

PSF reconstruction for extremely large telescopes

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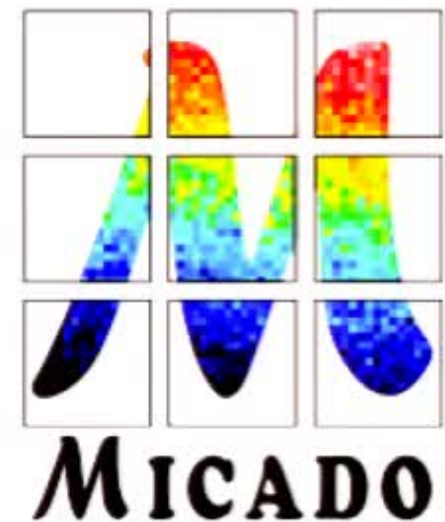
joint work with Ronny Ramlau and Daniela Saxenhuber

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Atlanta, October 27, 2015

Outline

- E-ELT, MICADO and AO
- Atmospheric Tomography
- PSF reconstruction
- Application



European Extremely Large Telescope (E-ELT)

- built by ESO - European Southern Observatory
- operates several astronomical telescopes, e.g. VLT (4x8m, Atacama desert)
- E-ELT: Cerro Armazones (3060m), primary mirror $\varnothing \sim 39$ m, first light 2024
- correct for astronomical seeing: **Adaptive Optics**



MICADO

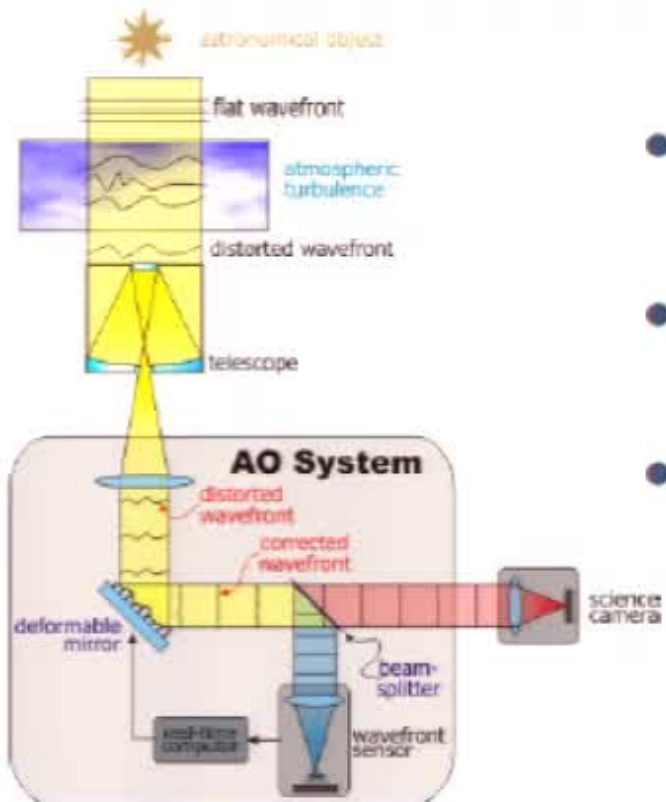
Multi-AO Imaging Camera for Deep Observations



Source: MICADO Consortium

- First light instrument of the E-ELT
- Certain AO modes
- High contrast imaging
- Motions of stars in globular clusters
- Exoplanet search

Adaptive Optics (AO)



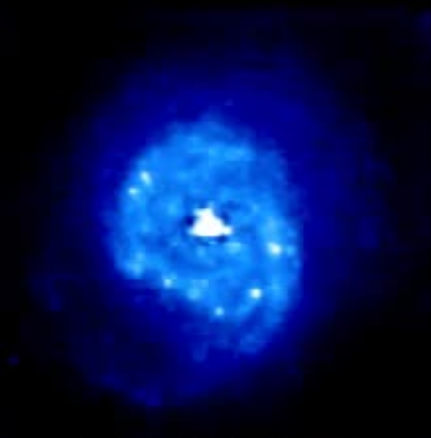
- light wavefronts coming from natural guide stars (NGS) or laser guide stars (LGS)
- atmospheric turbulences distort incoming wavefronts
- AO system:
 - commands for deformable mirror are calculated from wavefront sensor (WFS) measurements
 - deformable mirror adjusts for atmospheric turbulences
 - very short time frame $\sim 1\text{ms}$
- better image quality is achieved

