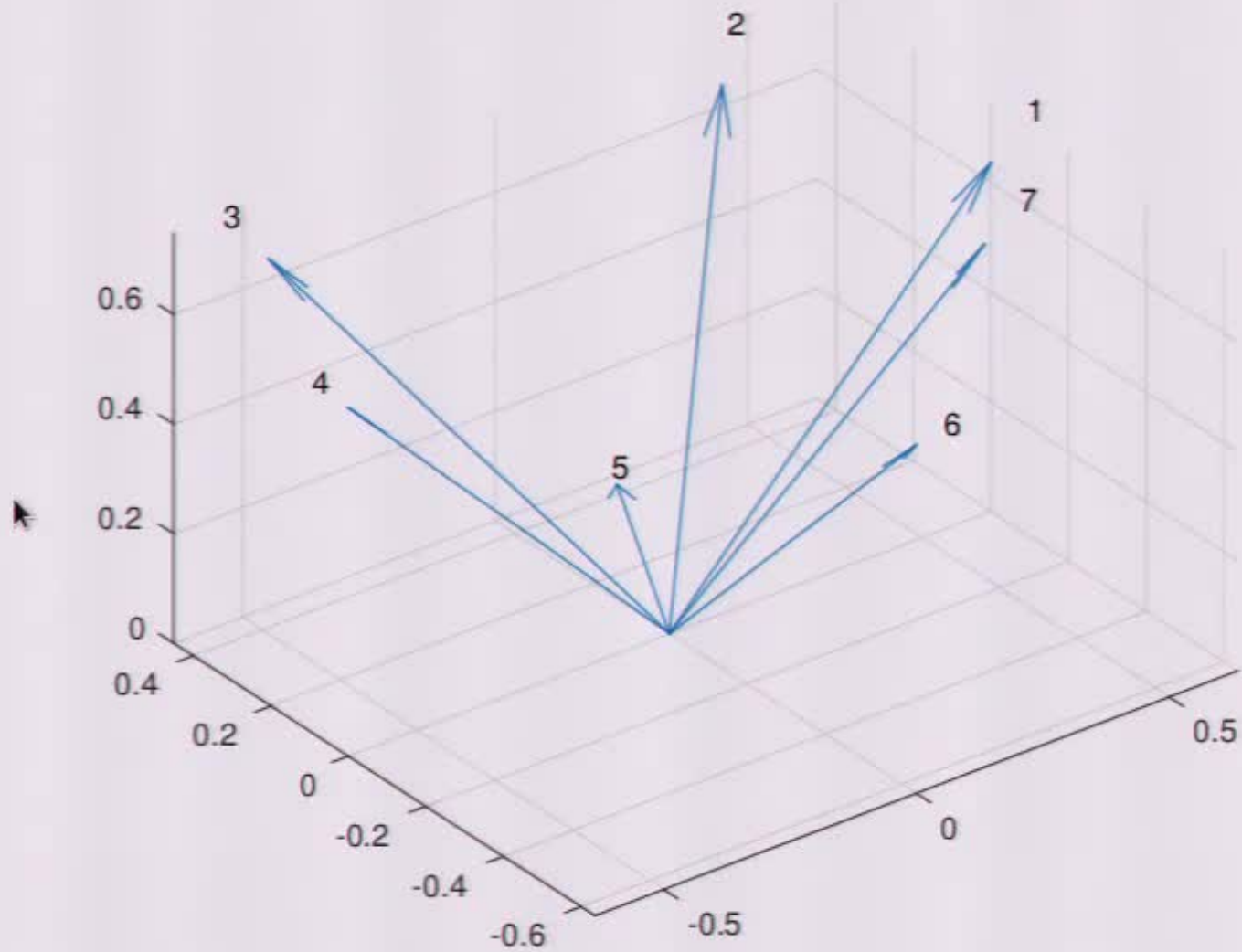
A numerical experiment



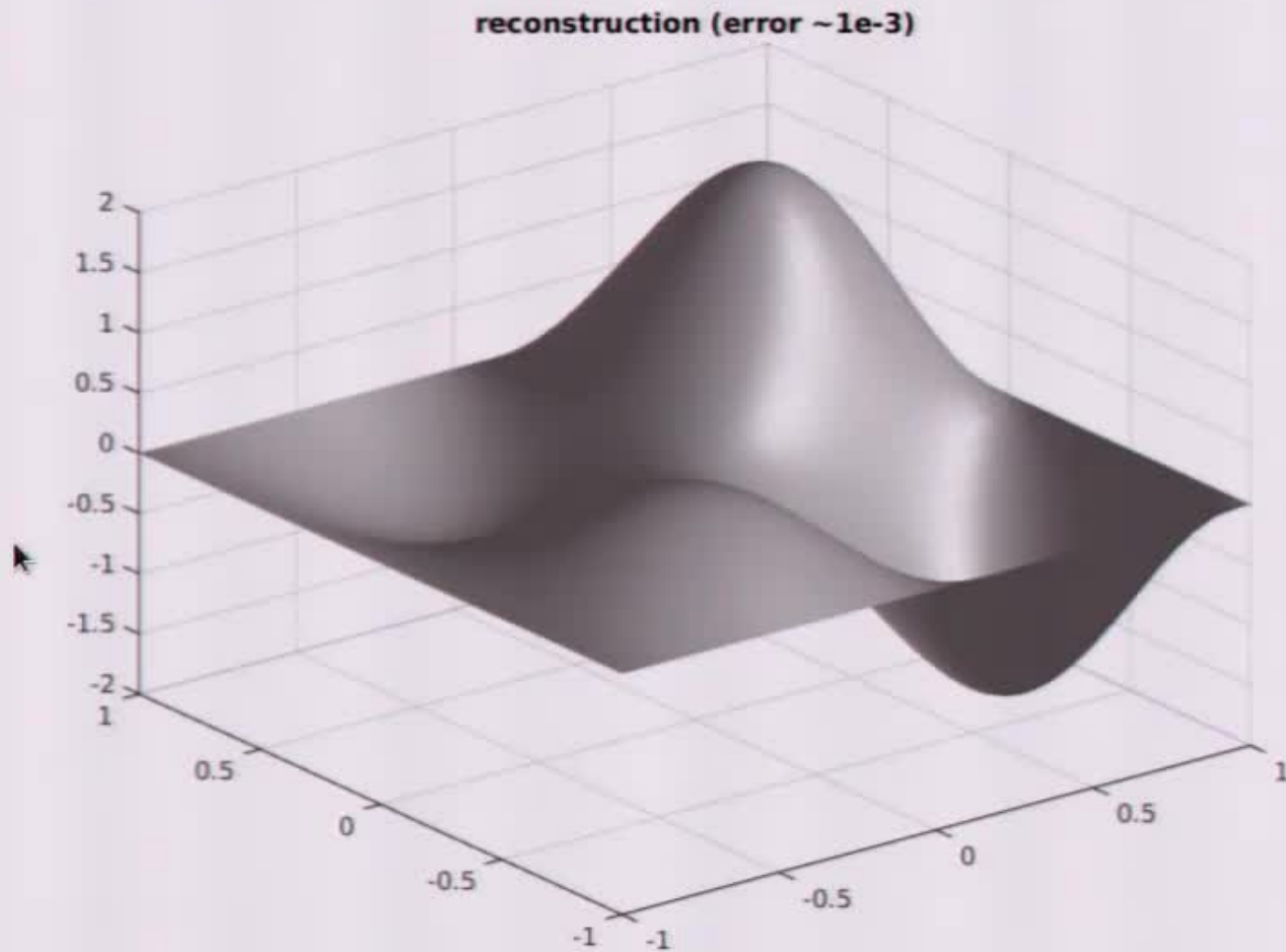
data set (100×100 pixels)