Source:

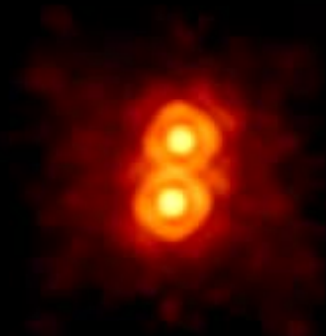
http://www.vikdhillon.staff.shef.ac.uk/teaching/phy217/telescopes/phy217_tel_adaptive.html



Without Adaptive Optics



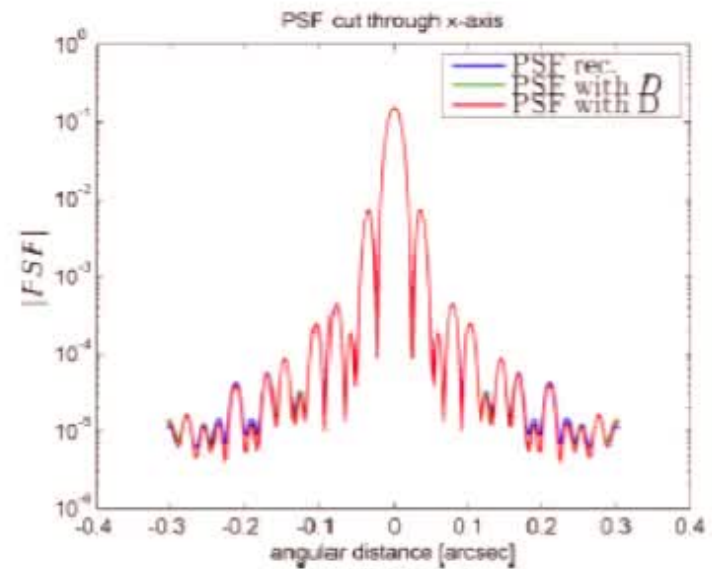
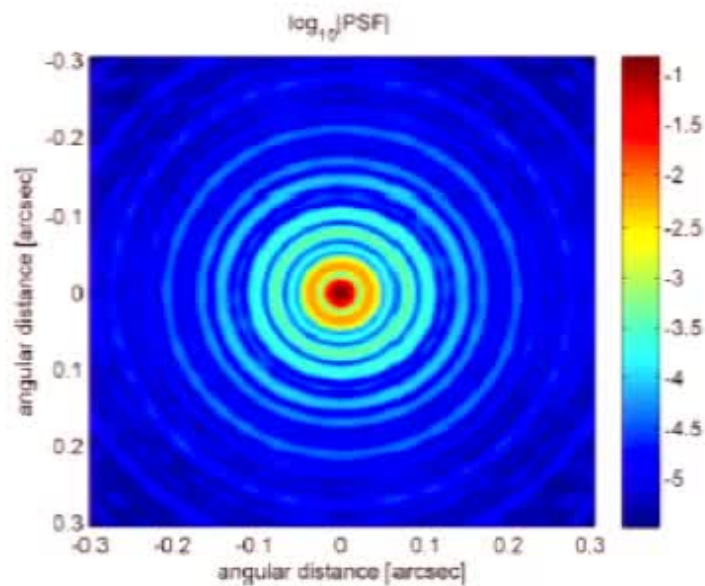
With Adaptive Optics



PSF reconstruction for MCAO

- Use tomographic reconstruction of the atmosphere from measured data (intermediate result of gradient-based method)
- Project through the atmosphere to get PSFs for each desired direction using A_{dir} (as in gradient-based method)
→ pseudo-wavefronts to calculate the corresponding $OTFs$.
- Simulate higher order terms not seen by WFS as before.
- Combine the three parts of the OTF together
- Apply the inverse Fourier transform.

PSF reconstruction for MCAO



Deconvolution