A numerical experiment

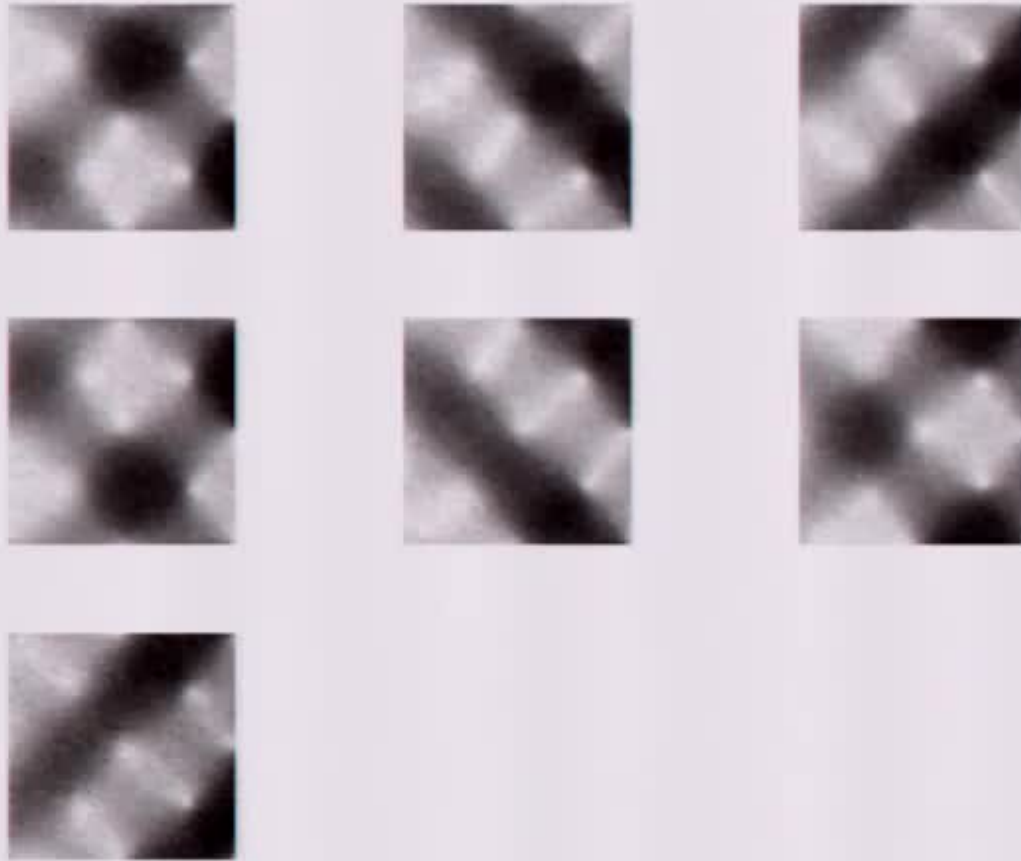
recovered rotated light directions (error $\sim 1e-13$)



A numerical experiment



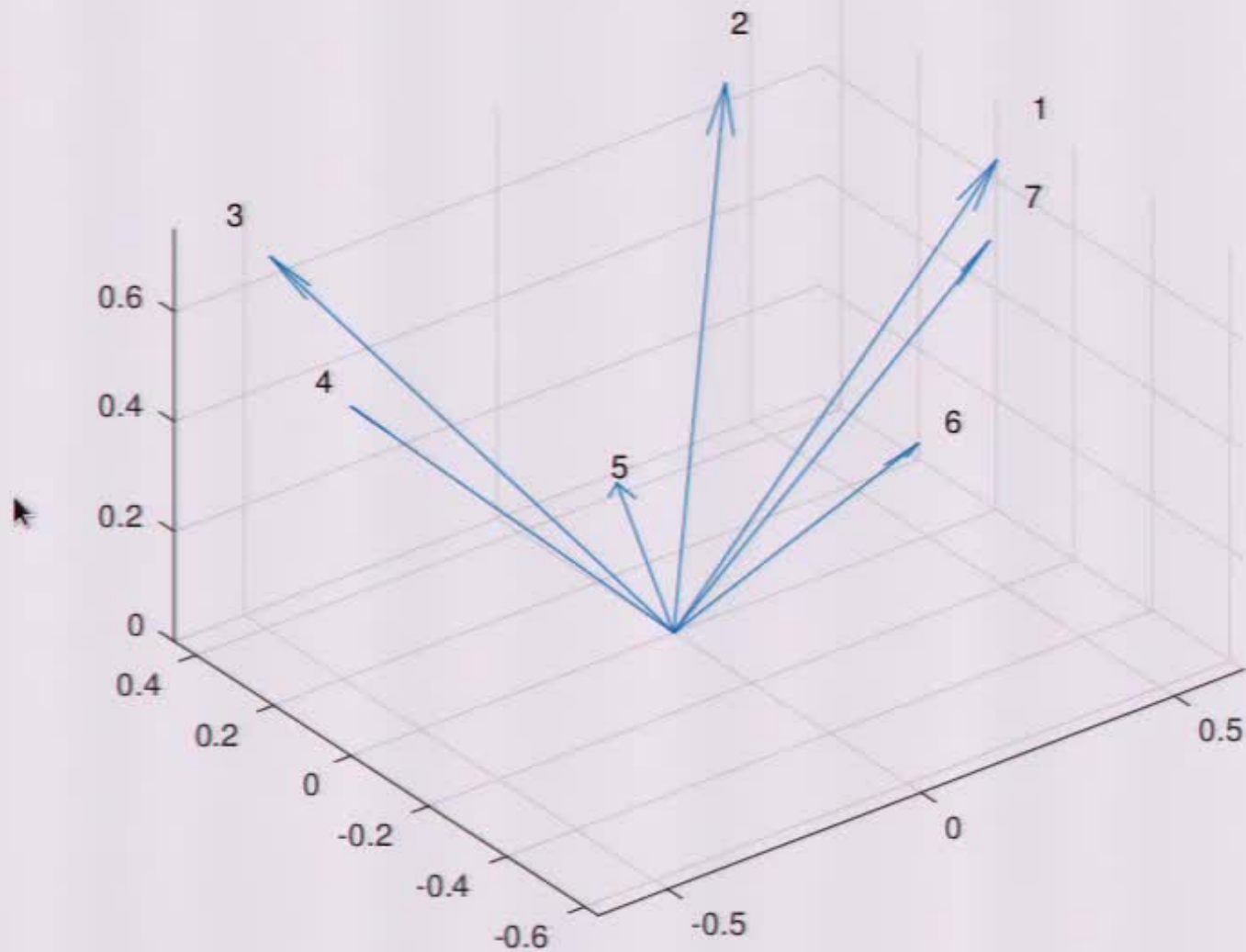
A numerical experiment with 10% white noise on M



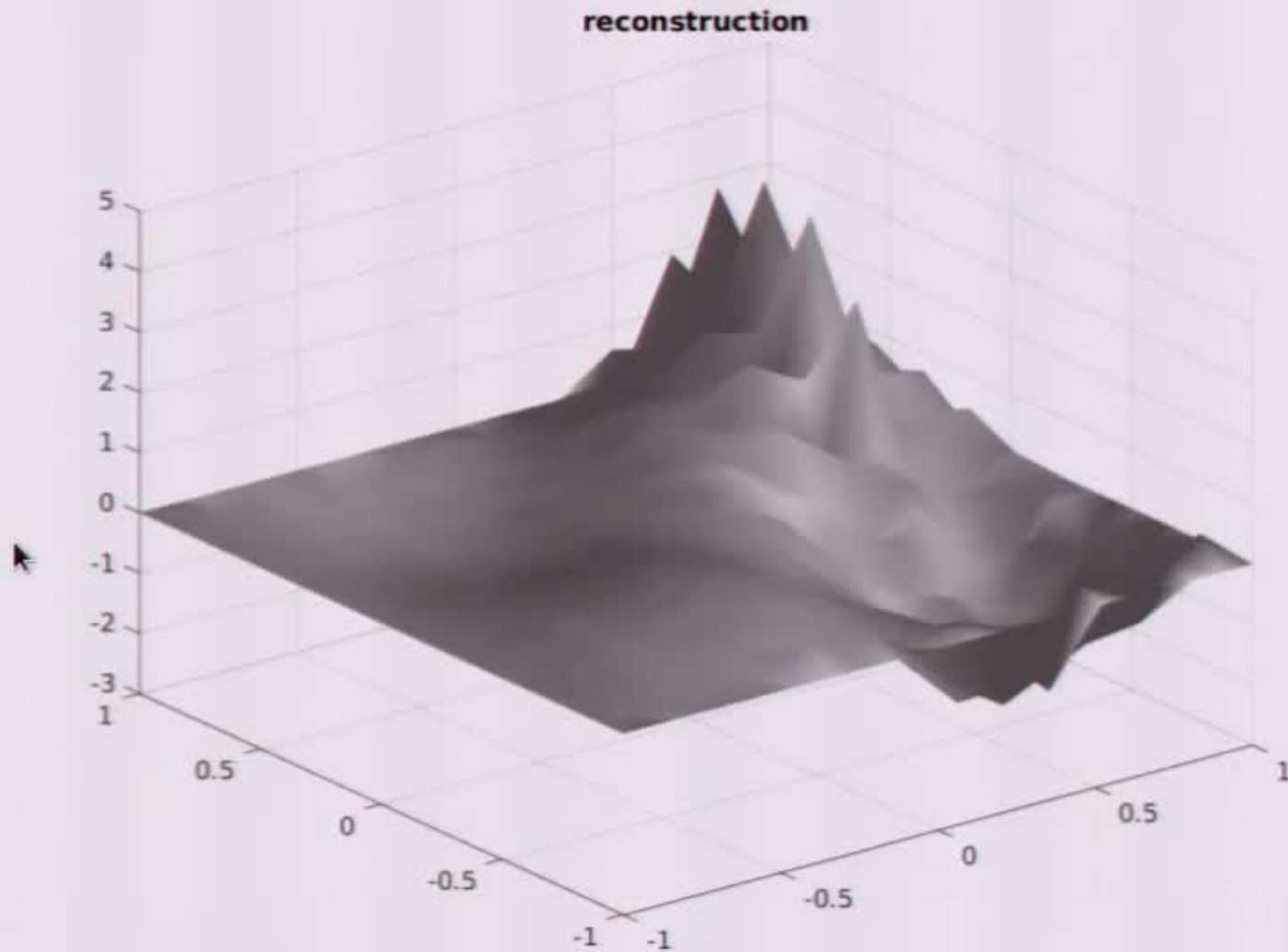
data set with 10% relative noise (100×100 pixels)

A numerical experiment with 10% white noise on M

recovered rotated light directions (error $\sim 5e-3$)



Solution by H-J with 10% white noise on M



Working in unideal light conditions



Real light sources are often far from being ideal

- 1 they **may be placed close to the surface** (especially in narrow locations, like caves or excavation sites)
- 2 they are **differently attenuated** at different points of the object, according to distance

A differential model which keeps into account these effects has been studied in [Mecca et al., SJIS 2014].

The first problem, causes a **spherical deformation** in the reconstructed surface, which can be partially “cured” by the **SVD**, if the light directions are known and the standard model is used.

Working in unideal light conditions

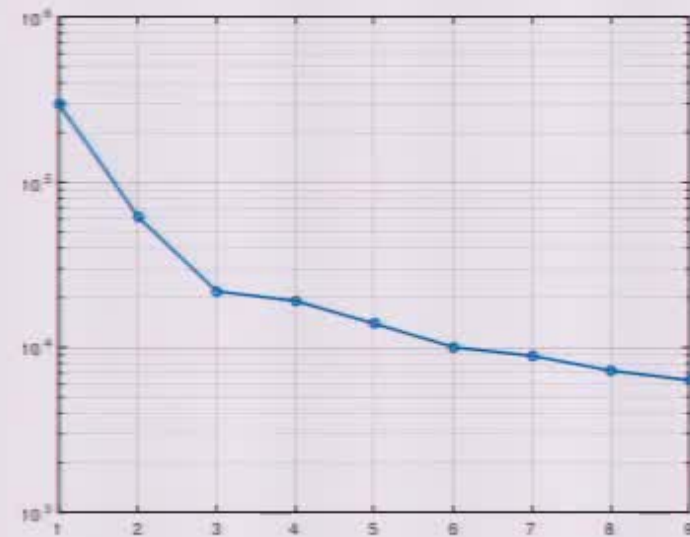


reconstructed surface Z



flattened surface $Z - U_1 \Sigma_1 V_1^T$

The *scholar* data set (close lights with different intensities)

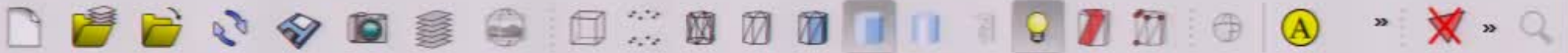


singular values of M

The next steps

- Develop a model for lights at finite distance
- Identify the lights position in the new model

Thank you for your attention!



FOV: 60
FPS: 18.9

Mesh: moneta.ply
Vertices: 140525
Faces: 279552